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# NUCLEAR STRUCTURE AND REACTIONS FROM CHIRAL INTERACTIONS

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Progress in *Ab Initio* Techniques in Nuclear Physics, TRIUMF BC Canada, February 23<sup>th</sup> 2016.



### INTRODUCTION



Nuclear astrophysics

Materials science



### EQUAL TREATMENT OF BOUND AND RESONANT STATES: COUPLE NCSM AND NCSM/RGM (NCSMC)

 Methods develop in this presentation to solve the many body problem





### EQUAL TREATMENT OF BOUND AND RESONANT STATES: COUPLE NCSM AND NCSM/RGM (NCSMC)

 Methods develop in this presentation to solve the many body problem



$$\Psi_{NCSMC}^{(A)} = \sum_{\lambda} c_{\lambda} |A\lambda J^{\pi}T\rangle + \sum_{\nu} \int d\vec{r} g_{\nu}(\vec{r}) \hat{A}_{\nu} \left| \Phi_{\nu\vec{r}}^{(A-a,a)} \right\rangle$$

elastic/inelastic



### EQUAL TREATMENT OF BOUND AND RESONANT STATES: COUPLE NCSM AND NCSM/RGM (NCSMC)

S. Baroni, P. Navrátil and S. Quaglioni PRL110 (2013); PRC93 (2013)

 Methods develop in this presentation to solve the many body problem



 The many body <u>quantum</u> problem is best described by the superposition of both type of wave functions

$$\Psi_{NCSMC}^{(A)} = \sum_{\lambda} c_{\lambda} |A\lambda J^{\pi}T\rangle + \sum_{\nu} \int d\vec{r} g_{\nu}(\vec{r}) \hat{A}_{\nu} \left| \Phi_{\nu \vec{r}}^{(A-a,a)} \right\rangle$$

NCSMC



### COUPLED NCSMC EQUATIONS

S. Baroni, P. Navrátil and S. Quaglioni PRL110 (2013); PRC93 (2013)



Scattering matrix (and observables) from matching solutions to known asymptotic with microscopic R-matrix on Lagrange mesh

# *n*-<sup>4</sup>He SCATTERING: NN VERSUS 3N INTERACTIONS

G. Hupin, J. Langhammer *et al.* PRC88 (2013); G. Hupin, S. Quaglioni and P. Navrátil, to be published in Physica ScriptaSpecial Edition - Nobel Prize '75 anniversary



n-4He scattering

### Two scenarii of nuclear Hamiltonians



- The 3N interactions influence mostly the *P* waves.
- The largest splitting between *P* waves is obtained with NN+3N.

Comparison between NN+3N -ind and NN+3N at N<sub>max</sub>=13 with six <sup>4</sup>He states and 14 <sup>5</sup>He states.

### AB INITIO n-4He SCATTERING

G. Hupin, J. Langhammer *et al.* PRC88 (2013); G. Hupin, S. Quaglioni and P. Navrátil, to be published in Physica ScriptaSpecial Edition - Nobel Prize '75 anniversary









- The convergence pattern looks good.
- The experimental phase-shifts are well reproduced.

### *n*-<sup>4</sup>He ELASTIC CROSS-SECTIONS

G. Hupin, S. Quaglioni and P. Navrátil, to be published as a contribution to the Special Physica Scripta Edition - 40 year anniversary - Nobel Prize '75



Comparison of the elastic cross-section between NN and NN+3N with <sup>4</sup>He (g.s.)



*n*-<sup>4</sup>He elastic cross-section for NN+3N-induced, NN+3N potentials compared to expt. and ENDF evaluation.



• We obtained a better agreement with data when using NN+3N.

 The 3N force is constitutive to the reproduction of the <sup>3</sup>/<sub>2</sub><sup>+</sup> resonance.

### PREDICTION FOR ELASTIC RECOIL DETECTION (ERD)

G. Hupin, S. Quaglioni and P. Navrátil, PRC90 (2014)



p-4He scattering

### Cross-section compared to experiments focused on proton recoil analysis



## NEUTRON-RICH HALO NUCLEUS <sup>11</sup>BE



## 

In the shell model picture g.s. expected to be  $J^{\pi}=1/2^{-1}$ 

(Z=6, N=7)  $^{13}C$  and (Z=8, N=7)  $^{15}O$  have  $J^{\pi}{=}1/2^{-}$  g.s.

- In reality, <sup>11</sup>Be g.s. is **J<sup>π</sup>=1/2**<sup>+</sup> -- parity inversion
- Very weakly bound: E<sub>th</sub>=-0.5 MeV Halo state -dominated by <sup>10</sup>Be-n in the S-wave
- The 1/2<sup>-</sup> state also bound -- only by 180 keV

### Z=4 N=7



Single particle interpretation using nuclear shell model



### Can we describe <sup>11</sup>Be in *ab initio* calculations?

- Continuum must be included
- Does the 3N interaction play a role in the parity inversion?

