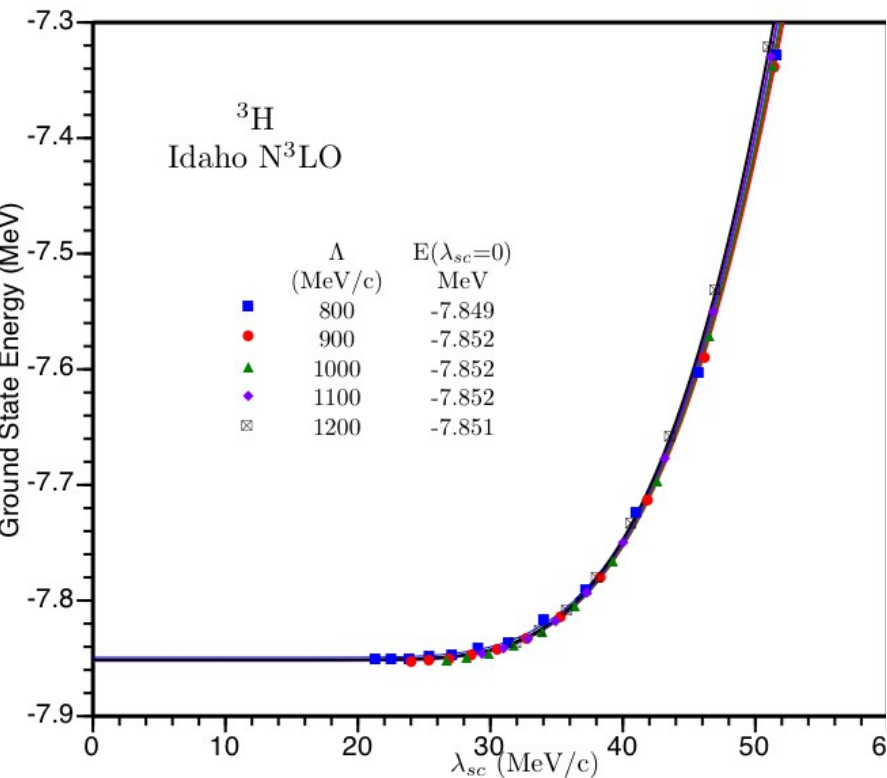


Universal properties of infrared extrapolations in a harmonic oscillator basis

Sidney A Coon and Michael K.G. Kruse



S. A. Coon, M. I. Avetian, M. K. G. Kruse, U. Van Kolck, P. Maris, J. P. Vary
Archive:1205.3230, PRC 86, 054002 (2012)

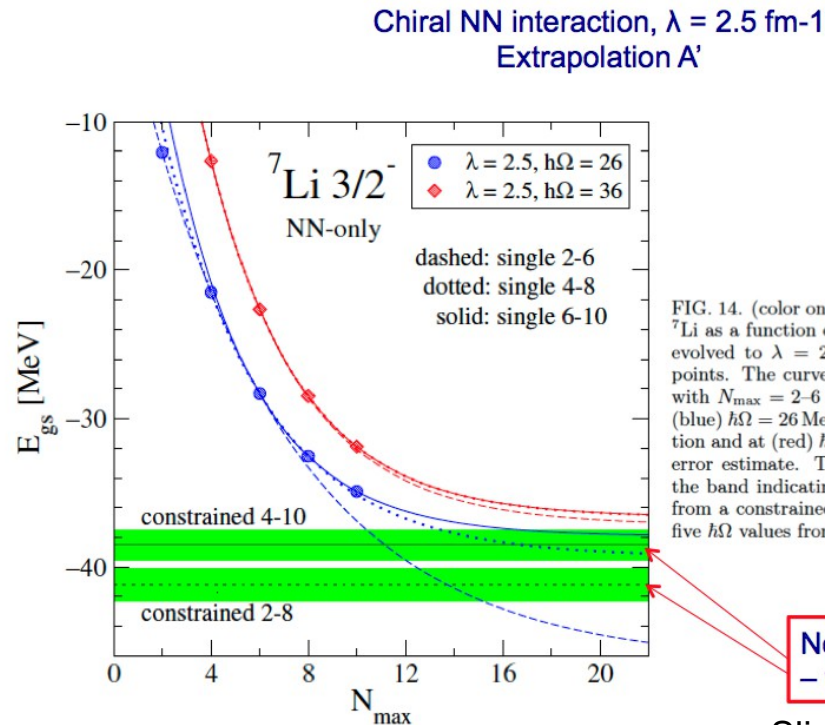


FIG. 14. (color online) Ground-state energy extrapolations of ${}^7\text{Li}$ as a function of N_{max} with an $N^3\text{LO}$ NN interaction [21] evolved to $\lambda = 2.5 \text{ fm}^{-1}$. The symbols are the calculated points. The curves show single extrapolations using Eq. (3) with $N_{\text{max}} = 2-6$ (dashed), 4-8 (dotted) and 6-10 (solid) at (blue) $\hbar\Omega = 26 \text{ MeV}$ which minimize the amount of extrapolation and at (red) $\hbar\Omega = 36 \text{ MeV}$ which minimize the numerical error estimate. The horizontal dotted and solid lines, with the band indicating the associated error bars, are the result from a constrained fit following the procedure of Ref. [1] for five $\hbar\Omega$ values from 22 to 30 MeV.

Note the inconsistency
– what's happening?

Slide by James Vary

E.D. Jurgenson, P. Maris, R.J. Furnstahl, P. Navratil, W.E. Ormand, J.P. Vary,
submitted to PRC; arXiv: 1302.5473



The No-Core Shell Model (NCSM)



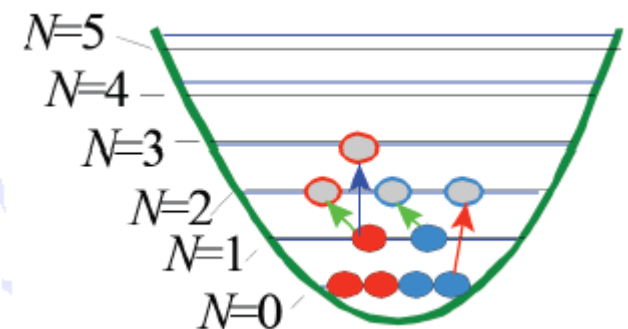
Starting Hamiltonian is translationally invariant.

$$H_A = \frac{1}{A} \sum_{i < j}^A \frac{(\vec{p}_i - \vec{p}_j)^2}{2m} + \sum_{i < j}^A V_{NN,ij}$$

Provided interaction is “soft” we don't need to do any renormalization of interaction,

It's that “simple”.

NCSM has two parameters:
Nmax and Ω

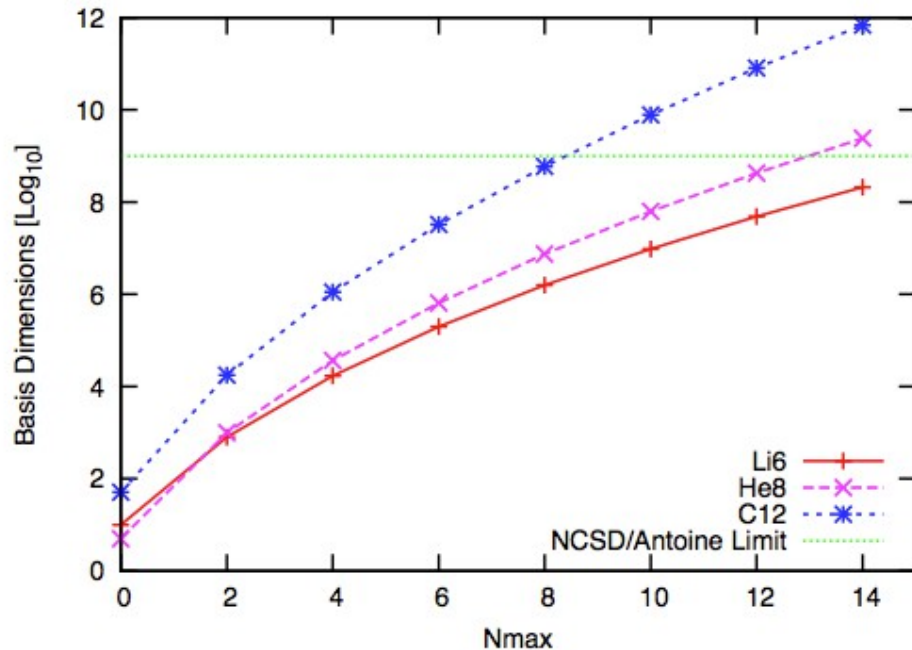


If we now use a single-particle basis, we have to remove the spurious CM states. Translational invariance is automatic if HO basis depends on Jacobi coordinates.

Advantage in m-scheme: Antisymmetry is easy to implement.

Disadvantage in m-scheme: Number of basis states is much larger than JT basis

M-scheme basis dimensions



- Size of the m-scheme basis grows rapidly with increasing Nmax.
- Switch to HO JT coupled basis? Possibly, but painful.
- Difficulties with such an approach, e.g. Jacobi coordinates or rewrite codes.
- Even if techniques like SRG potentials are used, you still can't perform converged calculations all the time.
- But why stick with the HO basis?
- Only basis where center of mass and intrinsic states can be completely decoupled.

The Variational Approach

- One can view a shell model calculation as a variational calculation, and is thus expanding the configuration space merely serves to improve the trial wavefunction.
- The traditional shell-model calculation involves trial wavefunctions which are linear combinations of Slater determinants.

Irvine, J. M. et al. "Nuclear Shell-Model Calculations and Strong Two-Body Correlations"

2.1.2. Linear Trial Functions

We next consider a trial function in the form of a linear expansion:

$$\psi_T = \sum_{i=1}^N a_i \varphi_i \quad (2.10)$$

Variational energy as a function of oscillator energy $\hbar\omega$ for fixed number of quanta
 Number of quanta increases by two for each curve

