

# Nuclear Structure Effects in Light Muonic Atoms

Nir Nevo Dinur<sup>1</sup>

Chen Ji<sup>2</sup>, Javier Hernandez<sup>2,3</sup>, Sonia Bacca<sup>2,3</sup>, Nir Barnea<sup>1</sup>

<sup>1</sup>The Hebrew University of Jerusalem, Israel

<sup>2</sup>TRIUMF, Vancouver, BC, Canada

<sup>3</sup>University of Manitoba, Winnipeg, Canada

TRIUMF Workshop — Feb. 17th 2015

Phys. Rev. Lett. **111**, 143402 (2013)

Phys. Rev. C **89**, 064317 (2014)

Few-Body Syst. **55**, 917 (2014)

Phys. Lett. B **736**, 344 (2014)



האוניברסיטה העברית בירושלים  
The Hebrew University of Jerusalem



- **Introduction**

- The proton radius puzzle
- Lamb shift, charge radius & polarization

- **Methods**

- *Ab-initio* calculation of nuclear polarization
- The LSR method

- **Results**

- $\mu^4\text{He}^+$
- $\mu\text{D}$
- $\mu^3\text{He}^+$  &  $\mu\text{T}$  (preliminary)

- **Summary**

- **Outlook**

# Proton radius puzzle

## How large is the proton?

- $r_p$  from electron-proton interaction

1.  $e$ - $p$  scattering:  $r_p = 0.875(10)$  fm
2. Hydrogen spectroscopy:  $r_p = 0.8768(69)$  fm
3.  $\implies$  CODATA-2010:  $r_p = \mathbf{0.8775(51)}$  fm

Mohr *et al.*, Rev. Mod. Phys. (2012)



## How large is the proton?

### ● $r_p$ from electron-proton interaction

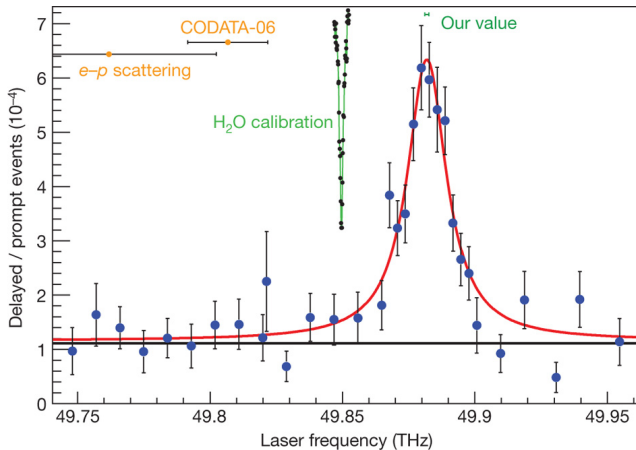
1.  $e$ - $p$  scattering:  $r_p = 0.875(10)$  fm
2. Hydrogen spectroscopy:  $r_p = 0.8768(69)$  fm
3.  $\implies$  CODATA-2010:  $r_p = \mathbf{0.8775(51)}$  fm  
Mohr *et al.*, Rev. Mod. Phys. (2012)

### ● $r_p$ from $\mu$ H Lamb shift (2S-2P)

1.  $\mu$ H  $2S_{1/2}^{F=1} - 2P_{3/2}^{F=2}$ :  $r_p = 0.84184(67)$  fm ( $5\sigma$ )  
Pohl *et al.*, Nature (2010)
2. Combined with  
 $\mu$ H  $2S_{1/2}^{F=0} - 2P_{3/2}^{F=1}$ :  $r_p = \mathbf{0.84087(39)}$  fm ( $7\sigma$ )  
Antognini *et al.*, Science (2013)

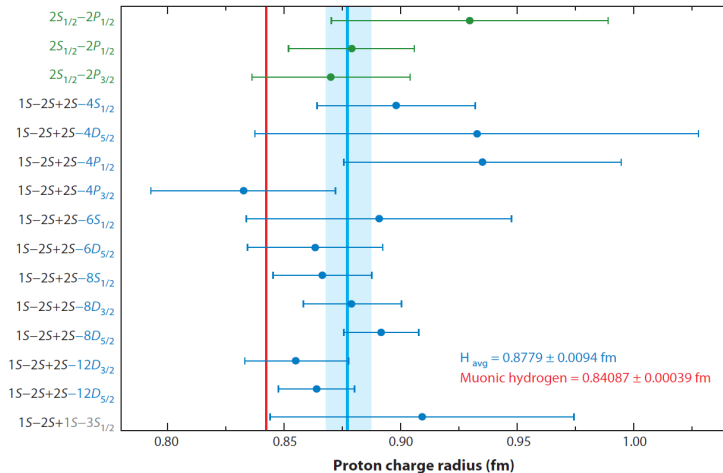


# Proton radius puzzle



# Origins of the discrepancy? — old data

## Hydrogen spectroscopy



Pohl *et al.*, *Annu. Rev. Nucl. Part. Sci.* '13

# Origins of the discrepancy? — old data

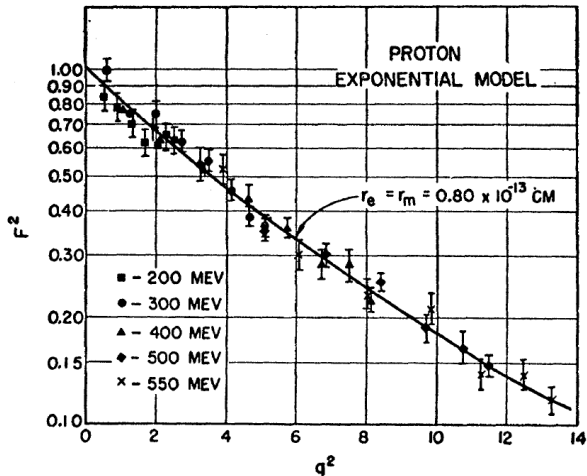
*ep* scattering measurements / fits?

( $Q^2$  not small enough / floating normalization)

$$G_E^p(Q^2) = 1 - \frac{1}{6} r_p^2 Q^2 + \dots$$

# Origins of the discrepancy? — old data

*ep* scattering measurements / fits?

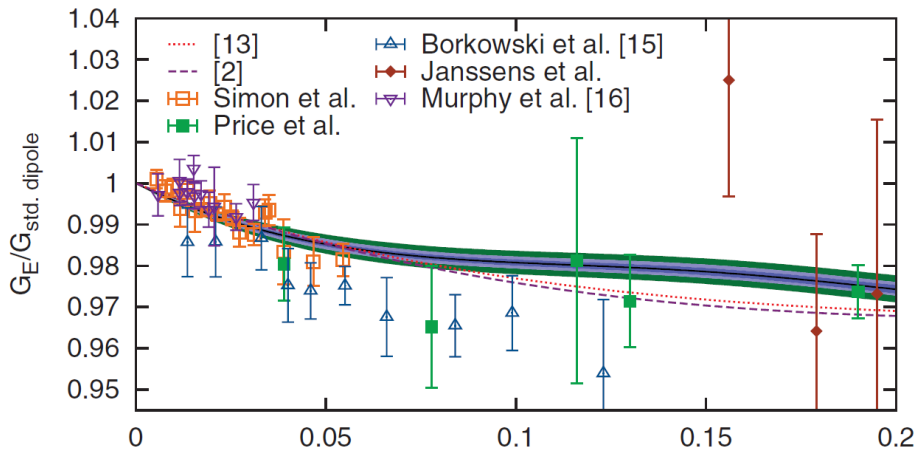


Chambers and Hofstadter, Phys. Rev. 103, 1454 (1956)



# Origins of the discrepancy? — old data

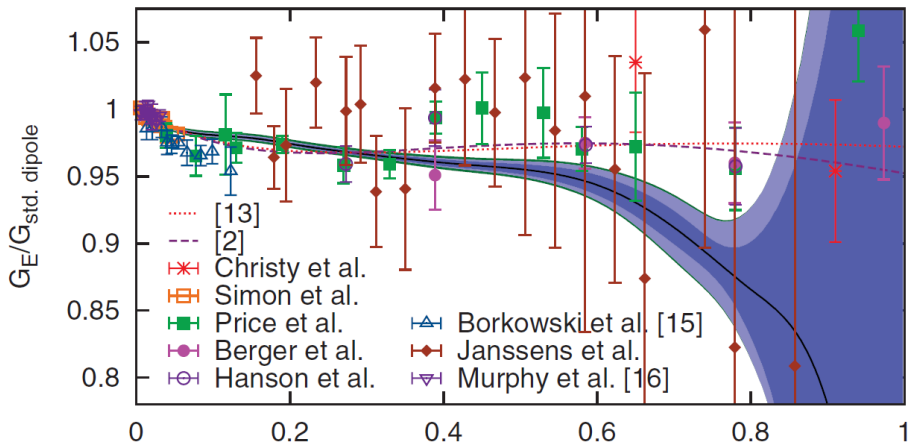
*ep* scattering measurements / fits?



Bernauer *et al.*, *Phys. Rev. Lett.* **105**, 242001 (2010)

# Origins of the discrepancy? — old data

*ep* scattering measurements / fits?