**FRIUMF** 

### p-<sup>10</sup>C SCATTERING: STRUCTURE OF <sup>11</sup>N RESONANCES

4.0060

<sup>9</sup>B+p



### Limited information about the structure of proton rich <sup>11</sup>N – mirror nucleus of <sup>11</sup>Be halo nucleus

- Incomplete knowledge of <sup>10</sup>C unbound excited states
- Importance of 3N force effects and continuum
- Can structure of exotic nuclei discriminate among different nuclear force models?









### New experiment at ISAC TRIUMF with reaccelerated <sup>10</sup>C

- The first ever <sup>10</sup>C beam at TRIUMF
- Angular distributions measured at  $E_{CM} \sim 4.16$  MeV and 4.4 MeV





### p-10C SCATTERING: STRUCTURE OF 11N RESONANCES

A. Calci, P. Navratil, G. Hupin, S. Quaglioni, R. Roth et al. with IRIS collaboration, in preparation

# **TRIUMF**

NCSMC calculations with chiral NN+3N [N<sup>3</sup>LO NN+N<sup>2</sup>LO 3NF(400), NNLOsat]

- <sup>10</sup>C: 0<sup>+</sup>, 2<sup>+</sup>, 2<sup>+</sup> NCSM eigenstates
- <sup>11</sup>N:  $\geq$ 4 ( $\pi$ =-1= and  $\geq$ 3 ( $\pi$ =+1) NCSM eigenstates





### p-<sup>10</sup>C phase shifts with NN+3N ( $\Lambda$ =400)





14

<sup>10</sup>C+p



### p-10C SCATTERING: STRUCTURE OF 11N RESONANCES

A. Calci, P. Navratil, G. Hupin, S. Quaglioni, R. Roth et al. with IRIS collaboration, in preparation







FROM RESEARCH TO INDUSTRY p-10C SCATTERING: STRUCTURE OF 11N RESONANCES A. Calci, P. Navratil, G. Hupin, S. Quaglioni, R. Roth et al. with IRIS collaboration, in preparation **IRIS** collaboration: **FRIUMF** Area where 3N force effect can be observed A. Kumar, R. Kanungo, A. Sanetullaev et al. 180 180 <sup>10</sup>C+p 10 'C+p 150 150 <sup>6</sup>P<sub>5/2</sub> <sup>2</sup>P<sub>1/2</sub>  $^{2}P$ <sup>2</sup>Р<sub>3/></sub> P 120 120  $^{2}P$ 90 90 δ [deg] δ [deg]  $3/2^{+}$ 60 60 30 30  $3/2^{+}$ chiral NN chiral NN+3NF400 -30 -30 -60L -60L 4 E<sub>kin</sub> [MeV] E<sub>kin</sub> [MeV] 2 3 5 6 3 6 p-<sup>10</sup>C  $p^{-10}C$ E<sub>kin</sub>=4.16 MeV Ekin=4.4 MeV 100 100 NCSMC NCSMC dS/dW [mb/sr] ds/dW [mb/sr] 10 chiral NN 10 chiral NN chiral NN+3NF400 chiral NN+3NF400 120 O<sub>CM</sub> [deg] 60 90 120 150 180 60 90 150 180 Q<sub>CM</sub> [deg]

16



### STRUCTURE OF <sup>11</sup>Be FROM CHIRAL NN+3N FORCES

A. Calci, P. Navratil, G. Hupin, S. Quaglioni, R. Roth et al. with IRIS collaboration, in preparation

# **RIUMF**

NCSMC calculations including chiral 3N (N<sup>3</sup>LO NN+N<sup>2</sup>LO 3NF400)

- <sup>10</sup>Be: 0<sup>+</sup>, 2<sup>+</sup>, 2<sup>+</sup> NCSM eigenstates
  - <sup>11</sup>Be:  $\geq 6$  ( $\pi$ =-1) and  $\geq 3$  ( $\pi$ =+1) NCSM eigenstates



Cez

### <sup>11</sup>Be WITHIN NCSMC: DISCRIMINATION AMONG CHIRAL NUCLEAR FORCES



A. Calci, P. Navratil, G. Hupin, S. Quaglioni, R. Roth et al. with IRIS collaboration, in preparation



### p-10C SCATTERING: STRUCTURE OF 11N RESONANCES

A. Calci, P. Navratil, G. Hupin, S. Quaglioni, R. Roth et al. with IRIS collaboration, in preparation



### <sup>4</sup>He(*d*,*d*)<sup>4</sup>He COMPARISON OF INTERACTION

G. Hupin, S. Quaglioni and P. Navrátil, PRL114 (2015)

potential

with







NN+3N-induced, NN+3N-full

 $\lambda = 2.0 \text{ fm}^{-1}$ .