1969 H atom up to 10 quanta

M. MOSHINSKY

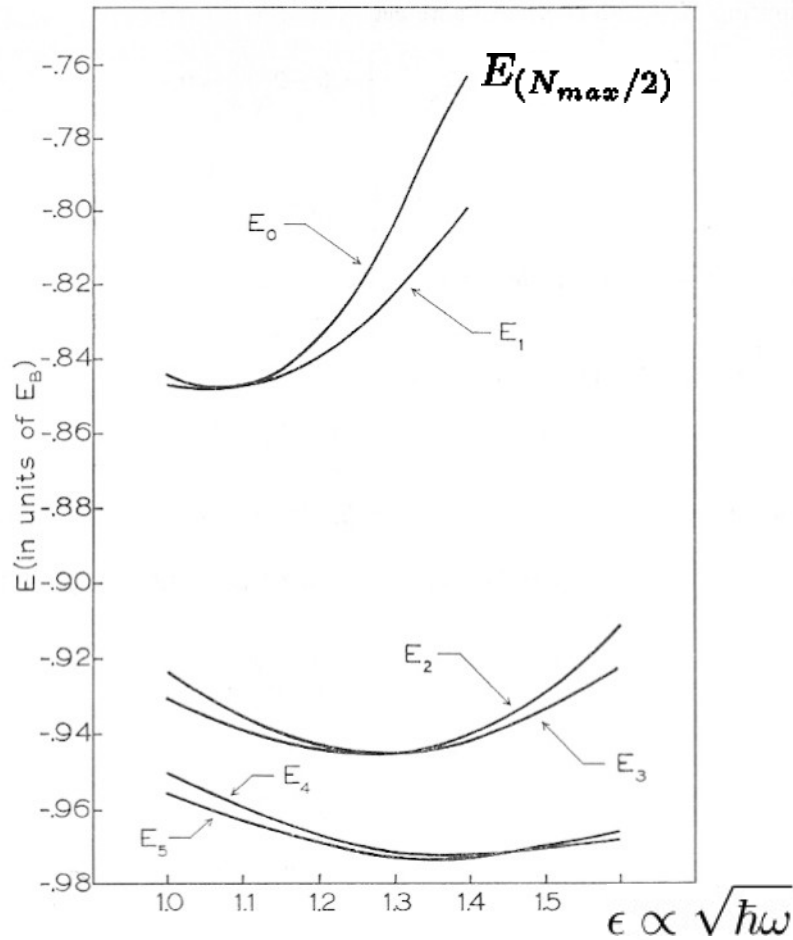
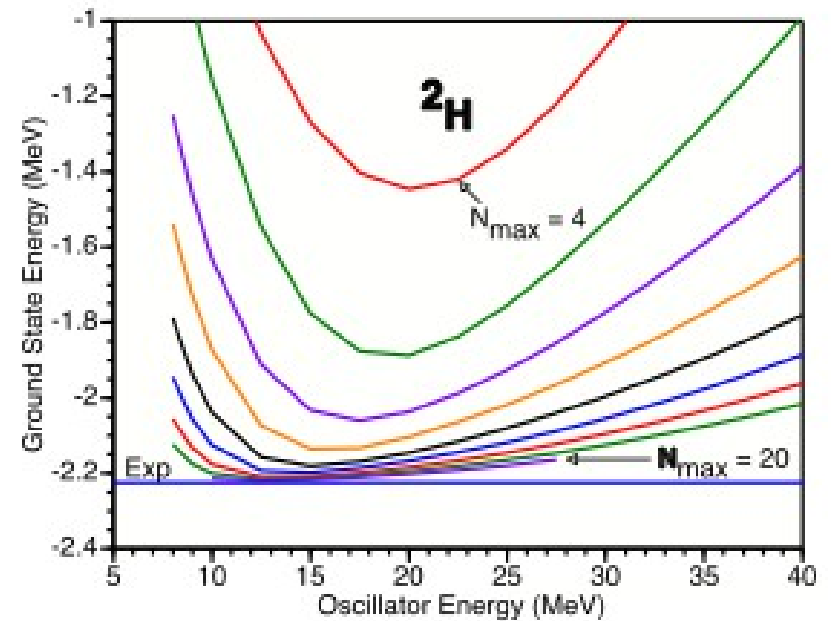
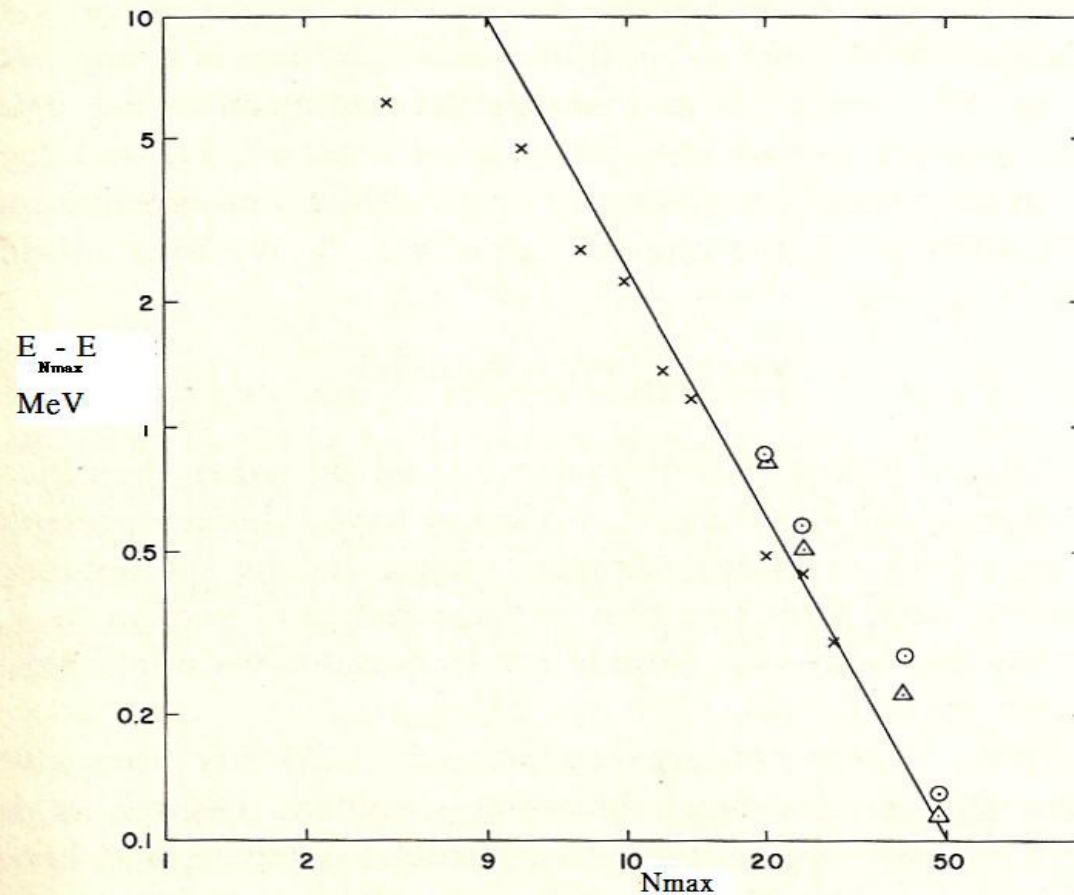


FIG. 1. Energy of the ground state of the H atom as a function of the parameter ϵ for the variational analysis discussed in Section 3. This energy $E_p(\epsilon)$, $p = 0, 1, 2, 3, 4, 5$ is associated with a trial wave function $\psi_p = \sum_{n=0}^p a_n^{(p)} |n00\rangle$, where $|n00\rangle$ is a harmonic-oscillator state of frequency $\hbar\omega = (me^4/2\hbar^2)\epsilon^2$.

2009 deuteron up to 20 quanta



No-core full configuration method of
 Maris, Vary, Shirokov



L.M Delves; in Advances
In Nuclear Physics vol 5 1972

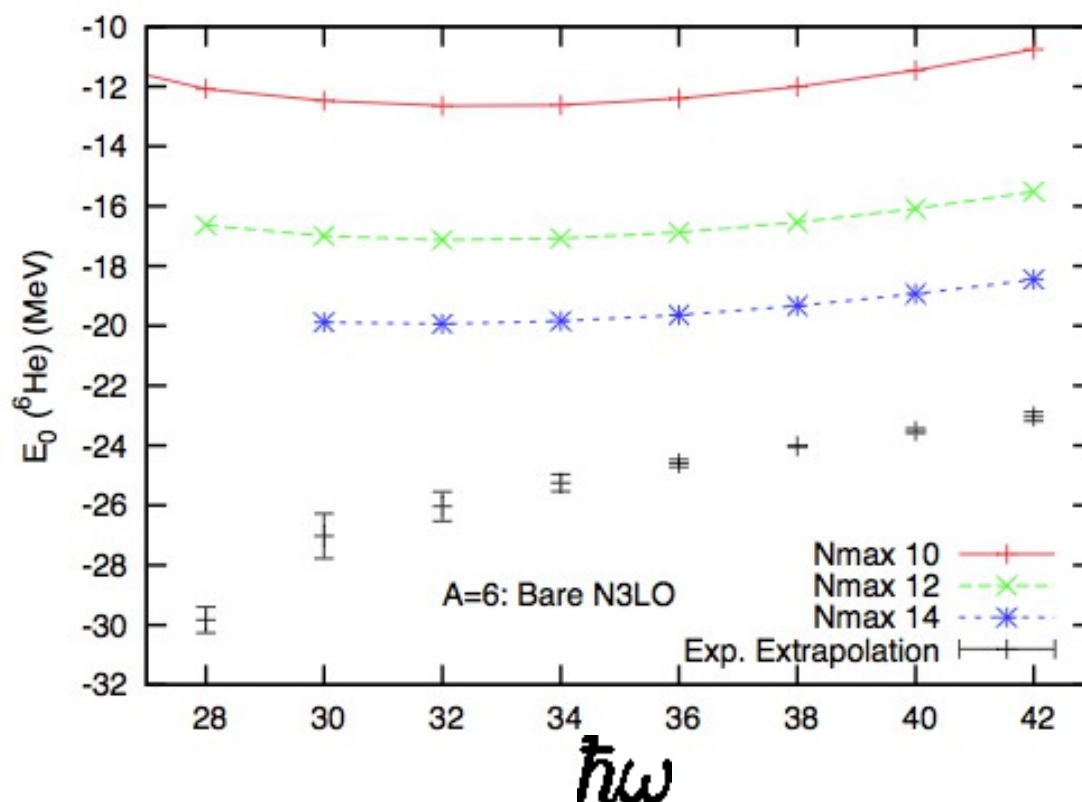
$$E_{N_{max}} = E + P(N_{max})^{-2}$$

“nonsmooth potentials” like Yukawa

Fig. 8. Convergence rates for variational calculations with a harmonic oscillator basis. \odot Deuteron, Yamaguchi potential; \times triton, Yamaguchi potential; and \triangle deuteron, Reid potential. Results taken from (JLS 70). The solid line has a slope of -2.0 .

“These results are independent of the dimensionality of the problem, that is, of the number of particles, provided that the appropriate N_{max} is used. ... The extrapolated results of these authors have been used for E . On the logarithmic scale used, these differences are predicted by our crude theory to lie on a straight line of slope 2 for the Reid potential; it is not clear to what extent we should expect the nonlocal [separable] Yamaguchi potential to be ‘smooth’.”

The current method is to seek the minimum of $E(N_{\text{max}}, \hbar\omega)$.



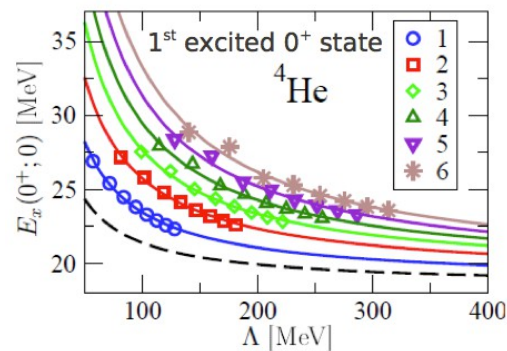
Extrapolation A of Maris et al PRC 79, 014308 (2009)

Current Method is Unsatisfactory...

- ...from an effective field theory point of view.
- Results are oscillator frequency dependent.
- No clear control of ultra-violet or infra-red nuclear physics.
- The goal is to investigate an alternate way from a more formal view point.

UV & IR CUTOFFS INTRODUCED

Construction of an effective field theory
within the No Core Shell Model



-> calculation at **Leading order** :

two N-N contact interactions in the 3S_1 , 1S_0 channel and a three-body contact interaction in the 3-nucleon $S_{1/2}$ channel

-> coupling constants fitted to the binding energy of the deuteron, triton and ^4He .