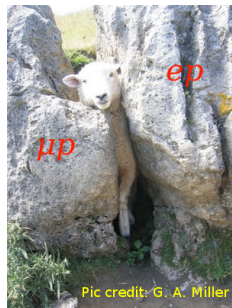


Bernauer *et al.*, Phys. Rev. Lett. **105**, 242001 (2010)

# Origins of the discrepancy? — new ideas

Study the discrepancy between  $r_p$  from  $ep$  and  $\mu p$  experiments

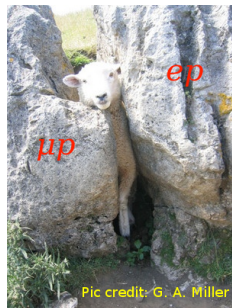
- Problems with electronic measurements ?
  - $eH$  spectroscopy
  - $ep$  scattering measurements / fits?



# Origins of the discrepancy? — new ideas

## Study the discrepancy between $r_p$ from $ep$ and $\mu p$ experiments

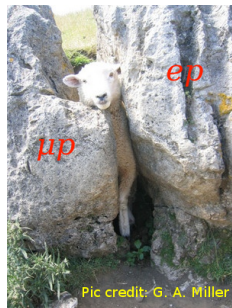
- Problems with electronic measurements ?
  - $eH$  spectroscopy
  - $ep$  scattering measurements / fits?
    - dispersion relations:  $r_p = 0.84(1)$ ,  $\chi_r^2 \approx 2.2$   
Lorenz *et al.*, EPJA '12



# Origins of the discrepancy? — new ideas

## Study the discrepancy between $r_p$ from $ep$ and $\mu p$ experiments

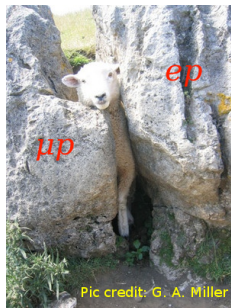
- Problems with electronic measurements ?
  - $eH$  spectroscopy
  - $ep$  scattering measurements / fits?
    - dispersion relations:  $r_p = 0.84(1)$ ,  $\chi_r^2 \approx 2.2$  1.4  
Lorenz *et al.*, EPJA '12; PLB '14; PRD '15



# Origins of the discrepancy? — new ideas

## Study the discrepancy between $r_p$ from $ep$ and $\mu p$ experiments

- Problems with electronic measurements ?
  - $eH$  spectroscopy
  - $ep$  scattering measurements / fits?
    - dispersion relations:  $r_p = 0.84(1)$ ,  $\chi_r^2 \approx 2.2$  1.4  
Lorenz *et al.*, EPJA '12; PLB '14; PRD '15

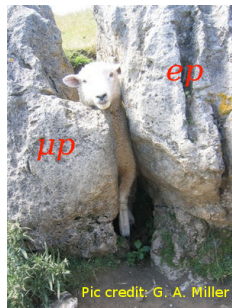


Parametrization	MAMI (1422 data points)	World data incl. MAMI (1922 data points)
Unconstrained $z$ expansion	$r_E = 0.64, r_M = 1.97, (\chi_r^2 = 1.12)$	$r_E = 0.85, r_M = 0.98, (\chi_r^2 = 1.17)$
$z$ expansion, $ e_k  < 10$	$r_E = 0.91, r_M = 0.79, (\chi_r^2 = 1.17)$	$r_E = 0.89, r_M = 0.77, (\chi_r^2 = 1.23)$
DR approach	$r_E = 0.84, r_M = 0.85, (\chi_r^2 = 1.41)$	$r_E = 0.84, r_M = 0.85, (\chi_r^2 = 1.32)$
Combination of the above	$r_E = 0.84, r_M = 0.85, (\chi_r^2 = 1.38)$	$r_E = 0.84, r_M = 0.85, (\chi_r^2 = 1.30)$

# Origins of the discrepancy? — new ideas

## Study the discrepancy between $r_p$ from $ep$ and $\mu p$ experiments

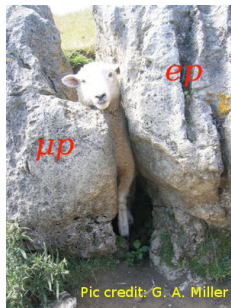
- Problems with electronic measurements ?
  - $eH$  spectroscopy
  - $ep$  scattering measurements / fits?
    - dispersion relations:  $r_p = 0.84(1)$ ,  $\chi_r^2 \approx 2.2$  1.4  
Lorenz *et al.*, EPJA '12; PLB '14; PRD '15
    - non-dipole fit:  $r_p = 0.849(7)$   
Adamuscin *et al.*, Prog. Part. Nucl. Phys. '12



# Origins of the discrepancy? — new ideas

## Study the discrepancy between $r_p$ from $ep$ and $\mu p$ experiments

- Problems with electronic measurements ?
  - $eH$  spectroscopy
  - $ep$  scattering measurements / fits?
    - dispersion relations:  $r_p = 0.84(1)$ ,  $\chi_r^2 \approx 2.2$  1.4  
Lorenz *et al.*, EPJA '12; PLB '14; PRD '15
    - non-dipole fit:  $r_p = 0.849(7)$   
Adamuscin *et al.*, Prog. Part. Nucl. Phys. '12
- exotic hadronic structure?
  - Birse & McGovern, EPJA '12 *vs.* Miller, PLB '13
  - Hill & Paz, PRD '10; PRL '11 & Jentschura PRA '13

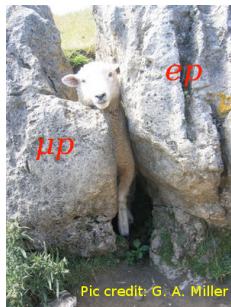




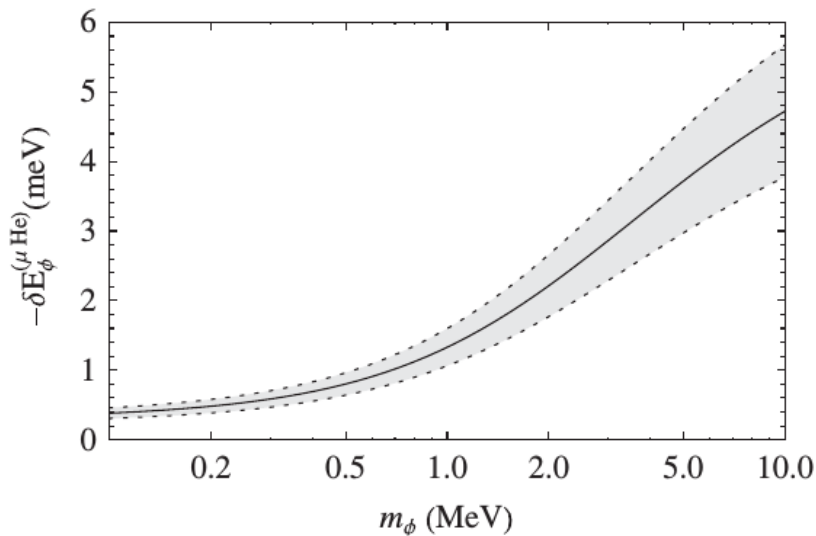
# Origins of the discrepancy? — new ideas

## Study the discrepancy between $r_p$ from $ep$ and $\mu p$ experiments

- Problems with electronic measurements ?
  - $eH$  spectroscopy
  - $ep$  scattering measurements / fits?
    - dispersion relations:  $r_p = 0.84(1)$ ,  $\chi_r^2 \approx 2.2$  1.4  
Lorenz *et al.*, EPJA '12; PLB '14; PRD '15
    - non-dipole fit:  $r_p = 0.849(7)$   
Adamuscin *et al.*, Prog. Part. Nucl. Phys. '12
- exotic hadronic structure?
  - Birse & McGovern, EPJA '12 *vs.* Miller, PLB '13
  - Hill & Paz, PRD '10; PRL '11 & Jentschura PRA '13
- beyond-standard-model physics?
  - new force carriers, e.g., dark photon: interact differently with  $e$  and  $\mu$
  - address both  $r_p$  puzzle &  $g_\mu$  puzzle  
Tucker-Smith & Yavin, PRD '11; Batell, McKeen & Pospelov, PRL '11;  
Carlson & Rislow, PRD '12; PRD '14



# Tucker-Smith & Yavin's prediction for $\mu^4\text{He}^+$



# Origins of the discrepancy?

## New experiments to shed light on the puzzle

- **Jefferson Lab**

- $ep$  scattering for  $Q^2$  from  $10^{-4} \text{ GeV}^2$  to  $10^{-2} \text{ GeV}^2$

# Origins of the discrepancy?

## New experiments to shed light on the puzzle

### ● Jefferson Lab

- $ep$  scattering for  $Q^2$  from  $10^{-4}$  GeV<sup>2</sup> to  $10^{-2}$  GeV<sup>2</sup>

### ● MUSE collaboration at PSI

- $\mu p$  scattering experiment (in development)
  - in the presence of both  $e$  &  $\mu$  beams: reduce systematic uncertainty
  - measure  $e^\pm p$  and  $\mu^\pm p$ : can study  $2\gamma$  exchange

# Origins of the discrepancy?

## New experiments to shed light on the puzzle

### ● Jefferson Lab

- $ep$  scattering for  $Q^2$  from  $10^{-4}$  GeV<sup>2</sup> to  $10^{-2}$  GeV<sup>2</sup>

### ● MUSE collaboration at PSI

- $\mu p$  scattering experiment (in development)
  - in the presence of both  $e$  &  $\mu$  beams: reduce systematic uncertainty
  - measure  $e^\pm p$  and  $\mu^\pm p$ : can study  $2\gamma$  exchange

### ● “Ordinary” Hydrogen

- and other electronic systems — Rydberg const. ...

