Role of the particle continuum and of threenucleon forces in oxygen isotopes

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### Overview

- 1. Introduction
- 2. Neutron-rich oxygen isotopes
- 3. Some comments on benchmark results
- 4. Summary

### Energy scales and relevant degrees of freedom



Energy

### Nuclear forces from chiral effective field theory [Weinberg; van Kolck; Epelbaum, Krebs, Meissner; Entem & Machleidt; ...]



## Is <sup>28</sup>O a bound nucleus?

#### **Experimental situation**

- "Last" stable oxygen isotope <sup>24</sup>O
- <sup>25</sup>O unstable (Hoffman et al 2008)
- <sup>26</sup>O unstable (Lunderberg et al 2012)
- <sup>31</sup>F exists (adding on proton shifts drip line by 6 neutrons!?)

#### Theoretical description challenging (proximity of continuum, 3NFs, subtle effects)

- Continuum shell model [Volya & Zelevinsky, PRL 94, 052501 (2005)]
- Chiral NN interactions [Hagen et al., PRC 80, 021306 (2009)].
- Effects of 3NFs [Otsuka, Suzuki, Holt, Schwenk, Akaishi, PRL 105, 032501 (2010)]
- More complete calculation necessary (3NFs, continuum, large model space, minimum adjustments to interaction)

### Coupled-cluster method (in CCSD approximation)

Ansatz:

$$|\Psi\rangle = e^{T}|\Phi\rangle$$
  

$$T = T_{1} + T_{2} + \dots$$
  

$$T_{1} = \sum_{ia} t_{i}^{a} a_{a}^{\dagger} a_{i}$$
  

$$T_{2} = \sum_{ijab} t_{ij}^{ab} a_{a}^{\dagger} a_{b}^{\dagger} a_{j} a_{i}$$

- Scales gently (polynomial) with increasing problem size o<sup>2</sup>u<sup>4</sup>.
- $\ensuremath{\textcircled{\odot}}$  Truncation is the only approximation.
- © Size extensive (error scales with A)

⊗ Most efficient for doubly magic nuclei

Correlations are *exponentiated* 1p-1h and 2p-2h excitations. Part of np-nh excitations included!



Coupled cluster equations  $E = \langle \Phi | \overline{H} | \Phi \rangle$   $0 = \langle \Phi_i^a | \overline{H} | \Phi \rangle$   $0 = \langle \Phi_{ij}^{ab} | \overline{H} | \Phi \rangle$  $\overline{H} \equiv e^{-T} H e^T = \left( H e^T \right)_c = \left( H + HT_1 + HT_2 + \frac{1}{2} HT_1^2 + ... \right)_c$ 

### Toward medium-mass nuclei Chiral N<sup>3</sup>LO (500 MeV) by Entem & Machleidt, NN only



- Chiral NN forces yield saturation, lack about 0.4 MeV per nucleon in binding energy.
- Chiral three-nucleon forces expected to yield 0.4 MeV per nucleon?!

#### Binding energy per nucleon

Nucleus	CCSD	A-CCSD(T)	Experiment
⁴He	5.99	6.39	7.07
<sup>16</sup> O	6.72	7.56	7.97
<sup>40</sup> Ca	7.72	8.63	8.56
<sup>48</sup> Ca	7.40	8.28	8.67

Our CC results for <sup>16</sup>O are confirmed by UMOA [Fujii et al., PRL 103, 182501 (2009)] and by IT-NCSM [R. Roth et al., PRL 107, 072501 (2011)]

	ССМ	(IT)-NCSM	UMOA
	E/A	E/A	E/A
⁴He	-6.39(5)	-6.35	
<sup>16</sup> O	-7.56(8)	-7.48(4)	-7.47

[Hagen, TP, Dean, Hjorth-Jensen, PRL 101, 092502 (2008)]

### Gamow shell model



(N. Michel et al, PRL 89 (2002) 042502)



#### complex-symmetric eigenvalue problem for hermitian hamiltonian

Review: Michel et al, J. Phys. G 36, 013101 (2009)

### How to economically include the effects of 3NFs?

Monopole shifts of shell-model interactions from 3NF as densitydependent NN force. [A.Zuker, PRL 90, 42502 (2003)]

3NFs included as normal-ordered contributions [Hagen, TP, Dean, Schwenk, Nogga, Wloch, Piecuch, PRC 76, 034302 (2007); Hebeler & Schwenk PRC 82, 014314 (2010); J.W. Holt, Kaiser, Weise PRC, 024002 (2010); Otsuka, Suzuki, J. D. Holt, Schwenk, Akaishi, PRL 105 032501 (2010)]

In-medium modifications of NN interactions (Brown-Rho scaling) [J. W. Holt, Brown, Kuo, J. D. Holt, Machleidt, PRL 100, 062501 (2008)]

Our approach: In-medium modifications of NN forces, inspired by EFT



3NFs as in-medium effective two-nucleon forces

**Parameters**:  $k_F = 1.0 \text{ fm}^{-1}$ ,  $c_E = 0.7$ ,  $c_D = -0.2$  from binding energies of  $^{16,22}O$ 

**Rationale**: The integration over symmetric nuclear matter is an approximation for finite nuclei and for nonzero isospin. Tuning of parameters generates corrections and takes into account higher-order effects.

# Can one do this?

Q1: Should one not keep the low-energy coefficients  $c_E$ ,  $c_D$  as determined in light systems?

A1: The summation over symmetric nuclear matter is an approximation for finite nuclei, and adjustments in  $c_E$  aim to correct for this approximation.

Q2: What about the combinatorial factors [Hebeler & Schwenk 2010]?

A2: A1

Q3: Will this work in <sup>4</sup>He?

A3: This approximation is probably limited to heavier nuclei.

Let's see, to what extent this works in practice.



Ground-state energies of oxygen isotopes

### Spectra (1p-1h excitations only)





## Matter and charge radii



Data: [Ozawa et al., NPA 693, 32 (2001); Kanungo et al., PRC 84, 061304(R) (2011)]

### Summary & Outlook

- Proposed and tested efficient approximation for the inclusion of 3NFs; in keeping with spirit of EFT and a sense of reality
- Much-improved binding energies, separation energies, and spectra
- <sup>24</sup>O is dripline nucleus with several long-lived resonances at ~7 MeV of excitation energy
- Continuum effects yield about 0.3 MeV of additional binding for resonances and bound states close to threshold
- Progress in computational capabilities since 2001 benchmark
- Agreement on error estimates necessary for credibility and impact