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Physics Letters B 653 (2007) 358–362

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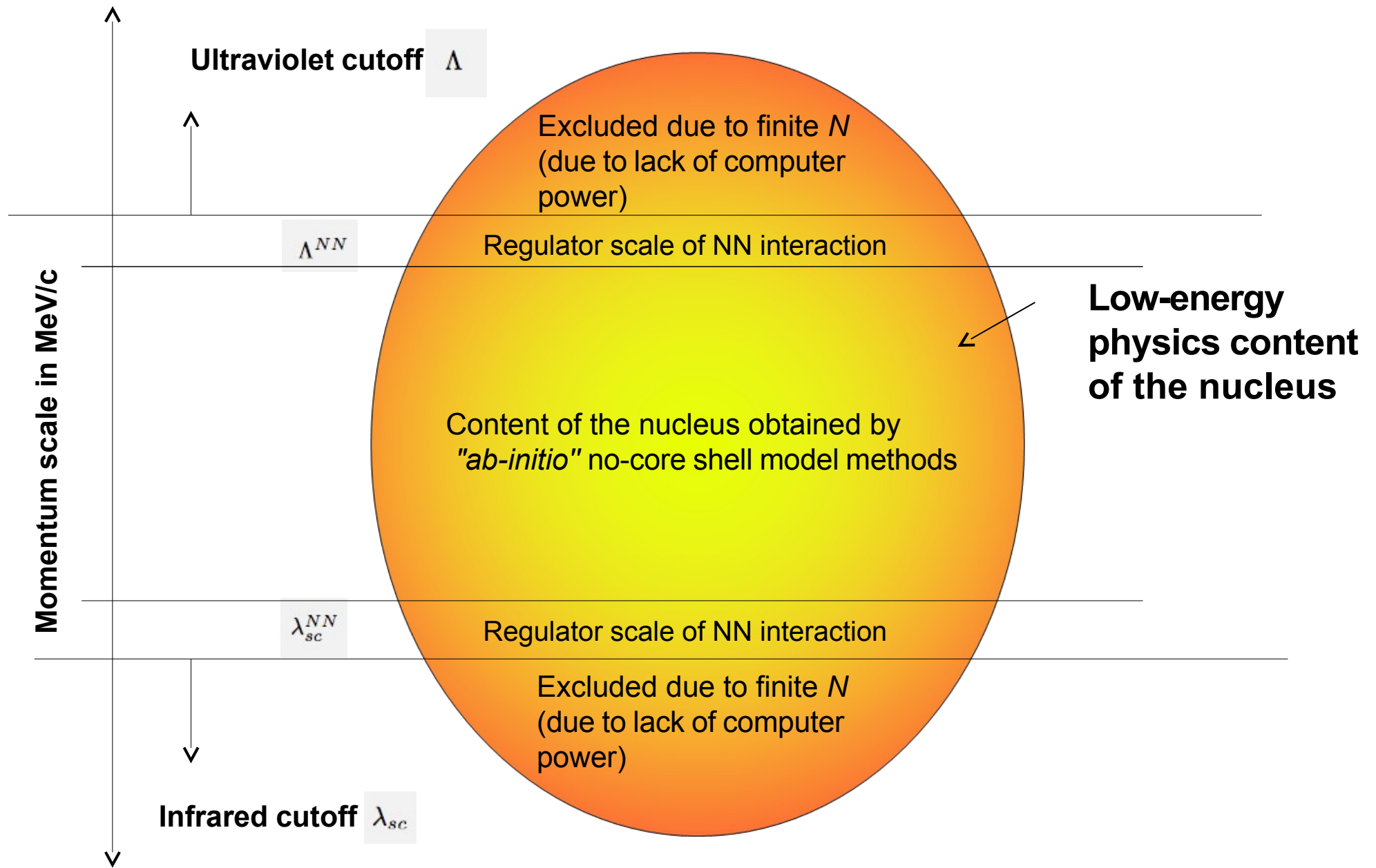
No-core shell model in an effective-field-theory framework

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Received 5 March 2007; received in revised form 8 July 2007; accepted 31 July 2007



Definition of Momentum Cutoffs

Ultraviolet momentum cutoff

$$\Lambda = \sqrt{m_N(N + 3/2)\hbar\omega}$$

Stetcu et al. PLB **653** (2007)

“Kallio momentum” $\Lambda_a = \sqrt{2}\Lambda$

JPVary lectures (2012)

Furnstahl et al. PRC **86** (2012)

Infrared momentum cutoff
defined by discrete energy
levels of the HO

$$\lambda = \sqrt{(m_N \hbar \omega)}$$

Stetcu et al. PLB **653** (2007)

Alternate infrared cutoff
defined by maximal radial
extent of highest sp state
of model space

$$\lambda_{SC} = \sqrt{(m_N \hbar \Omega)/(N + 3/2)}$$

Jurgenson et al. PRC (2011)

$$\lambda_{SC} = \lambda^2 / \Lambda$$

Alternate λ_{sc} in Furnstahl et al. PRC **86** (2012),

More et al. PRC **87** (2013),

Furnstahl et al. ArXiv:1312.6876

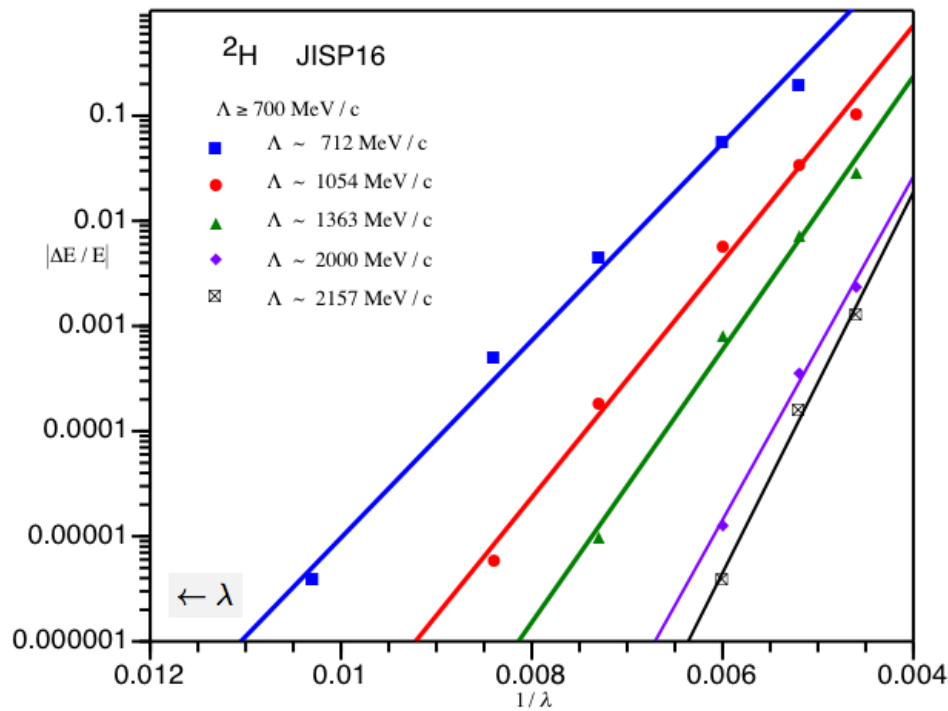
Strategy:

minimize $E(\Lambda, \lambda)$ or $E(\Lambda, \lambda_{sc})$

Works only if the answer improves
as we lower the infrared cutoff
at a fixed ultraviolet cutoff

Or if answer improves as we raise
UV cutoff at a fixed ir cutoff

Remove IR effects by decreasing value of IR momentum cutoff in the function chosen as an extrapolator whilst keeping the UV cutoff undisturbed.



$B/\sqrt{\Lambda}$ (i.e., multiplier of $1/\sqrt{\lambda_{sc}}$) is constant to within 5 %.

The momentum cutoff λ will remove IR effects. Indeed, any momentum cutoff $\lambda_{sc} \leq \lambda_{IR} \leq \Lambda$ will remove IR effects, but the IR regulator which is independent of the UV cutoff is some function of λ_{sc} . It is λ_{sc} which causes the IR effects and one does not need to decrease a IR cutoff below that of λ_{sc} to remove IR effects (i.e. extrapolate to zero).

- Extrapolator is clearly the exponential function.

- $$\frac{E(\lambda) - E(\lambda=0)}{E(\lambda=0)} \equiv \frac{\Delta E}{E} = A \exp(-B/\lambda)$$

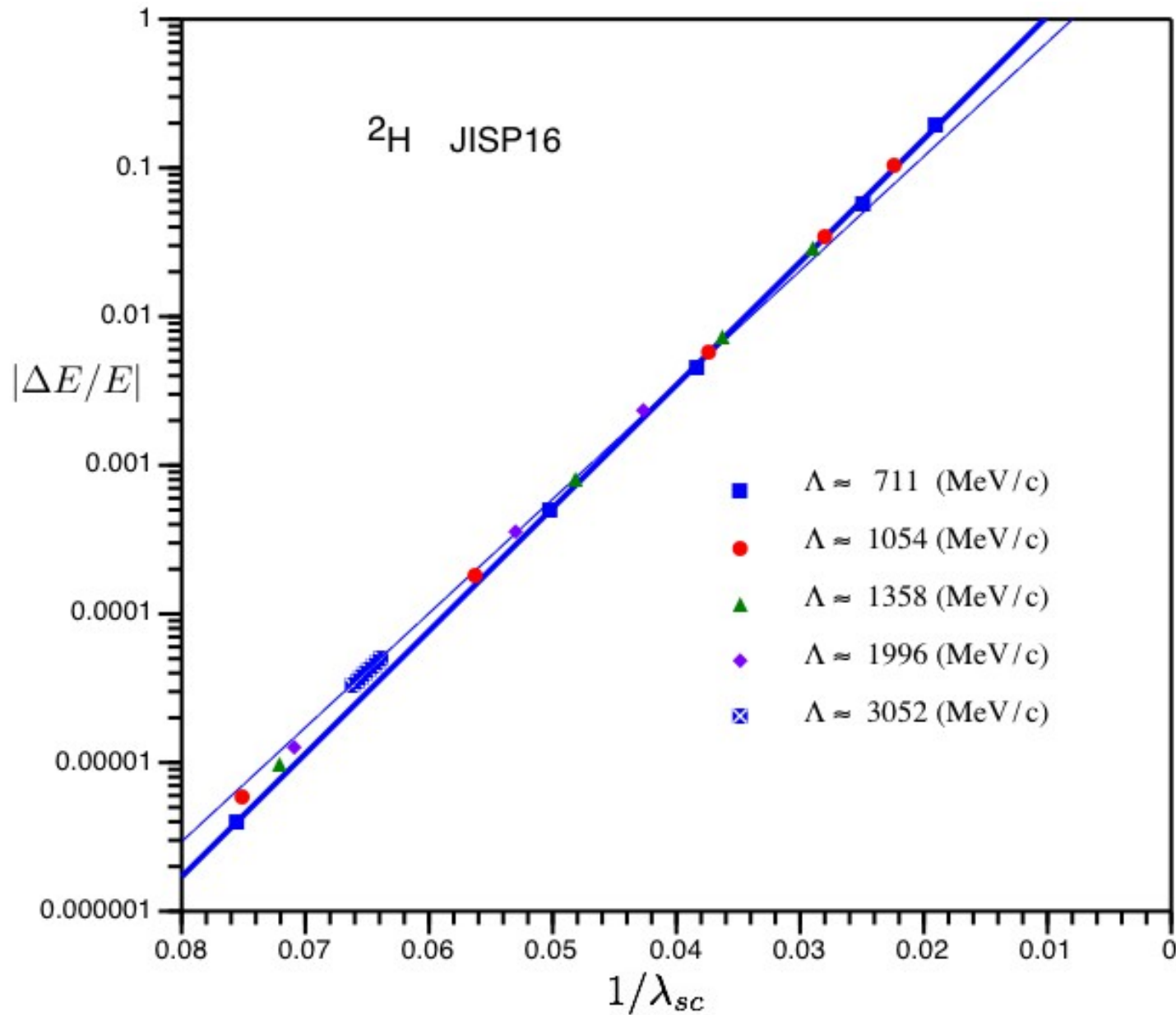
- B is a function of the UV cutoff Λ
- The IR cutoff *cannot* be aware of the UV cutoff.
- Remove dependence upon Λ