# Origins of the discrepancy?

## New experiments to shed light on the puzzle

### ● Jefferson Lab

- $ep$  scattering for  $Q^2$  from  $10^{-4}$  GeV<sup>2</sup> to  $10^{-2}$  GeV<sup>2</sup>

### ● MUSE collaboration at PSI

- $\mu p$  scattering experiment (in development)
  - in the presence of both  $e$  &  $\mu$  beams: reduce systematic uncertainty
  - measure  $e^\pm p$  and  $\mu^\pm p$ : can study  $2\gamma$  exchange

### ● “Ordinary” Hydrogen

- and other electronic systems — Rydberg const. ...

### ● CREMA collaboration at PSI

- Lamb shift (2S-2P) & isotope shift (1S-2S) in  $\mu\text{D}$  (finishing)
- Lamb shift in  $\mu^4\text{He}^+$  and  $\mu^3\text{He}^+$  (both measured in 2014)

# Origins of the discrepancy?

## New experiments to shed light on the puzzle

### ● Jefferson Lab

- $ep$  scattering for  $Q^2$  from  $10^{-4}$  GeV<sup>2</sup> to  $10^{-2}$  GeV<sup>2</sup>

### ● MUSE collaboration at PSI

- $\mu p$  scattering experiment (in development)
  - in the presence of both  $e$  &  $\mu$  beams: reduce systematic uncertainty
  - measure  $e^\pm p$  and  $\mu^\pm p$ : can study  $2\gamma$  exchange

### ● “Ordinary” Hydrogen

- and other electronic systems — Rydberg const. ...

### ● CREMA collaboration at PSI

- Lamb shift (2S-2P) & isotope shift (1S-2S) in  $\mu\text{D}$  (finishing)
- Lamb shift in  $\mu^4\text{He}^+$  and  $\mu^3\text{He}^+$  (both measured in 2014)

high-precision measurements  $\iff$  accurate theoretical inputs

# Extract nuclear charge radius

$\langle r^2 \rangle$  from Lamb shift

$$\Delta E_{LS} = \delta_{QED} + \delta_{size}(r) + \delta_{pol}$$

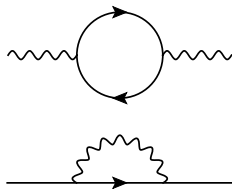


# Extract nuclear charge radius

## $\langle r^2 \rangle$ from Lamb shift

$$\Delta E_{LS} = \delta_{QED} + \delta_{size}(r) + \delta_{pol}$$

- QED corrections:
  - vacuum polarization
  - lepton self energy
  - relativistic recoil effects
- Calculations for  $\mu^{3,4}\text{He}^+$  reexamined



# Extract nuclear charge radius

## $\langle r^2 \rangle$ from Lamb shift

$$\Delta E_{LS} = \delta_{QED} + \delta_{size}(r) + \delta_{pol}$$

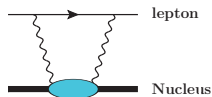
- Nuclear finite-size corrections (elastic):
  - leading term:  $\frac{m_r^3}{12} (Z\alpha)^4 \langle r^2 \rangle$
  - 3rd Zemach moment:  $-\frac{m_r^4}{24} (Z\alpha)^5 \langle r^3 \rangle_{(2)} \propto \langle r^2 \rangle^{3/2}$

# Extract nuclear charge radius

## $\langle r^2 \rangle$ from Lamb shift

$$\Delta E_{LS} = \delta_{QED} + \delta_{size}(r) + \delta_{pol}$$

- Nuclear **A** polarization corrections (inelastic):
  - exchange of two virtual photons
  - dominant contribution  $\sim (Z\alpha)^5$
- Nucleon **p/n** polarization corrections (inelastic)



# Uncertainty in nuclear polarization

$$\Delta E_{LS} = \delta_{QED} + \delta_{size}(r) + \delta_{pol}$$

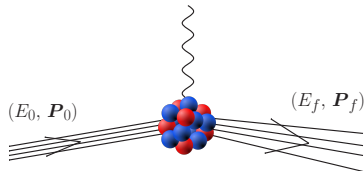
- The accuracy in determining  $\langle r^2 \rangle$  relies on the accuracy of  $\delta_{pol}$

# Uncertainty in nuclear polarization

$$\Delta E_{LS} = \delta_{QED} + \delta_{size}(r) + \delta_{pol}$$

- The accuracy in determining  $\langle r^2 \rangle$  relies on the accuracy of  $\delta_{pol}$
- Nuclear polarization  $\implies$  inputs from nuclear response functions

$$S_O(\omega) = \sum_f |\langle \psi_f | \hat{O} | \psi_0 \rangle|^2 \delta(E_f - E_0 - \omega)$$

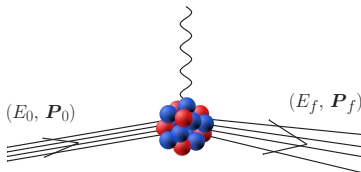


# Uncertainty in nuclear polarization

$$\Delta E_{LS} = \delta_{QED} + \delta_{size}(r) + \delta_{pol}$$

- The accuracy in determining  $\langle r^2 \rangle$  relies on the accuracy of  $\delta_{pol}$
- Nuclear polarization  $\implies$  inputs from nuclear response functions

$$S_O(\omega) = \sum_f |\langle \psi_f | \hat{O} | \psi_0 \rangle|^2 \delta(E_f - E_0 - \omega)$$



- Early calculations of  $\delta_{pol}$  in muonic atoms:  
 $\implies S_O(\omega)$  inputs were not accurate enough

# Previous calculations of $S_O(\omega)$ & $\delta_{\text{pol}}$

## • $\mu\text{D}$

- Modern forces: AV14 - Leidemann & Rosenfelder '95
- State-of-the-art: AV18 - Pachucki '11
  - underestimates nuclear physics uncertainty
  - includes nucleon polarizability  $\delta_{\text{pol}}^p$  (incorrect?)
- $\not\propto$ EFT: zero-range expansion - Friar '13
  - estimated uncertainty 1–2%
  - includes nucleon-size corrections (incorrect?)
- From scattering: Carlson, Gorchtein, Vanderhaeghen '14
  - estimate nuclear uncertainty  $\sim 47\%$

# Previous calculations of $S_O(\omega)$ & $\delta_{\text{pol}}$

## • $\mu\text{D}$

- Modern forces: AV14 - Leidemann & Rosenfelder '95
- State-of-the-art: AV18 - Pachucki '11
  - underestimates nuclear physics uncertainty
  - includes nucleon polarizability  $\delta_{\text{pol}}^p$  (incorrect?)
- $\not\chi\text{EFT}$ : zero-range expansion - Friar '13
  - estimated uncertainty 1–2%
  - includes nucleon-size corrections (incorrect?)
- From scattering: Carlson, Gorchtein, Vanderhaeghen '14
  - estimate nuclear uncertainty  $\sim 47\%$

## • $\mu^{3,4}\text{He}^+$

- From photoabsorption: Bernabeu & Jarlskog '74; Rinker '76; Friar '77
- $\delta_{\text{pol}}^{A=3/4} = -4.9 / -3.1 \text{ meV} \pm 20\%$  (c.f. experimental requirement  $\sim \pm 5\%$ )



# Previous calculations of $S_O(\omega)$ & $\delta_{\text{pol}}$

## • $\mu\text{D}$

- Modern forces: AV14 - Leidemann & Rosenfelder '95
- State-of-the-art: AV18 - Pachucki '11
  - underestimates nuclear physics uncertainty
  - includes nucleon polarizability  $\delta_{\text{pol}}^p$  (incorrect?)
- $\not\propto$ EFT: zero-range expansion - Friar '13
  - estimated uncertainty 1–2%
  - includes nucleon-size corrections (incorrect?)
- From scattering: Carlson, Gorchtein, Vanderhaeghen '14
  - estimate nuclear uncertainty  $\sim 47\%$

## • $\mu^{3,4}\text{He}^+$

- From photoabsorption: Bernabeu & Jarlskog '74; Rinker '76; Friar '77
- $\delta_{\text{pol}}^{A=3/4} = -4.9 / -3.1 \text{ meV} \pm 20\%$  (c.f. experimental requirement  $\sim \pm 5\%$ )

## • $\mu\text{T}$

- Very crude theoretical estimation: C. Joachain '61, in French

# Previous calculations of $S_O(\omega)$ & $\delta_{\text{pol}}$

## • $\mu\text{D}$

- Modern forces: AV14 - Leidemann & Rosenfelder '95
- State-of-the-art: AV18 - Pachucki '11
  - underestimates nuclear physics uncertainty
  - includes nucleon polarizability  $\delta_{\text{pol}}^p$  (incorrect?)
- $\not\propto$ EFT: zero-range expansion - Friar '13
  - estimated uncertainty 1–2%
  - includes nucleon-size corrections (incorrect?)
- From scattering: Carlson, Gorchtein, Vanderhaeghen '14
  - estimate nuclear uncertainty  $\sim 47\%$