90 90 45 -45 -45 -90 -135 0  $2 E_{kin}$  [MeV] 4 6

 ${}^{3}D_{3}$ 

180

135

- Best results in a decent model space (N<sub>max</sub>=11).
- The  ${}^{3}D_{3}$  resonance is reproduced but the  ${}^{3}D_{2}$  and  ${}^{3}D_{1}$  resonance positions are underestimated.
- The 3N force corrects the *D*-wave resonance positions by increasing the spin-orbit splitting.
- There is room for improvements.

### <sup>6</sup>Li SPECTRUM, NCSMC VS NCSM/RGM

G. Hupin, S. Quaglioni and P. Navrátil, PRL114 (2015)

### Comparison between NCSMC vs NCSM



- The 3N force is essential to get the correct <sup>6</sup>Li g.s. energy and splitting between the 3<sup>+</sup> and 2<sup>+</sup> states.
- The <sup>6</sup>Li g.s. is well reproduced.
- There is room for improvements, in particular regarding the 3<sup>+</sup> state.



<sup>4</sup>He(*d*,*d*)<sup>4</sup>He CROSS-SECTION G. Hupin, S. Quaglioni and P. Navrátil, PRL114 (2015)







The bulk of the cross section is well reproduced for a large set of kinetic energy and scattering angle.

<sup>4</sup>He(*d*, <sup>4</sup>He)*d* CROSS-SECTION G. Hupin, S. Quaglioni and P. Navrátil, PRL114 (2015)







Comparison to experiment of the *d*-<sup>4</sup>He elastic recoil differential cross section of NCSMC with NN+3N potential at  $\lambda$ =2.0 fm<sup>-1</sup>.



The 3<sup>+</sup> resonance is missed. As its width is very narrow, it has little impact and the bulk of the crosssection.

### FIRST STEPS TOWARDS AB INITIO CALCULATIONS OF FUSION WITH NCSM/RGM

P. Navrátil, S. Quaglioni, PRL108 (2012)



#### <sup>3</sup>H(*d*,*n*)<sup>4</sup>He astrophysical S-factor **BR51** $d^{+3}H \rightarrow n^{+4}He$ (a)**AR52** 30 CO52 AR54 He55 25 **GA56** b **BA57** S-factor [MeV GO61 20KO66 MC73 **MA75** JA 84 BR87 0d\*+0d'\* 7d\*+5d'\* Evidence of 9d\*+5d' incomplete model 100 1000 (nuclear force) $E_{\rm kin}$ [keV]

NCSM/RGM results for the  ${}^{3}\text{He}(d,n){}^{4}\text{He}$ astrophysical S-factor compared to beamtarget measurements.

Calculated S-factors converge with the inclusion of the virtual breakup of the deuterium, obtained by means of excited  ${}^{3}S_{1}-{}^{3}D_{1}$  (*d*<sup>\*</sup>) and  ${}^{3}D_{2}$  (*d*<sup>'\*</sup>) pseudo-states.

Incomplete nuclear interaction: requires 3N force (SRG-induced + "real")

### FIRST STEPS TOWARDS AB INITIO CALCULATIONS OF FUSION

G. Hupin, S. Quaglioni, P. Navrátil work in progress



### d-t fusion

# n-<sup>4</sup>He phaseshifts with NCSMC and the chiral two- and three-nucleon force



 $\lambda$ =2.0 fm<sup>-1</sup>, with eigenstates of <sup>5</sup>He at N<sub>max</sub> =9.

- Perspective to provide accurate t(d,n)<sup>4</sup>He fusion cross-section for the effort toward earth-based fusion energy generation.
- The *d*-t fusion is known to be very sensitive to the spin-orbit and isospin part of the nuclear interaction.

### FIRST STEPS TOWARDS AB INITIO CALCULATIONS OF FUSION



G. Hupin, S. Quaglioni, P. Navrátil work in progress

### Towards d-t fusion with NCSMC: comparison between effective interactions



 $n+^{4}$ He(g.s.) phase shifts with NN+3N potential, with eigenstates of <sup>5</sup>He.