- $$\lambda = \sqrt{\Lambda \lambda_{sc}} \implies$$

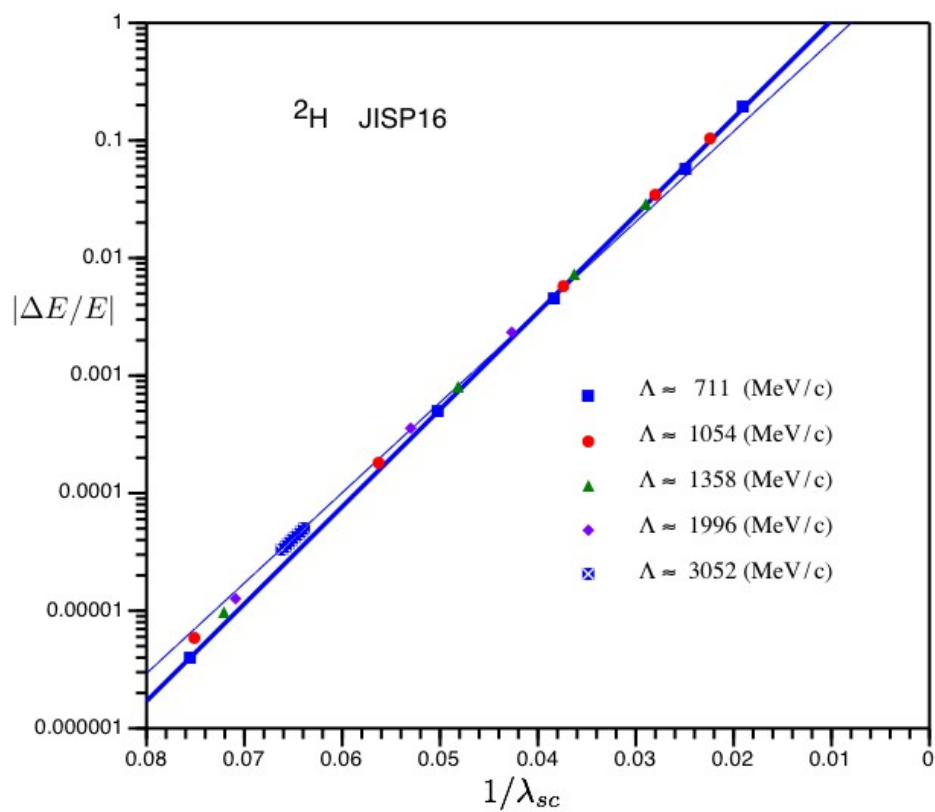
- $$\exp(-B/\lambda) = \exp(-B/\sqrt{\Lambda \lambda_{sc}}) = \exp\left(-\frac{B/\sqrt{\Lambda}}{\sqrt{\lambda_{sc}}}\right)$$

-

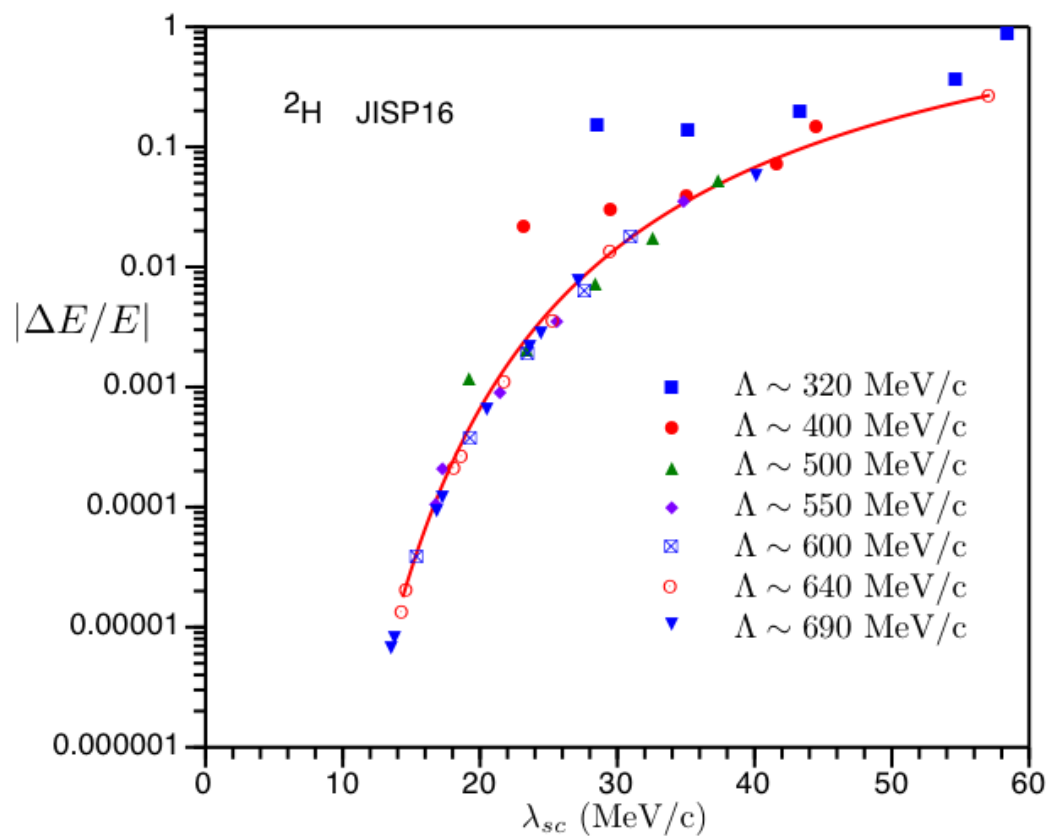
Result scales with $1/\lambda_{sc} = \Lambda/\lambda^2$ almost a universal behavior



$1/\lambda_{sc}$ has units of a length. $1/\lambda_{sc}$ is the maximal radial extent needed to encompass the system. One could call this radius "L" if one wanted a different name for $1/\lambda_{sc}$.

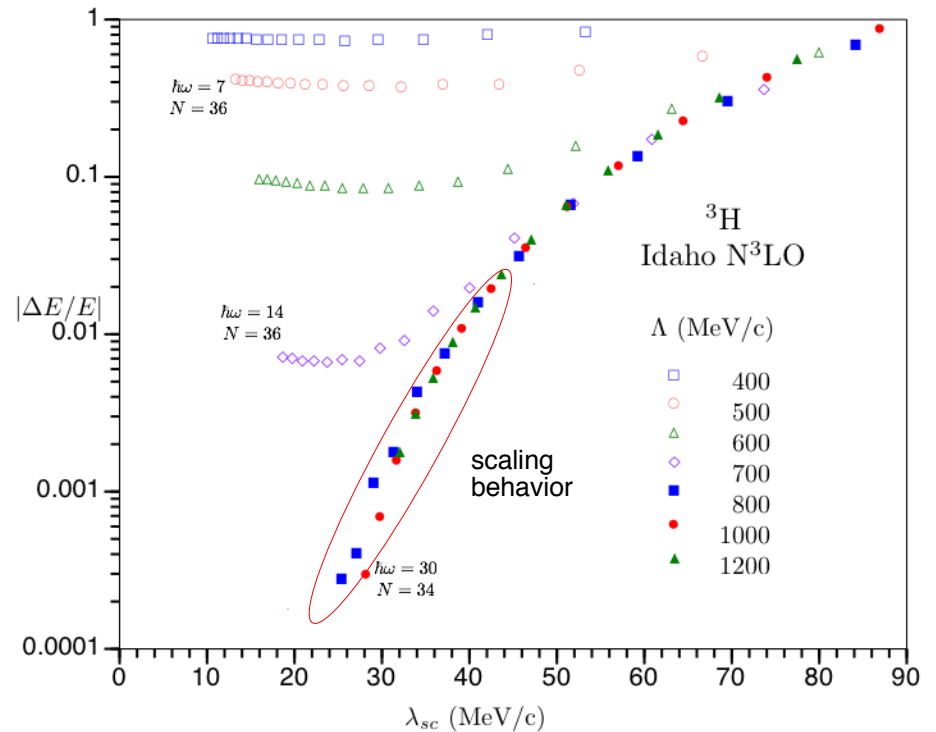
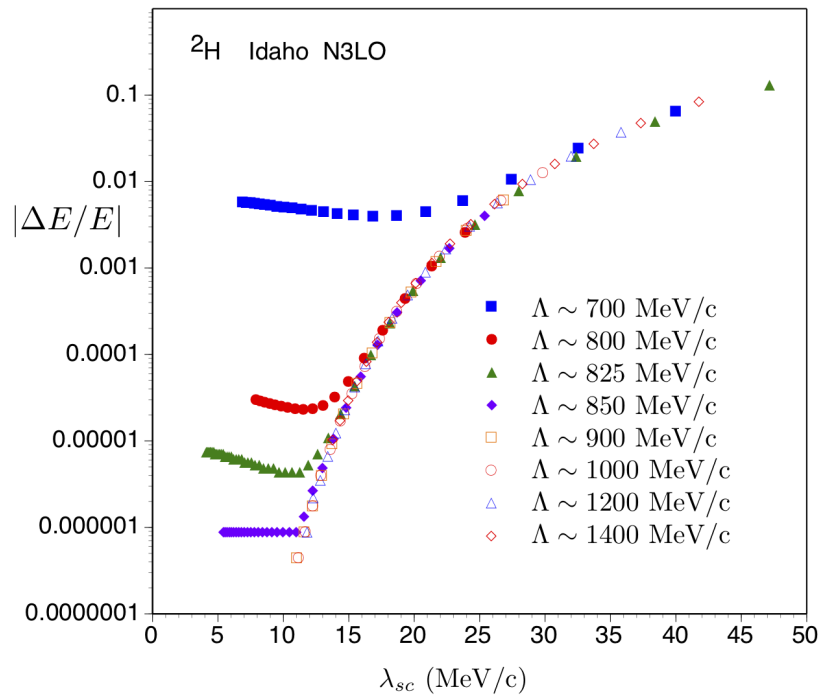


$\Lambda > 700$ MeV/c



$\Lambda < 700$ MeV/c

Running of $|\Delta E/E|$ with IR cutoff suggests an intrinsic UV scale of the NN interaction

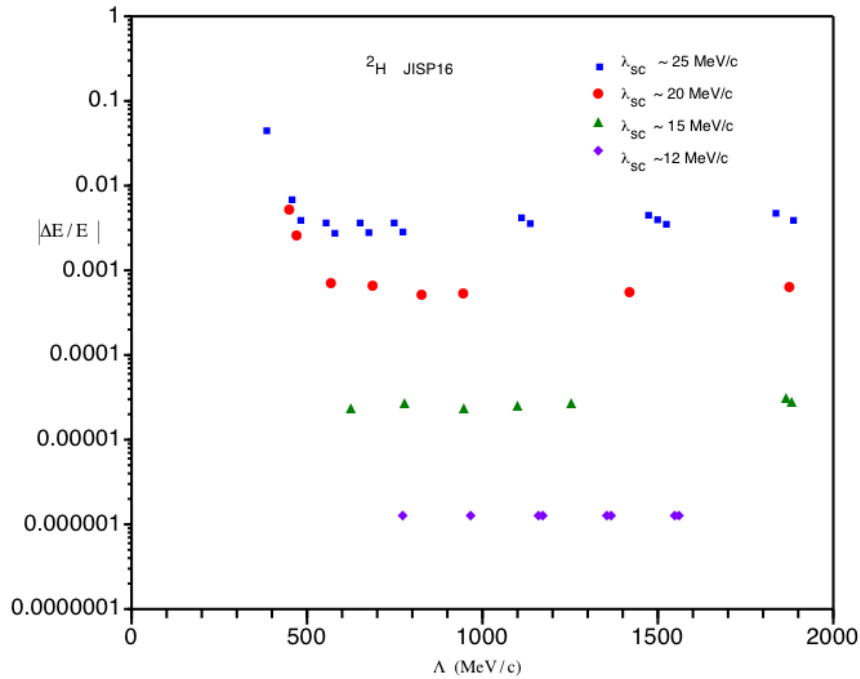


$|\Delta E/E|$ does not go to zero unless $\Lambda > \Lambda^{\text{NN}}$ where Λ^{NN} is some uv regulator scale of the NN interaction

For $\Lambda < \Lambda^{\text{NN}}$ there are missing contributions of size $|\Lambda - \Lambda^{\text{NN}}|/\Lambda^{\text{NN}}$ so “plateaus” appear as ir cutoff approaches 0.

Rise of plateaus suggests corrections are needed to Λ and λ , which are defined only to leading order in λ/Λ .