## • $\mu^{3,4}\text{He}^+$

- From photoabsorption: Bernabeu & Jarlskog '74; Rinker '76; Friar '77
- $\delta_{\text{pol}}^{A=3/4} = -4.9 / -3.1 \text{ meV} \pm 20\%$  (c.f. experimental requirement  $\sim \pm 5\%$ )

## • $\mu\text{T}$

- Very crude theoretical estimation: C. Joachain '61, in French

## • Status of $\delta_{\text{pol}}$ in light muonic atoms

- **experimental input** for  $S_O$  is unsatisfactory

# Previous calculations of $S_O(\omega)$ & $\delta_{\text{pol}}$

## • $\mu\text{D}$

- Modern forces: AV14 - Leidemann & Rosenfelder '95
- State-of-the-art: AV18 - Pachucki '11
  - underestimates nuclear physics uncertainty
  - includes nucleon polarizability  $\delta_{\text{pol}}^p$  (incorrect?)
- $\neq$ EFT: zero-range expansion - Friar '13
  - estimated uncertainty 1–2%
  - includes nucleon-size corrections (incorrect?)
- From scattering: Carlson, Gorchtein, Vanderhaeghen '14
  - estimate nuclear uncertainty  $\sim 47\%$

## • $\mu^{3,4}\text{He}^+$

- From photoabsorption: Bernabeu & Jarlskog '74; Rinker '76; Friar '77
- $\delta_{\text{pol}}^{A=3/4} = -4.9 / -3.1 \text{ meV} \pm 20\%$  (c.f. experimental requirement  $\sim \pm 5\%$ )

## • $\mu\text{T}$

- Very crude theoretical estimation: C. Joachain '61, in French

## • Status of $\delta_{\text{pol}}$ in light muonic atoms

- **experimental input** for  $S_O$  is unsatisfactory
- need to calculate  $\delta_{\text{pol}}$  using **modern potentials and *ab-initio* methods**

# Nuclear polarization in $\mu^4\text{He}^+$

We have performed the first *ab-initio* calculation of the nuclear polarization and 3rd Zemach moment in  $\mu^4\text{He}^+$  with state-of-the-art forces:

AV18+UIX and  $\chi\text{EFT}$

C. Ji, NND, S. Bacca & N. Barnea, PRL (2013); FBS (2014).

# Nuclear polarization in $\mu^4\text{He}^+$

We have performed the first *ab-initio* calculation of the nuclear polarization and 3rd Zemach moment in  $\mu^4\text{He}^+$  with state-of-the-art forces:

AV18+UIX and  $\chi\text{EFT}$

C. Ji, NND, S. Bacca & N. Barnea, PRL (2013); FBS (2014).

- **Few-body methods**

- **EIHH**: Effective Interaction Hyperspherical Harmonics
- **LIT**: Lorentz Integral Transform
- **LSR**: A new method we developed based on the Lanczos algorithm

# Nuclear polarization in $\mu^4\text{He}^+$

We have performed the first *ab-initio* calculation of the nuclear polarization and 3rd Zemach moment in  $\mu^4\text{He}^+$  with state-of-the-art forces:

AV18+UIX and  $\chi\text{EFT}$

C. Ji, NND, S. Bacca & N. Barnea, PRL (2013); FBS (2014).

- **Few-body methods**

- **EIHH**: Effective Interaction Hyperspherical Harmonics
- **LIT**: Lorentz Integral Transform
- **LSR**: A new method we developed based on the Lanczos algorithm

- **Calculation outline**

$$H_{nucl} \implies S_O \implies \delta_{pol}$$

# Nuclear polarization in $\mu^4\text{He}^+$

We have performed the first *ab-initio* calculation of the nuclear polarization and 3rd Zemach moment in  $\mu^4\text{He}^+$  with state-of-the-art forces:

AV18+UIX and  $\chi\text{EFT}$

C. Ji, NND, S. Bacca & N. Barnea, PRL (2013); FBS (2014).

- **Few-body methods**

- **EIHH**: Effective Interaction Hyperspherical Harmonics
- **LIT**: Lorentz Integral Transform
- **LSR**: A new method we developed based on the Lanczos algorithm

- **Calculation outline**

$$H_{nucl} \implies S_O \implies \delta_{pol}$$

- **Error estimation**

- $\delta_{pol}$  from different  $H_{nucl} \implies$  nuclear physics uncertainty
- also estimate numerical error and higher order corrections

# Nuclear polarization in $\mu^4\text{He}^+$

We have performed the first *ab-initio* calculation of the nuclear polarization and 3rd Zemach moment in  $\mu^4\text{He}^+$  with state-of-the-art forces:

AV18+UIX and  $\chi\text{EFT}$

C. Ji, NND, S. Bacca & N. Barnea, PRL (2013); FBS (2014).

- **Few-body methods**

- **EIHH**: Effective Interaction Hyperspherical Harmonics
- **LIT**: Lorentz Integral Transform
- **LSR**: A new method we developed based on the Lanczos algorithm

- **Calculation outline**

$$H_{nucl} \implies S_O \implies \delta_{pol}$$

- **Error estimation**

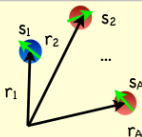
- $\delta_{pol}$  from different  $H_{nucl} \implies$  nuclear physics uncertainty
- also estimate numerical error and higher order corrections

- **Our Goal**

provide  $\delta_{pol}$  with accuracy comparable to the  $\pm 5\%$  experimental needs



# Nuclear potentials: two approaches



$$H_N |\psi_i\rangle = E_i |\psi_i\rangle$$

$$H_N = T + V_{NN} + V_{3N} + \dots$$

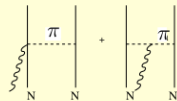
High precision two-nucleon potentials:  
well constrained on NN phase shifts

Three nucleon forces:  
less known, constraint on  $A > 2$  observables

Traditional Nuclear Physics  
AV18+UIX, ...,  $J_2$

Effective Field Theory  
 $N^2LO, N^3LO \dots$

$$J^\mu = J_N^\mu + J_{NN}^\mu + \dots$$



two-body currents (or MEC)  
subnuclear d.o.f.

$$J^\mu \text{ consistent with } V$$

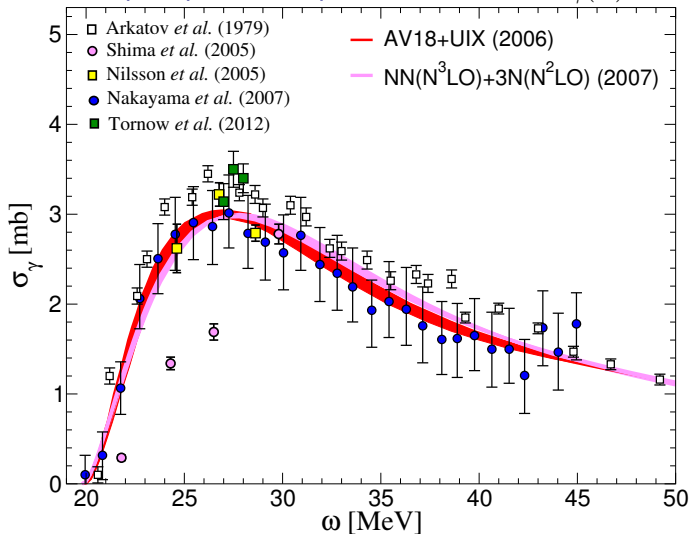
$$\nabla \cdot J = -i[V, \rho]$$

$$S(\omega) \propto |\langle \psi_f | J^\mu | \psi_0 \rangle|^2$$

Exact Initial state &  
Final state in the continuum at  
different energies and for different  $A$

# Nuclear potentials: $^4\text{He}$ Photoabsorption

electric dipole photoabsorption cross section  $\sigma_\gamma(\omega) = 4\pi^2\alpha\omega S_{D1}(\omega)$

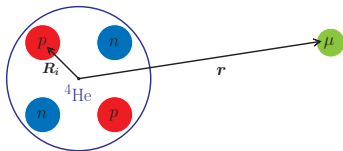


# Nuclear polarization: basic idea

- Hamiltonian for muonic atoms

$$H = H_{nucl} + H_{\mu} + \Delta H$$

$$H_{\mu} = \frac{p^2}{2m_r} - \frac{Z\alpha}{r}$$



- Corrections to the point Coulomb from protons

$$\Delta H = \alpha \sum_i^Z \Delta V(\mathbf{r}, \mathbf{R}_i) \equiv \alpha \sum_i^Z \left( \frac{1}{r} - \frac{1}{|\mathbf{r} - \mathbf{R}_i|} \right)$$

- Evaluate inelastic effects of  $\Delta H$  on muonic spectrum in  $2^{nd}$ -order perturbation theory

$$\delta_{pol} = \sum_{N \neq N_0, \mu} \langle N_0 \mu_0 | \Delta H | N \mu \rangle \frac{1}{E_{N_0} - E_N + \epsilon_{\mu_0} - \epsilon_{\mu}} \langle N \mu | \Delta H | \mu_0 N_0 \rangle$$

$|\mu_0\rangle$ : muon wave function for  $2S/2P$  state

# Nuclear polarization: contributions

## Systematic contributions to nuclear polarization

$\delta_{NR}$       **Non-Relativistic** limit

$\delta_L + \delta_T$       Longitudinal and **T**ransverse **relativistic** corrections

$\delta_C$       **Coulomb** distortions

$\delta_{NS}$       Corrections from **finite Nucleon Size**

# LSR: Lanczos sum rule method

- Nuclear polarization  $\implies$  energy-dependent sum rules of the response functions

$$\delta_{\text{pol}} \propto I_O = \int_0^{\infty} d\omega S_O(\omega) g(\omega)$$

# LSR: Lanczos sum rule method

- Nuclear polarization  $\implies$  energy-dependent sum rules of the response functions

$$\delta_{\text{pol}} \propto I_O = \int_0^{\infty} d\omega S_O(\omega) g(\omega)$$

- The Lanczos algorithm is an efficient method calculate the LIT of  $S_O$ .  
It can be extended to calculate the sum rules.