 $\lambda = 1.7 \text{ fm}^{-1} \text{ and } \hbar\Omega = 16 \text{ MeV}$ 

$N_{max}$	${}^{5}\text{He} \left(\frac{3}{2}^{+}\right)$	$^{4}\mathrm{He}$	$^{3}\mathrm{H}$	d
6	-5.1574	-28.1739	-8.2909	-2.0404
8	-7.2529	-28.3677	-8.3893	-2.092
10	-8.1861	-28.4348	-8.4455	-2.1654
12			-8.4546	-2.1781
$\infty$	-8.9373	-28.4693	-8.4565	-2.224
		I		I
	$\lambda = 2.0 \text{ fm}^3$	$^{-1}$ and $\hbar\Omega$	= 20  MeV	7
N <sub>max</sub>	$\lambda = 2.0 \text{ fm}$ <sup>5</sup> He $\left(\frac{3}{2}^{+}\right)$	$^{-1}$ and $\hbar\Omega$ $^{4}\text{He}$	$= 20 \text{ MeV}$ $^{3}\text{H}$	/ d
N <sub>max</sub> 6	$\lambda = 2.0 \text{ fm}$ $^{5}\text{He} (\frac{3}{2}^{+})$ $^{-1.9037}$	$ \begin{array}{c} ^{-1} \text{ and } \hbar\Omega \\ ^{4}\text{He} \\ ^{-27.7923} \end{array} $	= 20  MeV $^{3}\text{H}$ -8.0971	d -1.9199
N <sub>max</sub> 6 8	$\lambda = 2.0 \text{ fm}$ $^{5}\text{He} \left(\frac{3}{2}^{+}\right)$ $^{-1.9037}$ $^{-4.9122}$	$^{-1}$ and $\hbar\Omega$ $^{4}\text{He}$ $^{-27.7923}$ $^{-28.2341}$	$= 20 \text{ MeV}$ $^{3}\text{H}$ $^{-8.0971}$ $^{-8.2721}$	d -1.9199 -1.9633
N <sub>max</sub> 6 8 10	$\lambda = 2.0 \text{ fm}$ $5 \text{He} \left(\frac{3}{2}^{+}\right)$ $-1.9037$ $-4.9122$ $-6.6422$	$^{-1}$ and $\hbar\Omega$ $^{4}\text{He}$ $^{27.7923}$ $^{28.2341}$ $^{28.4078}$	$= 20 \text{ MeV}$ $^{3}\text{H}$ $^{-8.0971}$ $^{-8.2721}$ $^{-8.4099}$	d -1.9199 -1.9633 -2.1172
N <sub>max</sub> 6 8 10 12	$\lambda = 2.0 \text{ fm}$ <sup>5</sup> He ( $\frac{3}{2}^+$ ) -1.9037 -4.9122 -6.6422 -7.7062	$^{-1}$ and $\hbar\Omega$ $^{4}\text{He}$ $^{27.7923}$ $^{28.2341}$ $^{28.4078}$ $^{28.4438}$	$= 20 \text{ MeV}$ $^{3}\text{H}$ $^{-8.0971}$ $^{-8.2721}$ $^{-8.4099}$ $^{-8.4387}$	d -1.9199 -1.9633 -2.1172 -2.1351

NCSM convergence of compound and cluster states.

### FIRST STEPS TOWARDS AB INITIO CALCULATIONS OF FUSION

G. Hupin, S. Quaglioni, P. Navrátil work in progress







 $n+^{4}$ He(g.s.) phase shifts with NN+3N potential, with eigenstates of <sup>5</sup>He.

$\lambda = 1.7$	$\rm fm^{-1}$	and	$\hbar\Omega =$	16	MeV
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	_		_					
$N_{\rm max}$	${}^{5}\text{He}\left(\frac{3}{2}^{+}\right)$	$^{4}\mathrm{He}$	$^{3}\mathrm{H}$	d				
6	42.29	1.04	1.96	8.26				
8	18.85	0.36	0.80	5.94				
10	8.41	0.12	0.13	2.63				
12	-	-	0.02	2.06				
$\lambda = 2.0 \text{ fm}^{-1} \text{ and } \hbar\Omega = 20 \text{ MeV}$								
N <sub>max</sub>	${}^{5}\text{He}\left(\frac{3}{2}^{+}\right)$	$^{4}\mathrm{He}$	$^{3}\mathrm{H}$	d				
6	78.70	2.38	4.25	13.67				
8	45.04	0.83	2.18	11.72				
10	25.68	0.22	0.55	4.80				
12	13.78	0.09	0.21	4.00				
Relative error (%) with respect to								
converged value								

A smaller frequency allows us to capture the dilute nature of the  $3/2^+$  resonance.

### CONCLUSIONS AND OUTLOOK



Evolution of stars, birth, main sequence, death

We are extending the *ab initio* NCSM/RGM approach to describe low-energy reactions with two- and three-nucleon interactions.

We are able to describe:

- Nucleon-nucleus collisions with NN+3N interaction
- Deuterium-nucleus collisions with NN+3N interaction as the n-n
- NCSMC for single- and two-nucleon projectile

Work in progress:

- Fusion reactions with our best complete *ab initio* approach
- The present NNN force is "incomplete", need to go to N<sup>3</sup>LO