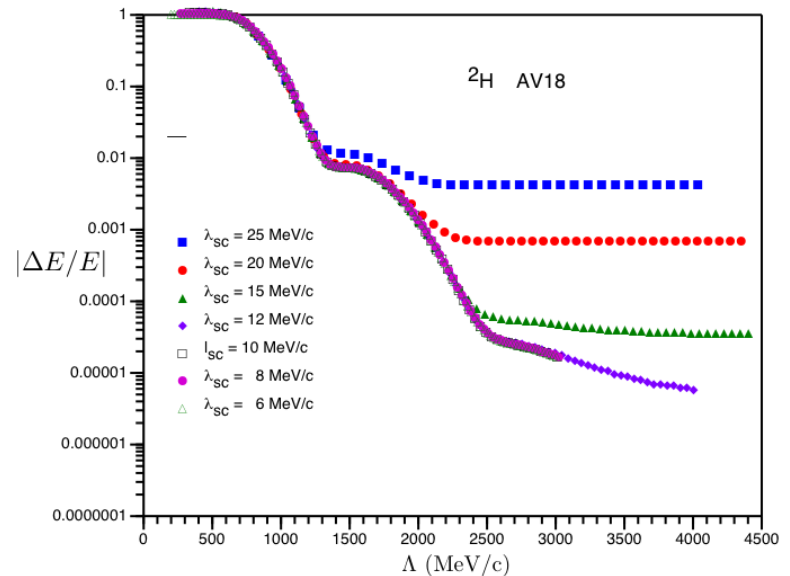
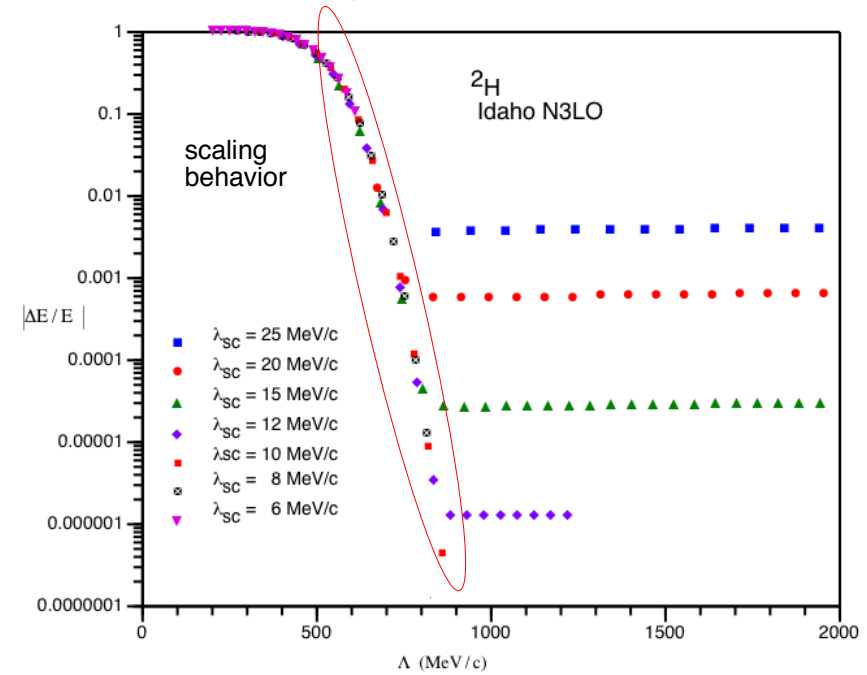
Running of $|\Delta E/E|$ with UV cutoff suggests an intrinsic IR scale of the NN interaction

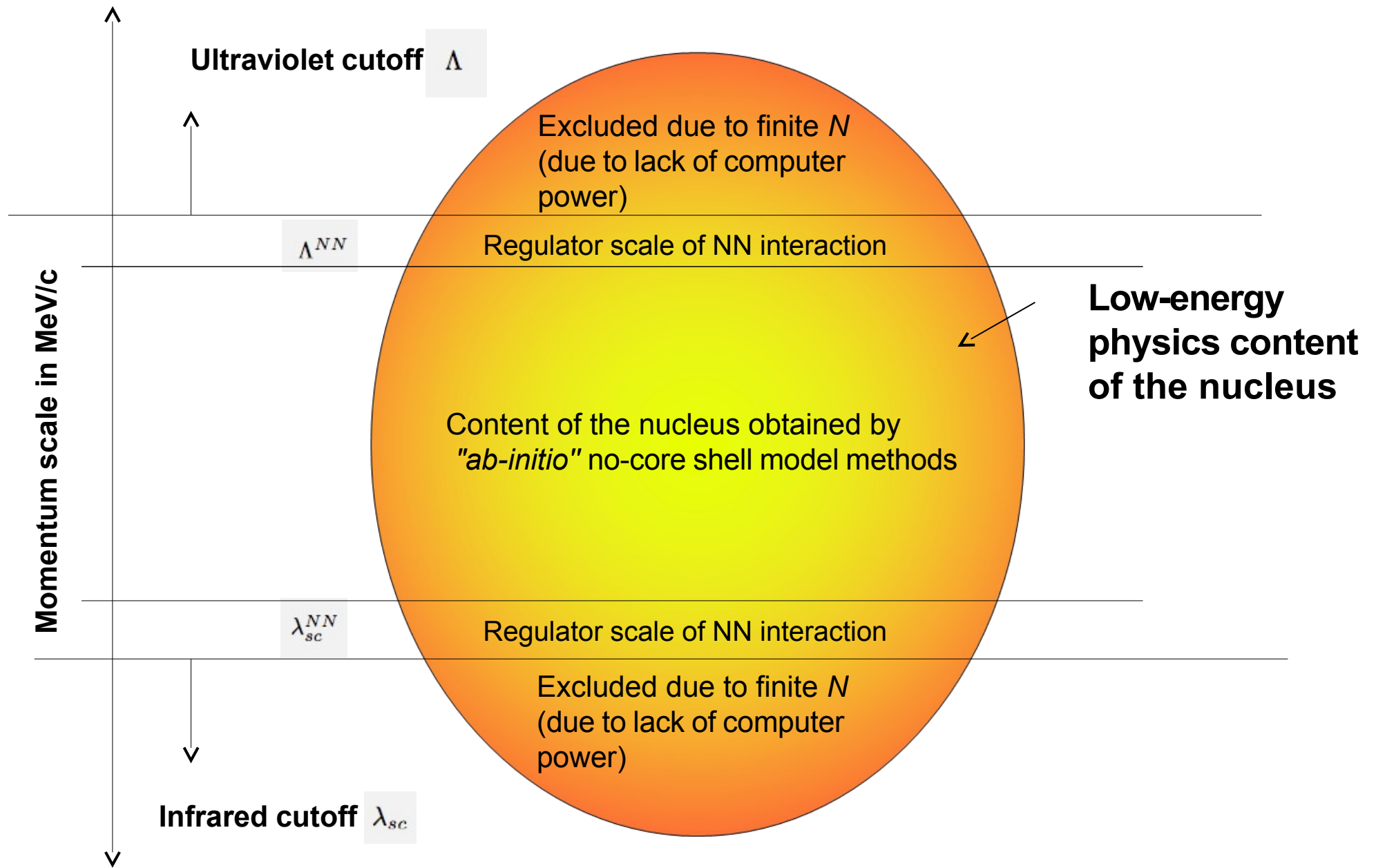


Plateau's ascribed to another “missing contributions” argument.

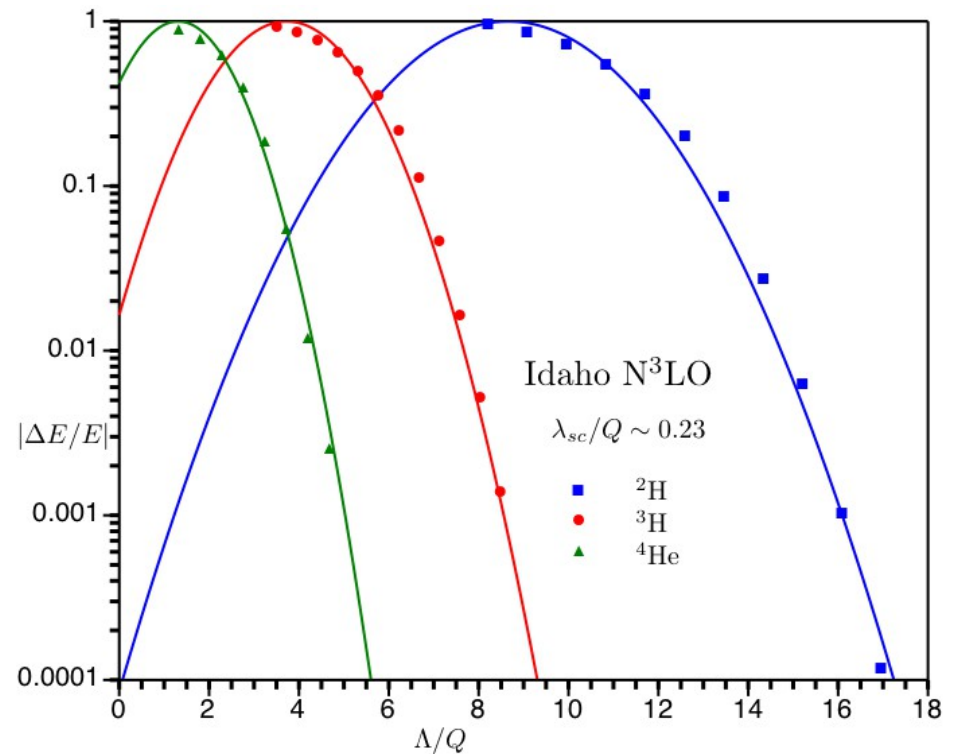
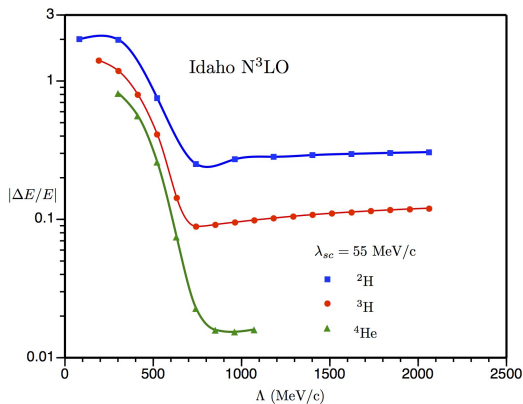
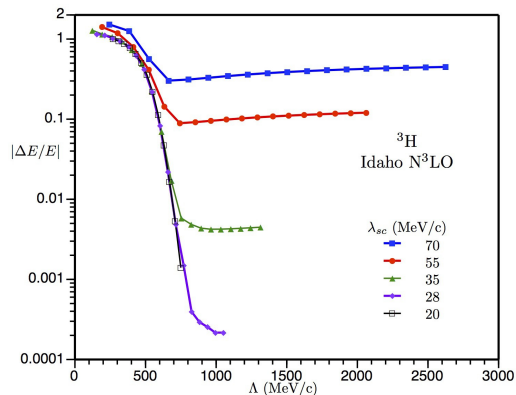
$|\Delta E/E|$ does not go to zero unless $\lambda_{\text{sc}} \leq \lambda_{\text{sc}}^{\text{NN}}$ where $\lambda_{\text{sc}}^{\text{NN}}$ is some ir regulator scale of the NN interaction.