# LSR: Lanczos sum rule method

- Nuclear polarization  $\implies$  energy-dependent sum rules of the response functions

$$\delta_{\text{pol}} \propto I_O = \int_0^{\infty} d\omega S_O(\omega) g(\omega)$$

- The Lanczos algorithm is an efficient method calculate the LIT of  $S_O$ . It can be extended to calculate the sum rules.
- With the Lanczos sum rule (LSR) method, we directly calculate  $I_O$ , without explicitly solving  $S_O$ .

# LSR: Lanczos sum rule method

- Nuclear polarization  $\implies$  energy-dependent sum rules of the response functions

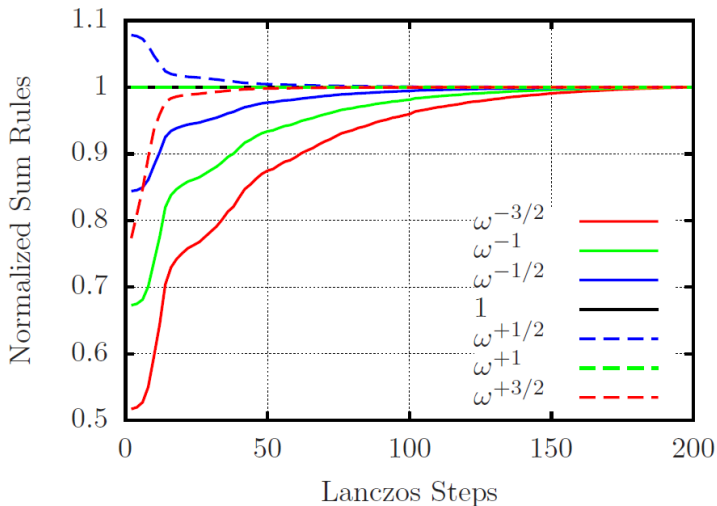
$$\delta_{\text{pol}} \propto I_O = \int_0^\infty d\omega S_O(\omega) g(\omega)$$

- The Lanczos algorithm is an efficient method calculate the LIT of  $S_O$ . It can be extended to calculate the sum rules.
- With the Lanczos sum rule (LSR) method, we directly calculate  $I_O$ , without explicitly solving  $S_O$ .
- The calculated  $I_O$  converges as the LIT of  $S_O$ , if  $g(\omega)$  is smooth.

NND, Ji, Bacca, Barnea, Phys. Rev. C **89**, 064317 (2014)



# LSR: Example — ${}^4\text{He}$ dipole response integrals



(Model space size  $M \sim 10^5$ )

NND, Barnea, Ji, and Bacca, PRC (2014)

# Nuclear polarization in $\mu^4\text{He}^+$

[meV]		AV18+UIX	$\chi^{\text{EFT}} \star$
$\delta^{(0)}$	$\delta_{D1}^{(0)}$	-4.418	-4.701
	$\delta_L^{(0)}$	0.289	0.308
	$\delta_T^{(0)}$	-0.126	-0.134
	$\delta_C^{(0)}$	0.512	0.546

$\star$   $NN$ :  $N^3\text{LO-EM}$   
 $3N$ :  $N^2\text{LO}$  ( $c_D=1$ ,  $c_E=-0.029$ )

# Nuclear polarization in $\mu^4\text{He}^+$

[meV]		AV18+UIX	$\chi^{\text{EFT}} \star$
$\delta^{(0)}$	$\delta_{D1}^{(0)}$	-4.418	-4.701
	$\delta_L^{(0)}$	0.289	0.308
	$\delta_T^{(0)}$	-0.126	-0.134
	$\delta_C^{(0)}$	0.512	0.546
$\delta^{(1)}$	$\delta_{R3pp}^{(1)}$	-3.442	-3.717
	$\delta_{Z3}^{(1)}$	4.183	4.526

$\star$  NN: N<sup>3</sup>LO-EM  
3N: N<sup>2</sup>LO ( $c_D=1$ ,  $c_E=-0.029$ )

# Nuclear polarization in $\mu^4\text{He}^+$

[meV]		AV18+UIX	$\chi^{\text{EFT}} \star$
$\delta^{(0)}$	$\delta_{D1}^{(0)}$	-4.418	-4.701
	$\delta_L^{(0)}$	0.289	0.308
	$\delta_T^{(0)}$	-0.126	-0.134
	$\delta_C^{(0)}$	0.512	0.546
$\delta^{(1)}$	$\delta_{R3pp}^{(1)}$	-3.442	-3.717
	$\delta_{Z3}^{(1)}$	4.183	4.526
$\delta^{(2)}$	$\delta_{R2}^{(2)}$	0.259	0.324
	$\delta_Q^{(2)}$	0.484	0.561
	$\delta_{D1D3}^{(2)}$	-0.666	-0.784

$\star$  NN: N<sup>3</sup>LO-EM

3N: N<sup>2</sup>LO ( $c_D=1$ ,  $c_E=-0.029$ )

# Nuclear polarization in $\mu^4\text{He}^+$

[meV]		AV18+UIX	$\chi^{\text{EFT}} \star$
$\delta^{(0)}$	$\delta_{D1}^{(0)}$	-4.418	-4.701
	$\delta_L^{(0)}$	0.289	0.308
	$\delta_T^{(0)}$	-0.126	-0.134
	$\delta_C^{(0)}$	0.512	0.546
$\delta^{(1)}$	$\delta_{R3pp}^{(1)}$	-3.442	-3.717
	$\delta_{Z3}^{(1)}$	4.183	4.526
$\delta^{(2)}$	$\delta_{R2}^{(2)}$	0.259	0.324
	$\delta_Q^{(2)}$	0.484	0.561
	$\delta_{D1D3}^{(2)}$	-0.666	-0.784
$\delta_{NS}$	$\delta_{R1pp}^{(1)}$	-1.036	-1.071
	$\delta_{Z1}^{(1)}$	1.753	1.811
	$\delta_{NS}^{(2)}$	-0.200	-0.210

$\star$  NN:  $\text{N}^3\text{LO-EM}$   
3N:  $\text{N}^2\text{LO}$  ( $c_D=1$ ,  $c_E=-0.029$ )

# Nuclear polarization in $\mu^4\text{He}^+$

[meV]		AV18+UIX	$\chi^{\text{EFT}} \star$
$\delta^{(0)}$	$\delta_{D1}^{(0)}$	-4.418	-4.701
	$\delta_L^{(0)}$	0.289	0.308
	$\delta_T^{(0)}$	-0.126	-0.134
	$\delta_C^{(0)}$	0.512	0.546
$\delta^{(1)}$	$\delta_{R3pp}^{(1)}$	-3.442	-3.717
	$\delta_{Z3}^{(1)}$	4.183	4.526
$\delta^{(2)}$	$\delta_{R2}^{(2)}$	0.259	0.324
	$\delta_Q^{(2)}$	0.484	0.561
	$\delta_{D1D3}^{(2)}$	-0.666	-0.784
$\delta_{NS}$	$\delta_{R1pp}^{(1)}$	-1.036	-1.071
	$\delta_{Z1}^{(1)}$	1.753	1.811
	$\delta_{NS}^{(2)}$	-0.200	-0.210
$\delta_{\text{pol}}$		-2.408	-2.542

$\star$  NN:  $\text{N}^3\text{LO-EM}$   
 3N:  $\text{N}^2\text{LO}$  ( $c_D=1$ ,  $c_E=-0.029$ )

# Nuclear polarization in $\mu^4\text{He}^+$

[meV]	AV18+UIX	$\chi^{\text{EFT}} \star$
$\delta^{(0)}$	-3.743	-3.981
$\delta^{(1)}$	0.741	0.809
$\delta^{(2)}$	0.077	0.101
$\delta_{NS}$	0.517	0.530
$\delta_{\text{pol}}$	-2.408	-2.542

- Convergence from  $\delta^{(0)}$  to  $\delta^{(2)}$  in a systematic expansion of  $\sqrt{2m_r\omega}|\mathbf{R} - \mathbf{R}'| \sim \sqrt{m_r/M_N} \approx 0.3$

★ NN: N<sup>3</sup>LO-EM  
3N: N<sup>2</sup>LO ( $c_D=1$ ,  $c_E=-0.029$ )

# Nuclear polarization in $\mu^4\text{He}^+$

[meV]	AV18+UIX	$\chi\text{EFT}^\star$
$\delta^{(0)}$	-3.743	-3.981
$\delta^{(1)}$	0.741	0.809
$\delta^{(2)}$	0.077	0.101
$\delta_{NS}$	0.517	0.530
$\delta_{\text{pol}}$	-2.408	-2.542

- Convergence from  $\delta^{(0)}$  to  $\delta^{(2)}$  in a systematic expansion of  $\sqrt{2m_r\omega}|\mathbf{R} - \mathbf{R}'| \sim \sqrt{m_r/M_N} \approx 0.3$
- AV18+UIX &  $\chi\text{EFT}$  differ:  $\sim 5.5\%$  (0.134 meV)

★ NN: N<sup>3</sup>LO-EM  
3N: N<sup>2</sup>LO ( $c_D=1$ ,  $c_E=-0.029$ )



# Nuclear physics uncertainty

${}^4\text{He}$ observable		AV18+UIX	$\chi\text{EFT}$	Difference
binding energy	$B_0$ [MeV]	28.422	28.343	0.28%
point-proton nuclear radius	$R_{pp}$ [fm]	1.432	1.475	3.0%
electric-dipole polarizability	$\alpha_E$ [fm <sup>3</sup> ]	0.0651	0.0694	6.4%
$\mu^4\text{He}^+$ nuclear polarization	$\delta_{\text{pol}}$ [meV]	-2.408	-2.542	5.5%

- $B_0$ ,  $R_{pp}$  &  $\alpha_E$  in good agreement with previous calculations  
*Kievsky et al. '08, Gazit et al. '06 & Stetcu et al. '09*

- systematic uncertainty in  $\delta_{\text{pol}}$  from nuclear physics:

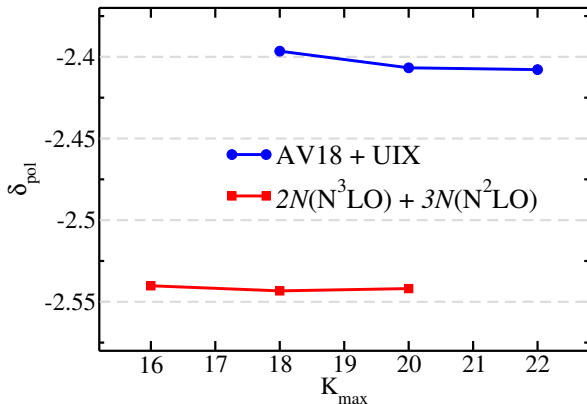
$$\frac{5.5\%}{\sqrt{2}} \implies \pm 4\% (1\sigma)$$

# Numerical accuracy

- **Convergence with model space size**

Compare  $\delta_{\text{pol}}^{(K_{\text{max}})}$  with  $\delta_{\text{pol}}^{(K_{\text{max}}-4)}$

- AV18+UIX  $\sim 0.4\%$
- $\chi\text{EFT} \sim 0.2\%$



# Error budget — $\mu^4\text{He}^+$

- **Nuclear physics**

4% from two potentials

- **Numerical accuracy**

0.4% from convergence

- **Additional corrections**

- $(Z\alpha)^6$  terms (beyond 2nd-order perturbation theory)

- Rel. & Coulomb corrections (other than dipole)

- higher-order nucleon-size corrections

⇒  $\sim 4\%$  estimated from additional corrections

- **Final result (quadratic sum)**

our prediction:  $\delta_{\text{pol}} = -2.47 \text{ meV} \pm 6\%$

previous estimates:  $\delta_{\text{pol}} = -3.1 \text{ meV} \pm 20\%$

experimental needs:  $\delta_{\text{pol}}$  uncertainty  $\sim 5\%$

# Reflections on $\mu^4\text{He}^+$

- The accuracy of  $\delta_{\text{pol}}$  in  $\mu^4\text{He}^+$  is limited by the nuclear physics (AV18+UIX vs.  $\chi\text{EFT-EM}$ )
- To study this we can further vary the nuclear potentials
  - Use  $\chi\text{EFT}$  at different orders to track the convergence
  - At each order vary the cutoff to estimate the theoretical error ( $\chi\text{EFT-EGM}$ : Epelbaum, Glöckle, Meißner, NPA '05)

# Reflections on $\mu\text{D}$ — See Javier Hernandez

- Pachucki only used AV18
  - ⇒ No nuclear physics uncertainty
  - ⇒ We can add  $\chi^{\text{EFT-EM}}$  &  $\chi^{\text{EFT-EGM}}$
- 2-body problem
  - ⇒ only  $NN$  interaction
  - ⇒ simple numerics
- Other issues
  - Pachucki did not include nucleon-size corrections
  - Pachucki did not treat nucleon-polarization correctly
  - We already reproduced Pachucki's leading term as a check for  $\mu^4\text{He}^+$
  - There is also a Magnetic contribution
  - It is beneficial to calculate  $\delta^\star \equiv |\delta_{\text{pol}} + \delta_{\text{Zem}}|$

# Work in progress

The work is not completed yet ...



# Preliminary: $\delta_{\text{pol}}$ in $\mu^3\text{He}^+$

[meV]	AV18+UIX	$\chi\text{EFT}^\star$
$\delta^{(0)}$	-5.361	-5.469
$\delta^{(1)}$	-0.460	-0.396
$\delta^{(2)}$	0.841	0.930
$\delta_{NS}$	0.797	0.805
$\delta_{Mag}$	0.078	0.047
$\delta_{\text{pol}}$	-4.104	-4.084

● Convergence from  $\delta^{(0)}$  to  $\delta^{(2)}$  ?

● AV18+UIX &  $\chi\text{EFT}$  agree:  
< 1% ( $\sim 0.02$  meV)

*c.f.*  $\mu^4\text{He}^+$ :

5.5% (0.134 meV)

★ NN:  $\text{N}^3\text{LO-EM}$   
3N:  $\text{N}^2\text{LO}$  ( $c_D=1$ ,  $c_E=-0.029$ )

# Preliminary: $\delta_{\text{pol}}$ in $\mu\text{T}$

[meV]	AV18+UIX	$\chi\text{EFT}^{\star}$
$\delta^{(0)}$	-0.680	-0.695
$\delta^{(1)}$	0.178	0.184
$\delta^{(2)}$	-0.025	-0.029
$\delta_{NS}$	0.053	0.054
$\delta_{Mag}$	0.010	0.006
$\delta_{\text{pol}}$	-0.465	-0.480

- Convergence from  $\delta^{(0)}$  to  $\delta^{(2)}$  ?
- AV18+UIX &  $\chi\text{EFT}$  agree:  
 $\sim 3\%$  ( $\sim 0.02$  meV)
- $\delta_{\text{pol}}^N$  of the nucleons ?
- Precise triton radius

★ NN: N<sup>3</sup>LO-EM  
3N: N<sup>2</sup>LO ( $c_D=1$ ,  $c_E=-0.029$ )



# Summary

- **Lamb shifts in muonic atoms**

- raise interesting questions about lepton universality
- probe isospin dependence of the proton radius puzzle
- allow high precision determination of the nuclear charge radius  $\langle r^2 \rangle$
- For  $A > 1$  the precision of  $\langle r^2 \rangle$  is bound by the nuclear polarization  $\delta_{\text{pol}}^A$

- **We perform the first *ab-initio* calculation of  $\delta^\star = \left| \delta_{\text{pol}}^A + \delta_{\text{Zem}} \right|$  in  $\mu^{3,4}\text{He}^+$  &  $\mu\text{T}$ , and improve the nuclear uncertainty in  $\mu\text{D}$**

$$\mu\text{D} \quad \delta^\star = 1.66(2) \text{ meV} \quad [\text{PLB } \mathbf{736}, 344 \text{ (2014)}]$$

$$\mu\text{T} \quad \delta^\star = 0.70(2) \text{ meV} \quad (\text{preliminary})$$

$$\mu^3\text{He}^+ \quad \delta^\star = 14.5(4) \text{ meV} \quad (\text{preliminary})$$

$$\mu^4\text{He}^+ \quad \delta^\star = 8.6(3) \text{ meV} \quad [\text{PRL } \mathbf{111}, 143402 \text{ (2013)}]$$

- more accurate than previous calculations
- will significantly improve the precision of  $\langle r^2 \rangle$  extracted from ongoing  $\mu^{3,4}\text{He}^+$  Lamb shift measurements

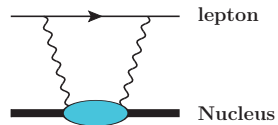
- Study higher-order terms
- Reduce nuclear physics uncertainty
  - understand why various nuclear potentials differ
  - further explore the various parameterizations (3NF?)
  - include higher-order or otherwise improved  $\chi$ EFT forces
- Investigate nuclear polarization in e.g.  $\mu^6\text{Li}^{+2}$ ,  $\mu^6\text{He}^+$ , ...
- Investigate nuclear polarization in HFS of electronic and muonic atoms

# The muonic Lamb shift is still puzzling !!!



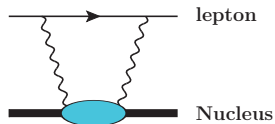
# Non-relativistic limit

- Neglect Coulomb interactions in the intermediate state



# Non-relativistic limit

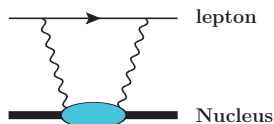
- Neglect Coulomb interactions in the intermediate state
- Expand muon matrix element in powers of  $\sqrt{2m_r\omega}|R - R'|$