Value of $\lambda_{\text{sc}}^{\text{NN}}$ is consistent with lowest energy configuration described by NN interaction;
 e.g. deuteron binding momentum $Q = 45$ MeV/c,
 or average of inverse scattering lengths ≈ 16 MeV/c





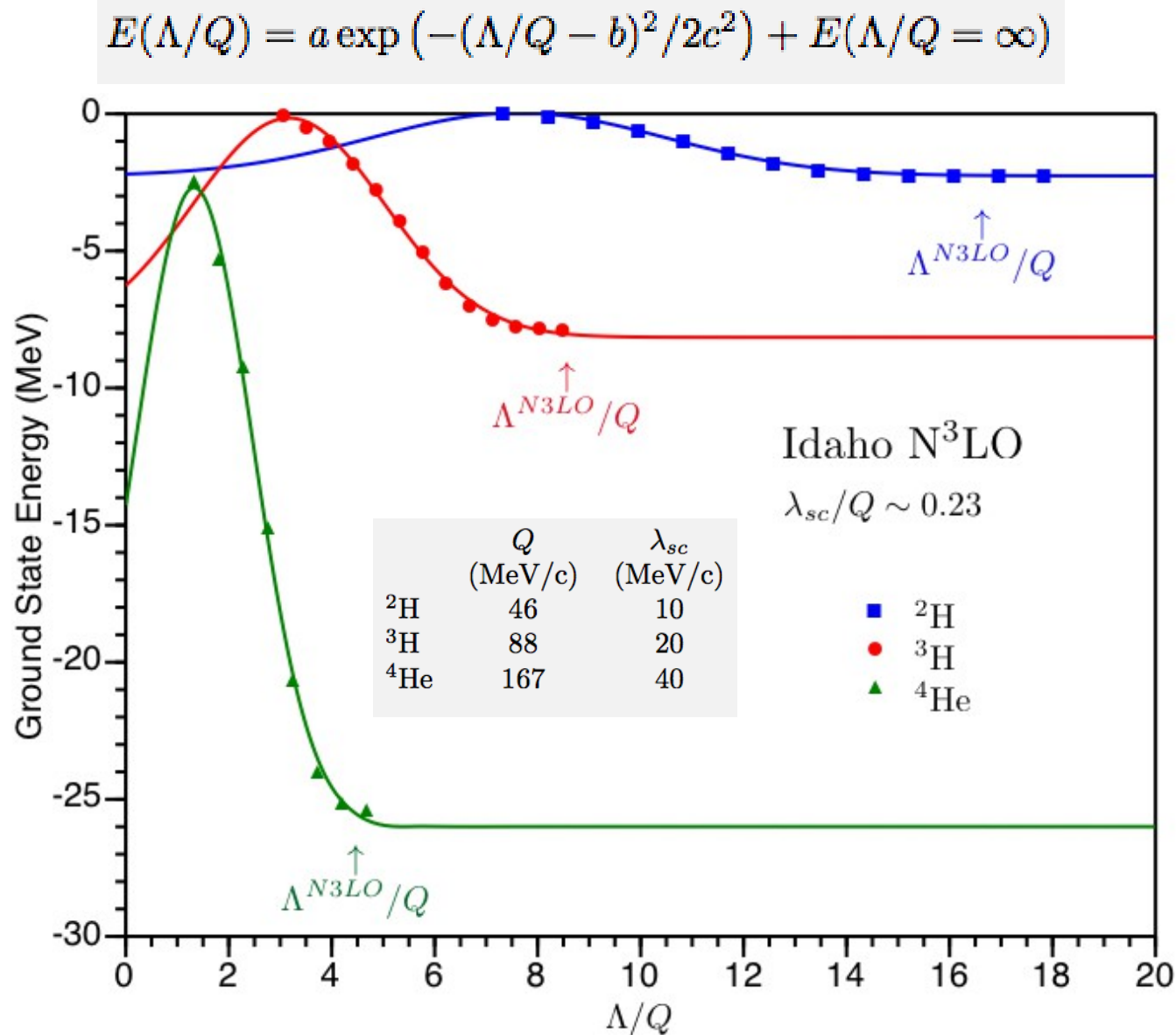
Single curve for $\lambda_{sc} \leq \lambda_{sc}^{NN}$ and $\Lambda \leq \Lambda^{NN}$



Scale each model space cutoff by binding Momentum of nucleus Q to demonstrate universal (i.e., independent of particle number) nature of the Gaussian fits to low values of UV cutoff Λ

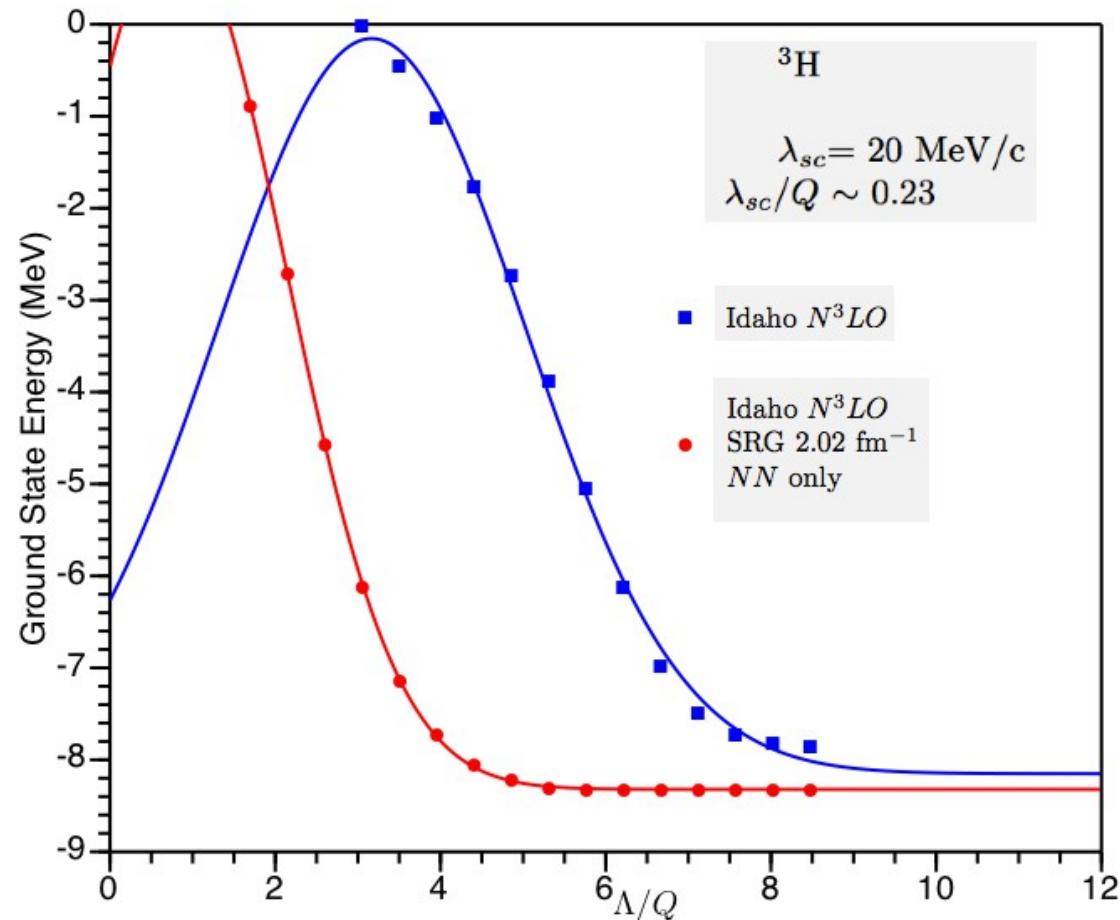
	Q (MeV/c)	λ_{sc} (MeV/c)
^2H	46	10
^3H	88	20
^4He	167	40

UV extrapolations with $\Lambda < \Lambda^{\text{NN}}$



Extrapolated energies do NOT agree with independent calculations but are *lower*:
 2 keV for deuteron, 300 keV (or 4%) for triton and 620 keV (or 2.4%) for alpha

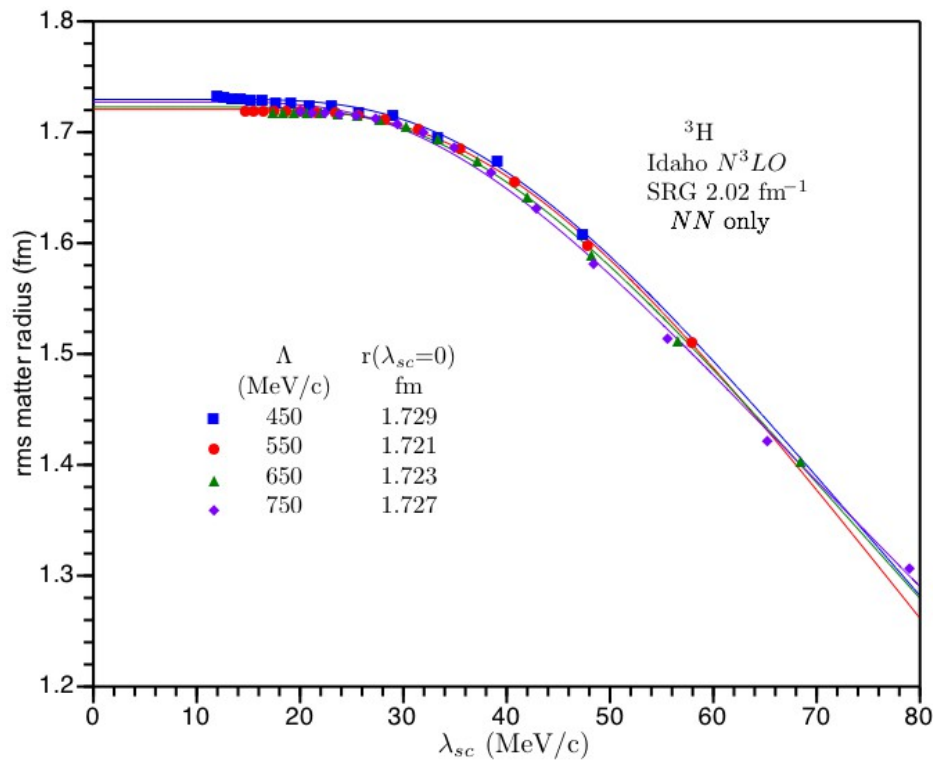
UV extrapolations with $\Lambda < \Lambda^{\text{NN}}$



- 1) Extrapolation agrees with independent calculations only for SRG transformed potential.
- 2) Extrapolation with other values of fixed λ_{sc} is neither reliable nor robust.

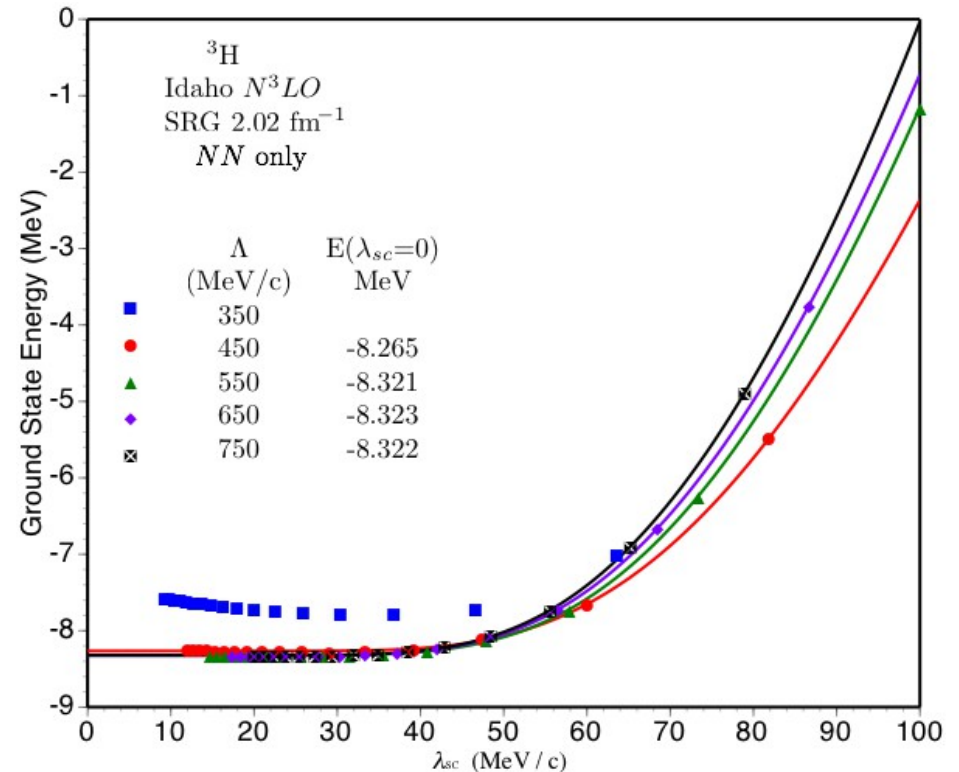
IR extrapolations with λ_{sc}

$\Lambda > \Lambda^{NN} \sim 500 \text{ MeV/c}$ for this SRG transformed interaction