# Non-relativistic limit

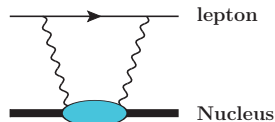
- Neglect Coulomb interactions in the intermediate state
- Expand muon matrix element in powers of  $\sqrt{2m_r\omega}|\mathbf{R} - \mathbf{R}'|$

- $|\mathbf{R} - \mathbf{R}'| \implies$  “virtual” distance the proton travels in  $2\gamma$  exchange
- uncertainty principal  $|\mathbf{R} - \mathbf{R}'| \sim 1/\sqrt{2m_N\omega}$
- $\sqrt{2m_r\omega}|\mathbf{R} - \mathbf{R}'| \sim \sqrt{\frac{m_r}{m_N}} \approx 0.3$



# Non-relativistic limit

- Neglect Coulomb interactions in the intermediate state
- Expand muon matrix element in powers of  $\sqrt{2m_r\omega}|\mathbf{R} - \mathbf{R}'|$



- $|\mathbf{R} - \mathbf{R}'| \implies$  “virtual” distance the proton travels in  $2\gamma$  exchange
- uncertainty principal  $|\mathbf{R} - \mathbf{R}'| \sim 1/\sqrt{2m_N\omega}$
- $\sqrt{2m_r\omega}|\mathbf{R} - \mathbf{R}'| \sim \sqrt{\frac{m_r}{m_N}} \approx 0.3$

$$P \simeq \frac{m_r^3(Z\alpha)^5}{12} \sqrt{\frac{2m_r}{\omega}} \left[ |\mathbf{R} - \mathbf{R}'|^2 - \frac{\sqrt{2m_r\omega}}{4} |\mathbf{R} - \mathbf{R}'|^3 + \frac{m_r\omega}{10} |\mathbf{R} - \mathbf{R}'|^4 \right]$$

$$\delta_{NR} = \delta_{NR}^{(0)} + \delta_{NR}^{(1)} + \delta_{NR}^{(2)} = \text{LO} \quad + \quad \text{NLO} \quad + \quad \text{N}^2\text{LO}$$

# NR limit at LO: $\delta_{NR}^{(0)}$

$$\delta_{NR} = \delta_{NR}^{(0)} + \delta_{NR}^{(1)} + \delta_{NR}^{(2)}$$

- $\delta_{NR}^{(0)} \propto |\mathbf{R} - \mathbf{R}'|^2$

$$\delta_{D1}^{(0)} = -\frac{2\pi m_r^3}{9} (Z\alpha)^5 \int_{\omega_{th}}^{\infty} d\omega \sqrt{\frac{2m_r}{\omega}} S_{D1}(\omega)$$

- $S_{D1}(\omega) \implies$  electric dipole response function [  $\hat{D}_1 = R Y_1(\hat{R})$  ]
- $\delta_{D1}^{(0)}$  is the dominant contribution to  $\delta_{pol}$



# NR limit at LO: $\delta_{NR}^{(0)}$

$$\delta_{NR} = \delta_{NR}^{(0)} + \delta_{NR}^{(1)} + \delta_{NR}^{(2)}$$

- $\delta_{NR}^{(0)} \propto |\mathbf{R} - \mathbf{R}'|^2$

$$\delta_{D1}^{(0)} = -\frac{2\pi m_r^3}{9} (Z\alpha)^5 \int_{\omega_{th}}^{\infty} d\omega \sqrt{\frac{2m_r}{\omega}} S_{D1}(\omega)$$

- $S_{D1}(\omega) \implies$  electric dipole response function [  $\hat{D}_1 = R Y_1(\hat{R})$  ]

- $\delta_{D1}^{(0)}$  is the dominant contribution to  $\delta_{pol}$

- $\implies$  Rel. and Coulomb corrections added at this order

# NR limit at NLO: $\delta_{NR}^{(1)}$

$$\delta_{NR} = \delta_{NR}^{(0)} + \delta_{NR}^{(1)} + \delta_{NR}^{(2)}$$

- $\delta_{NR}^{(1)} \propto |\mathbf{R} - \mathbf{R}'|^3$

$$\delta_{NR}^{(1)} = \delta_{R3pp}^{(1)} + \delta_{Z3}^{(1)}$$

$$\delta_{R3pp}^{(1)} = -\frac{m_r^4}{24} (Z\alpha)^5 \iint d\mathbf{R}d\mathbf{R}' |\mathbf{R} - \mathbf{R}'|^3 \langle N_0 | \hat{\rho}^\dagger(\mathbf{R}) \hat{\rho}(\mathbf{R}') | N_0 \rangle$$

$$\delta_{Z3}^{(1)} = \frac{m_r^4}{24} (Z\alpha)^5 \iint d\mathbf{R}d\mathbf{R}' |\mathbf{R} - \mathbf{R}'|^3 \rho_0(\mathbf{R}) \rho_0(\mathbf{R}')$$

- $\delta_{R3pp}^{(1)} \Rightarrow$  3rd-order proton charge correlation

- $\delta_{Z3}^{(1)} \Rightarrow$  3rd-order Zemach moment

# NR limit at NLO: $\delta_{NR}^{(1)}$

$$\delta_{NR} = \delta_{NR}^{(0)} + \delta_{NR}^{(1)} + \delta_{NR}^{(2)}$$

- $\delta_{NR}^{(1)} \propto |\mathbf{R} - \mathbf{R}'|^3$

$$\delta_{NR}^{(1)} = \delta_{R3pp}^{(1)} + \delta_{Z3}^{(1)}$$

$$\delta_{R3pp}^{(1)} = -\frac{m_r^4}{24} (Z\alpha)^5 \iint d\mathbf{R}d\mathbf{R}' |\mathbf{R} - \mathbf{R}'|^3 \langle N_0 | \hat{\rho}^\dagger(\mathbf{R}) \hat{\rho}(\mathbf{R}') | N_0 \rangle$$

$$\delta_{Z3}^{(1)} = \frac{m_r^4}{24} (Z\alpha)^5 \iint d\mathbf{R}d\mathbf{R}' |\mathbf{R} - \mathbf{R}'|^3 \rho_0(\mathbf{R}) \rho_0(\mathbf{R}') = \frac{m_r^4}{24} (Z\alpha)^5 \langle \mathbf{r}^3 \rangle_{(2)}$$

- $\delta_{R3pp}^{(1)} \Rightarrow$  3rd-order proton charge correlation

- $\delta_{Z3}^{(1)} \Rightarrow$  3rd-order Zemach moment  
cancels Zemach moment in finite-size corrections  
c.f. Pachucki '11 & Friar '13 ( $\mu\text{D}$ )

# NR limit at N<sup>2</sup>LO: $\delta_{NR}^{(2)}$

$$\delta_{NR} = \delta_{NR}^{(0)} + \delta_{NR}^{(1)} + \delta_{NR}^{(2)}$$

- $\delta_{NR}^{(2)} \propto |\mathbf{R} - \mathbf{R}'|^4$

$$\delta_{NR}^{(2)} = \frac{m_r^5}{18} (Z\alpha)^5 \int_{\omega_{\text{th}}}^{\infty} d\omega \sqrt{\frac{\omega}{2m_r}} \left[ S_{R^2}(\omega) + \frac{16\pi}{25} S_Q(\omega) + \frac{16\pi}{5} S_{D_1 D_3}(\omega) \right]$$

- $S_{R^2}(\omega) \implies$  monopole response function

- $S_Q(\omega) \implies$  quadrupole response function

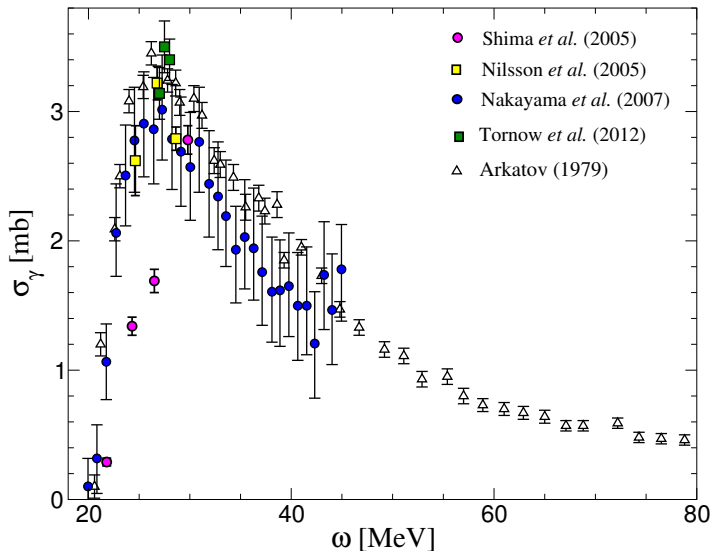
- $S_{D_1 D_3}(\omega) \implies$  interference between  $D_1$  and  $D_3$  [  $\hat{D}_3 = R^3 Y_1(\hat{R})$  ]

**BACK UP**



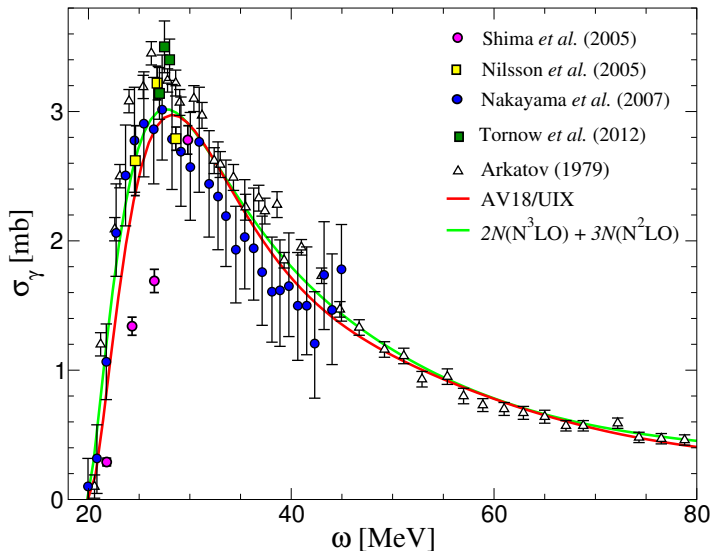
# $^4\text{He}$ photoabsorption cross sections

electric-dipole photoabsorption cross section  $\sigma_\gamma(\omega) = 4\pi^2\alpha\omega S_{D1}(\omega)$



# $^4\text{He}$ photoabsorption cross sections

electric-dipole photoabsorption cross section  $\sigma_\gamma(\omega) = 4\pi^2\alpha\omega S_{D1}(\omega)$





# Non-Relativistic Approximation

- Coulomb interactions in the (virtual) intermediate state are neglected
- Results are expressed in an expansion of  $\sqrt{2m_r\omega}|\mathbf{R} - \mathbf{R}'|$

$$\Delta E_{NR} = \Delta E_{NR}^{(2)} + \Delta E_{NR}^{(3)} + \Delta E_{NR}^{(4)}$$

# Non-Relativistic Approximation

- Coulomb interactions in the (virtual) intermediate state are neglected
- Results are expressed in an expansion of  $\sqrt{2m_r\omega}|\mathbf{R} - \mathbf{R}'|$

$$\Delta E_{NR} = \Delta E_{NR}^{(2)} + \Delta E_{NR}^{(3)} + \Delta E_{NR}^{(4)}$$

## 1. $\sim R^2$ term:

- $\Delta E_{NR}^{(2)}$  is the dominant polarizability contribution

$$\Delta E_{NR}^{(2)} = -\frac{2\pi m_r^3}{9} (Z\alpha)^5 \int_{\omega_{th}}^{\infty} d\omega \sqrt{\frac{2m_r}{\omega}} S^{D_1}(\omega)$$

- $S^{D_1}(\omega) = \frac{1}{2J_0+1} \sum_{N \neq N_0, J} |\langle N_0 J_0 | \hat{D}_1 | N J \rangle|^2 \delta(\omega - E_N + E_{N_0})$
- $\hat{D}_1 = \frac{1}{Z} \sum_i R_i Y_1(\hat{R}_i)$

# Non-Relativistic Approximation

- Coulomb interactions in the (virtual) intermediate state are neglected
- Results are expressed in an expansion of  $\sqrt{2m_r\omega}|\mathbf{R} - \mathbf{R}'|$

$$\Delta E_{NR} = \Delta E_{NR}^{(2)} + \Delta E_{NR}^{(3)} + \Delta E_{NR}^{(4)}$$

2.  $\sim R^3$  term:

$$\Delta E_{NR}^{(3)} = -\frac{m_r^4}{24}(Z\alpha)^5 \left[ \iint d^3R d^3R' |\mathbf{R} - \mathbf{R}'|^3 \langle N_0 | \hat{\rho}^\dagger(\mathbf{R}) \hat{\rho}(\mathbf{R}') | N_0 \rangle - \langle r^3 \rangle_{(2)} \right]$$

# Non-Relativistic Approximation

- Coulomb interactions in the (virtual) intermediate state are neglected
- Results are expressed in an expansion of  $\sqrt{2m_r\omega}|\mathbf{R} - \mathbf{R}'|$

$$\Delta E_{NR} = \Delta E_{NR}^{(2)} + \Delta E_{NR}^{(3)} + \Delta E_{NR}^{(4)}$$

2.  $\sim R^3$  term:

$$\Delta E_{NR}^{(3)} = -\frac{m_r^4}{24}(Z\alpha)^5 \left[ \iint d^3R d^3R' |\mathbf{R} - \mathbf{R}'|^3 \langle N_0 | \hat{\rho}^\dagger(\mathbf{R}) \hat{\rho}(\mathbf{R}') | N_0 \rangle - \langle r^3 \rangle_{(2)} \right]$$

- **1st term:** charge correlation function vanishes in point-nucleon limit

# Non-Relativistic Approximation

- Coulomb interactions in the (virtual) intermediate state are neglected
- Results are expressed in an expansion of  $\sqrt{2m_r\omega}|\mathbf{R} - \mathbf{R}'|$

$$\Delta E_{NR} = \Delta E_{NR}^{(2)} + \Delta E_{NR}^{(3)} + \Delta E_{NR}^{(4)}$$

2.  $\sim R^3$  term:

$$\Delta E_{NR}^{(3)} = -\frac{m_r^4}{24}(Z\alpha)^5 \left[ \iint d^3R d^3R' |\mathbf{R} - \mathbf{R}'|^3 \langle N_0 | \hat{\rho}^\dagger(\mathbf{R}) \hat{\rho}(\mathbf{R}') | N_0 \rangle - \langle r^3 \rangle_{(2)} \right]$$

- **1st term:** charge correlation function vanishes in point-nucleon limit

- **2nd term:** Zemach moment

$$\langle r^3 \rangle_{(2)} = \iint d^3R d^3R' |\mathbf{R} - \mathbf{R}'|^3 \rho_0(\mathbf{R}) \rho_0(\mathbf{R}') \\ \rho_0(\mathbf{R}) = \langle N_0 | \hat{\rho}(\mathbf{R}) | N_0 \rangle$$

cancels exactly the **Zemach term** in (elastic) finite-size corrections  
c.f. Pachucki PRL 2011 ( $\mu\text{D}$ )

# Non-Relativistic Approximation

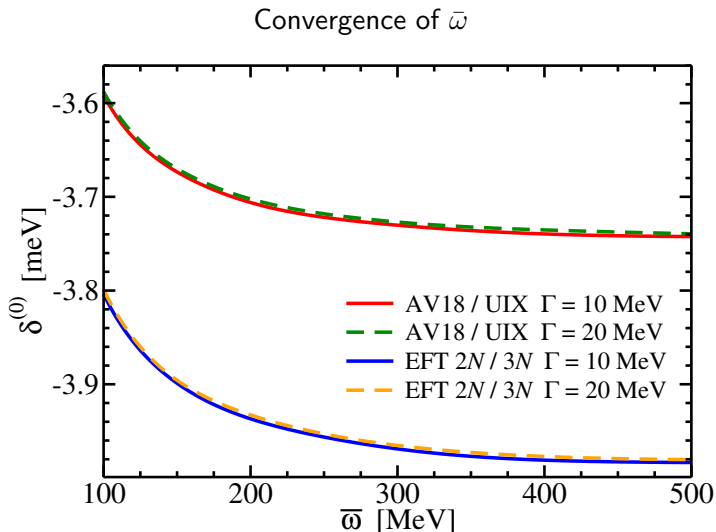
## 3. $\sim R^4$ term:

- $\Delta E_{NR}^{(4)}$  corresponds to higher-multipole corrections

$$\Delta E_{NR}^{(4)} = \frac{m_r^5}{18} (Z\alpha)^5 \int_{\omega_{th}}^{\infty} d\omega \sqrt{\frac{\omega}{2m_r}} \left[ S^{Q_0}(\omega) + \frac{16\pi}{25} S^{Q_2}(\omega) + \frac{16\pi}{5} S^{D_{13}}(\omega) \right]$$

- $$S^{R^2}(\omega) = \frac{1}{2J_0+1} \sum_{N \neq N_0, J} |\langle N_0 J_0 || \hat{R}^2 || N J \rangle|^2 \delta(\omega - E_N + E_{N_0})$$
$$S^{Q_2}(\omega) = \frac{1}{2J_0+1} \sum_{N \neq N_0, J} |\langle N_0 J_0 || \hat{Q}_2 || N J \rangle|^2 \delta(\omega - E_N + E_{N_0})$$
$$S^{D_{13}}(\omega) = \frac{1}{2J_0+1} \sum_{N \neq N_0, J} (-1)^{J_0-J} \times \text{Re} \left( \langle N_0 J_0 || \hat{D}_3 || N J \rangle \langle N J || \hat{D}_1 || N_0 J_0 \rangle \right) \delta(\omega - E_N + E_{N_0})$$
- $$\hat{R}^2 = \frac{1}{Z} \sum_i R_i^2 \qquad \hat{D}_1 = \frac{1}{Z} \sum_i R_i Y_1(\hat{R}_i)$$
$$\hat{Q}_2 = \frac{1}{Z} \sum_i R_i^2 Y_2(\hat{R}_i) \qquad \hat{D}_3 = \frac{1}{Z} \sum_i R_i^3 Y_3(\hat{R}_i)$$

# Convergence of Ab-initio calculations



# Convergence of Ab-initio calculations

$\delta^{(0)}$  convergence with the largest model space  $K_{max}$