Radius extrapolation

$$r(\lambda_{sc}) = -A \exp(-B/\lambda_{sc}) + r(\lambda_{sc} = 0)$$

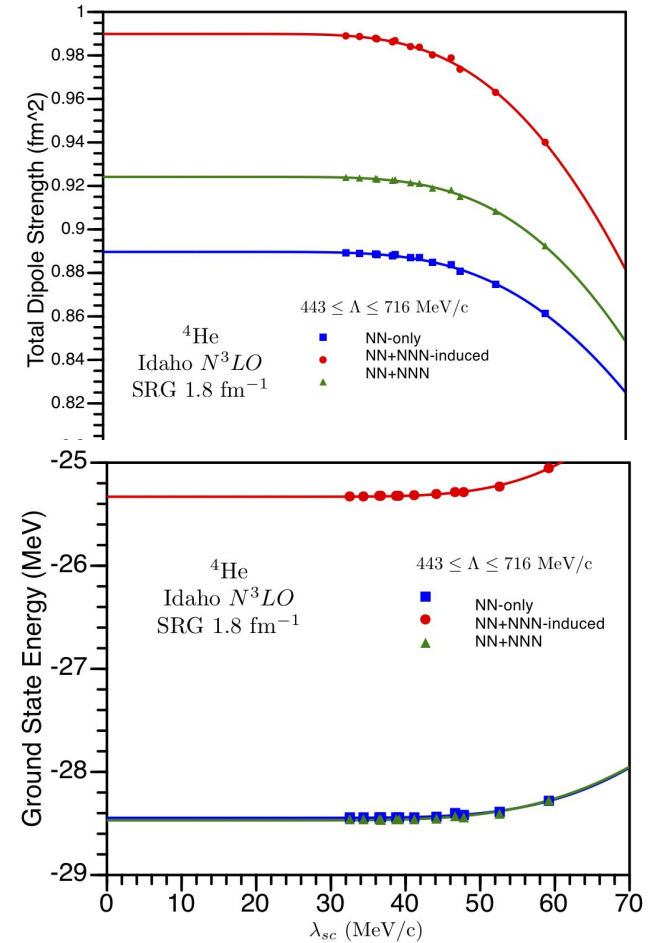
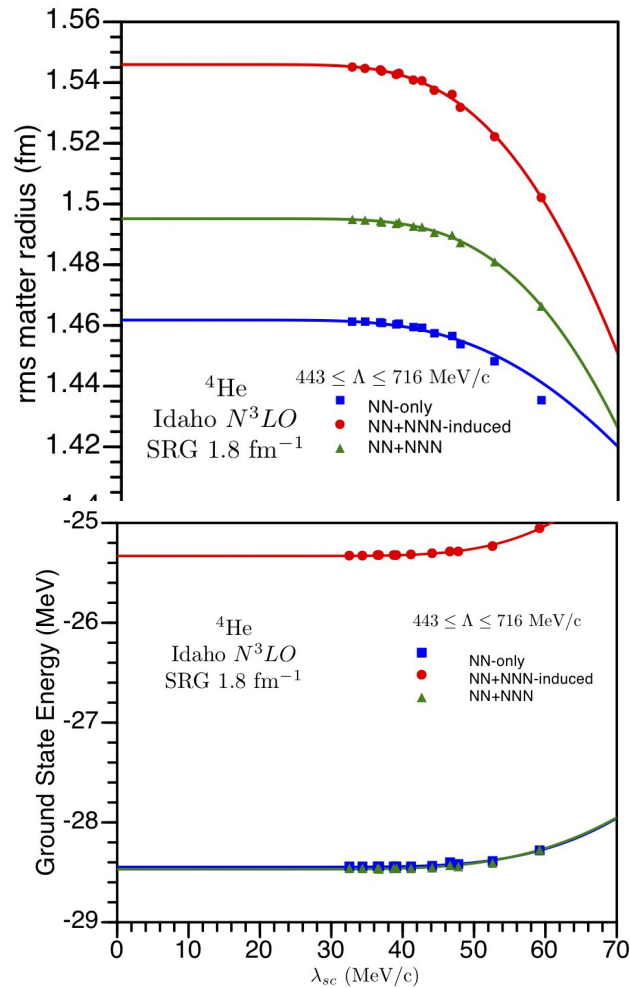


Energy extrapolation

$$E(\lambda_{sc}) = A \exp(-B/\lambda_{sc}) + E(\lambda_{sc} = 0)$$

IR extrapolations of energy, radius, and total dipole moment operator

Convergence of all three operators is the same with λ_{sc} extrapolation



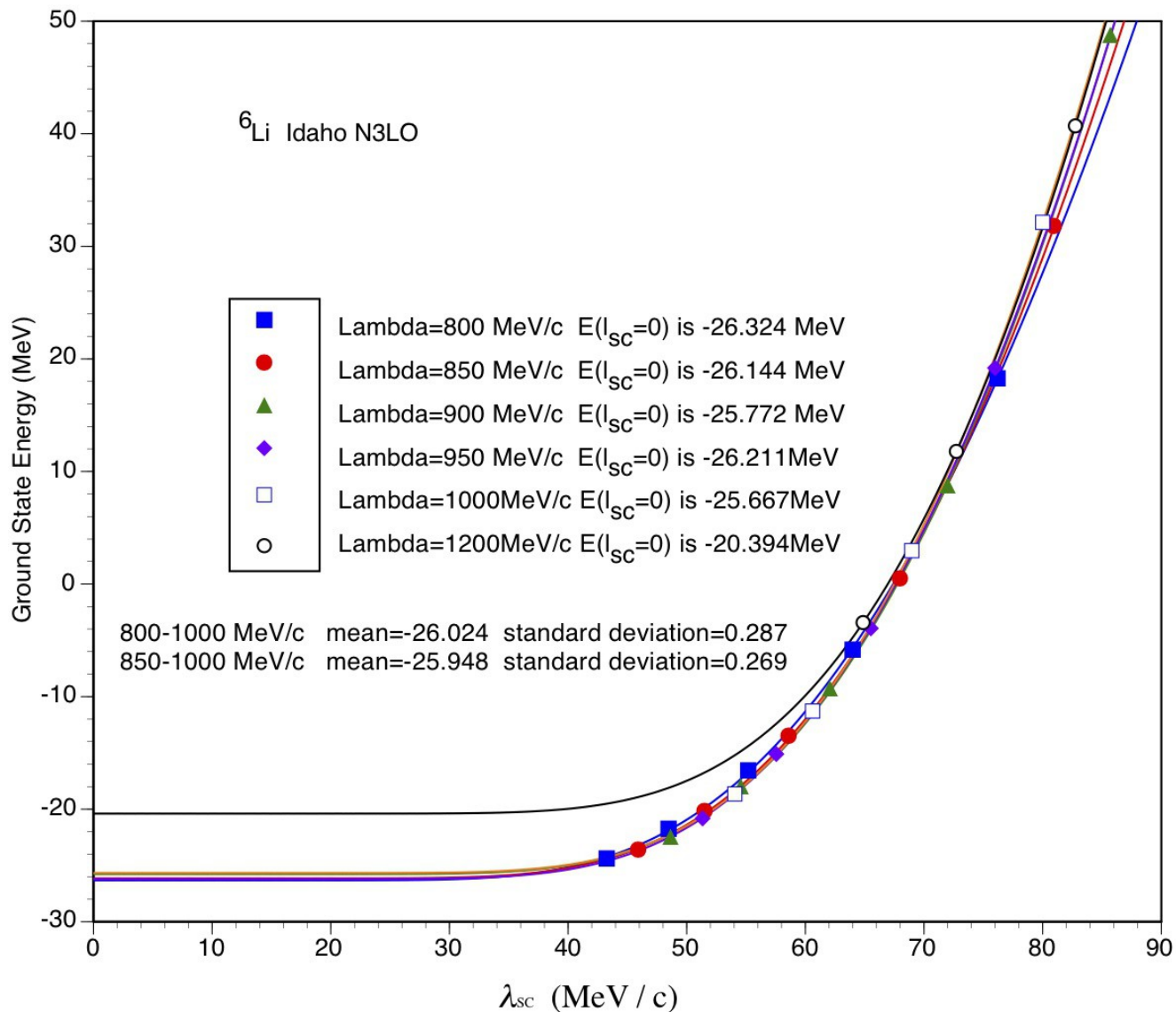
Plot all results at $\hbar\omega = 22$ and 28 MeV and Nmax up to 18

Data from M. D. Schuster, S. Quaglioni, C. W. Johnson and P. Navratil, ArXiv:1304:5491
"Ab Initio many-body calculation of the ${}^4\text{He}$ photo-absorption cross section"

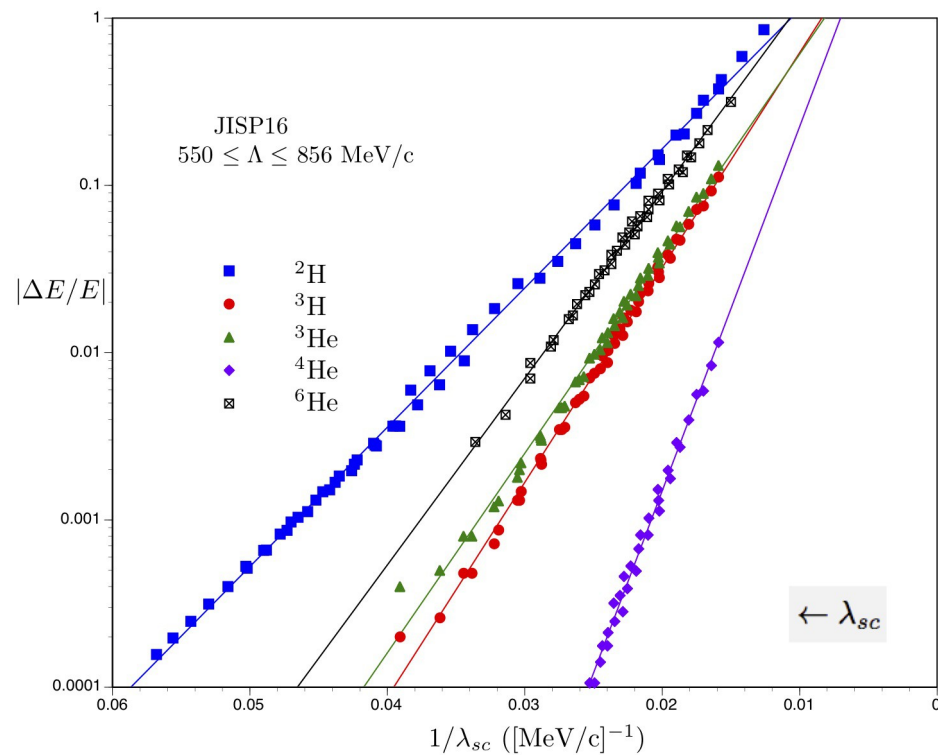
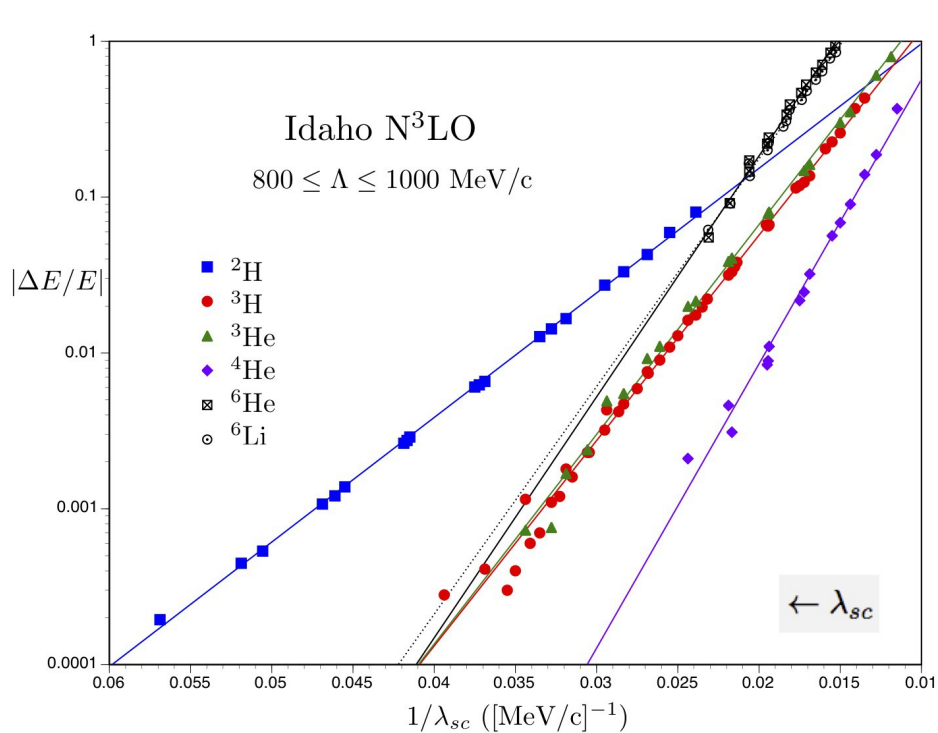
Basis: Nmax up through 16

2004 fixed $\hbar\omega$ extrapolation yields $E_{\text{gs}} = -28.2 \text{ MeV}$

We think IR extrapolation $E_{\text{gs}} = -26.0$ is correct

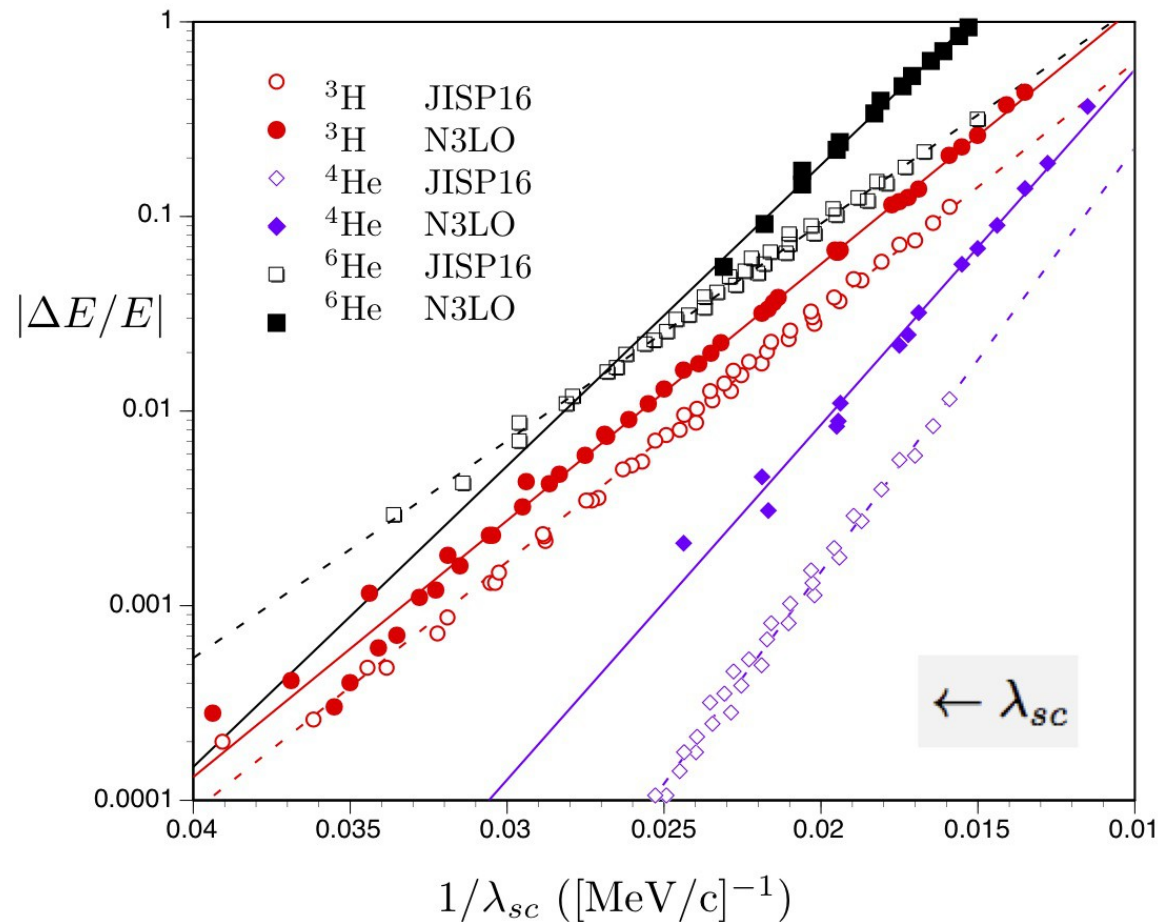


Universal property of IR extrapolation is nucleus dependent



$$\Delta E = E(\lambda_{sc}) - E(\lambda_{sc} = 0) = a \exp(-b/\lambda_{sc})$$

Universal property of IR extrapolation is potential dependent



$$\Delta E = E(\lambda_{sc}) - E(\lambda_{sc} = 0) = a \exp(-b/\lambda_{sc})$$

Single nucleon separation energies and slopes of $\Delta E/E$ vs $1/\lambda_{sc}$ plots

nucleus	exp.	JISP16 N ³ LO		
	S	$4\sqrt{m_N c^2 \bar{S}}$	bc	bc
	(MeV)	(MeV)	(MeV)	(MeV)
² H	2.22457	183	192	184
³ H	6.25723	307	296	304
³ He	5.49348	287	275	310
⁴ He	20.2 (ave)	551	502	420
⁶ He			257	356
⁶ Li	5.0 (ave)	275		336

Alternate λ_{sc} in [Furnstahl et al. PRC **86** \(2012\)](#),
[More et al. PRC **87** \(2013\)](#),
[Furnstahl et al. ArXiv:1312.6876](#)

model the slope “b” in the table by the momentum which corresponds to the separation energy of a single nucleon from the nucleus.

Extra Slides

How do intrinsic regulator scales control needed values of N and $\hbar\omega$ for a converged result?

$$\Lambda/\lambda_{sc} = N + 3/2$$

$$(\Lambda\lambda_{sc})/m_N = \hbar\omega$$

$$\Lambda \geq \Lambda^{NN} = 800$$

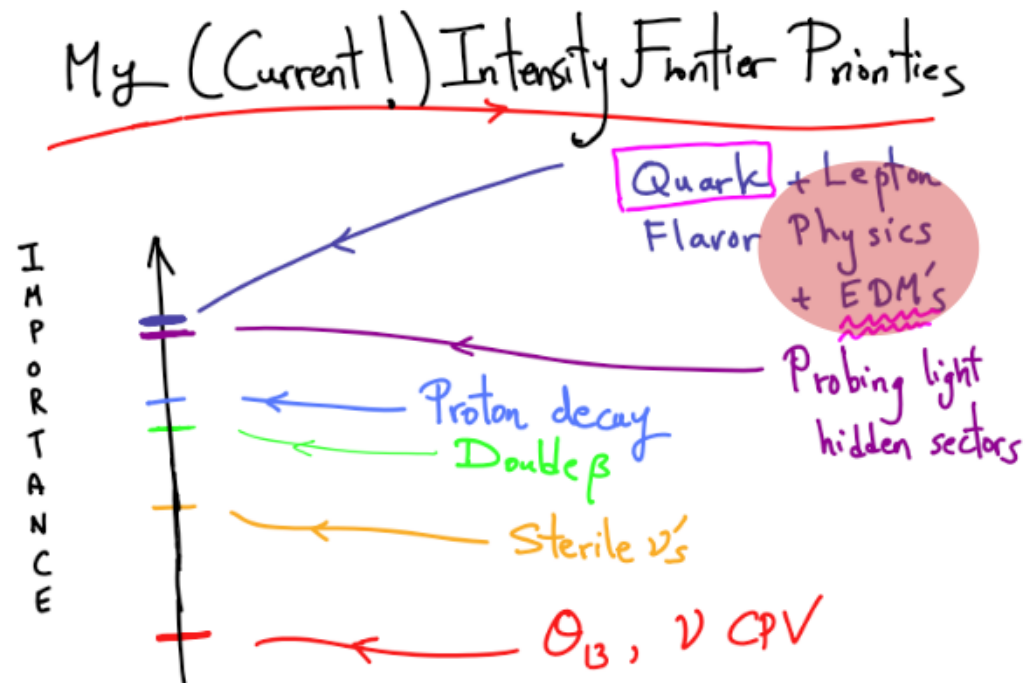
$\lambda_{sc} \approx 10$	$\lambda_{sc} \approx 20$	$\lambda_{sc} \approx 40$
$N \geq 80$	$N \geq 40$	$N \geq 20$
$\hbar\omega \succeq 8$	$\hbar\omega \succeq 16$	$\hbar\omega \succeq 32$

$$\Lambda \geq \Lambda^{NN} = 500$$

$\lambda_{sc} \approx 10$	$\lambda_{sc} \approx 20$	$\lambda_{sc} \approx 40$
$N \geq 50$	$N \geq 25$	$N \geq 12$
$\hbar\omega \succeq 5$	$\hbar\omega \succeq 10$	$\hbar\omega \succeq 20$

Conclusion: One must extrapolate for all but the lightest nuclei

Physics: Potential of EDMs



N. Arkani-Hamed (IAS, Princeton) at Intensity Frontier WS, USA (2011)

EDMs of d, ^3H , ^3He , ^6Li

G. Isidori – Symmetry Physics Implications

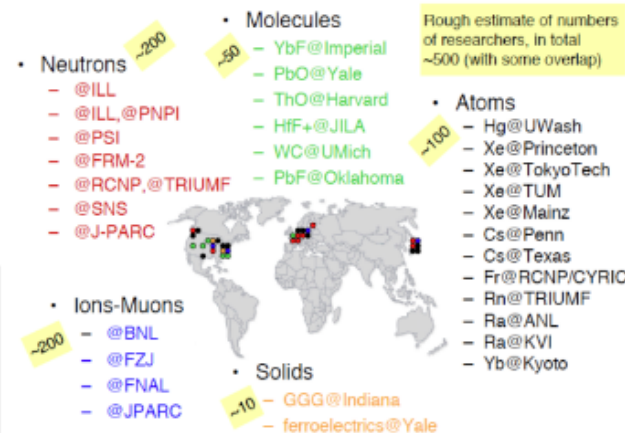
ESPP Open Symposium [Cracow, 10-12 Sep. 2011]

★ The key role of LFV and EDMs

The search for **E**lectric **D**ipole **M**oments of fundamental particles (n, e, μ , ... and, more generally, atoms or heavy nuclei), share the three main virtues of LFV searches:

- We know CP is not an exact symmetry of nature \Rightarrow non-vanishing EDMs
- Virtually no problems of SM backgrounds
- EDMs close to the present bounds in several realistic models

Observable	Exp. Current
$ d_{Tl} $ [e cm]	$< 9.0 \times 10^{-25}$
$ d_{Hg} $ [e cm]	$< 3.1 \times 10^{-29}$
$ d_n $ [e cm]	$< 2.9 \times 10^{-26}$



world-wide effort in trying to improve the limits by ~ 1 order of magnitude

G. Isidori at ESPP Open Symposium, Cracow (Sept. 2012)

EDM of charged particles: use of storage rings

PROCEDURE

- Place particles in a storage ring
- Align spin along momentum (\rightarrow freeze horizontal spin precession)
- Search for time development of vertical polarization

