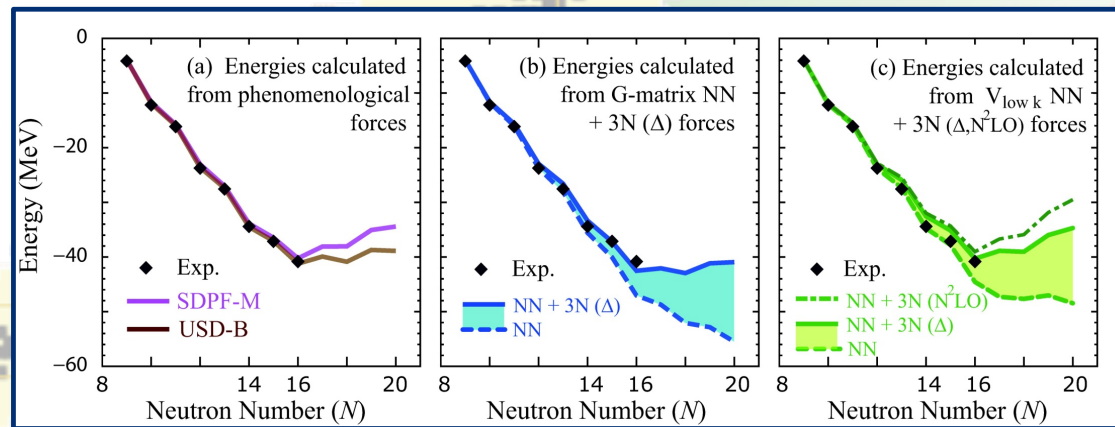


Three-Nucleon Forces and the Structure of Exotic Nuclei

Jason D. Holt

THE UNIVERSITY of
TENNESSEE **UT**
KNOXVILLE

126



OAK RIDGE NATIONAL LABORATORY

Drip Lines and Magic Numbers: 3N Forces in Medium-Mass Nuclei

Important in light nuclei, nuclear matter...

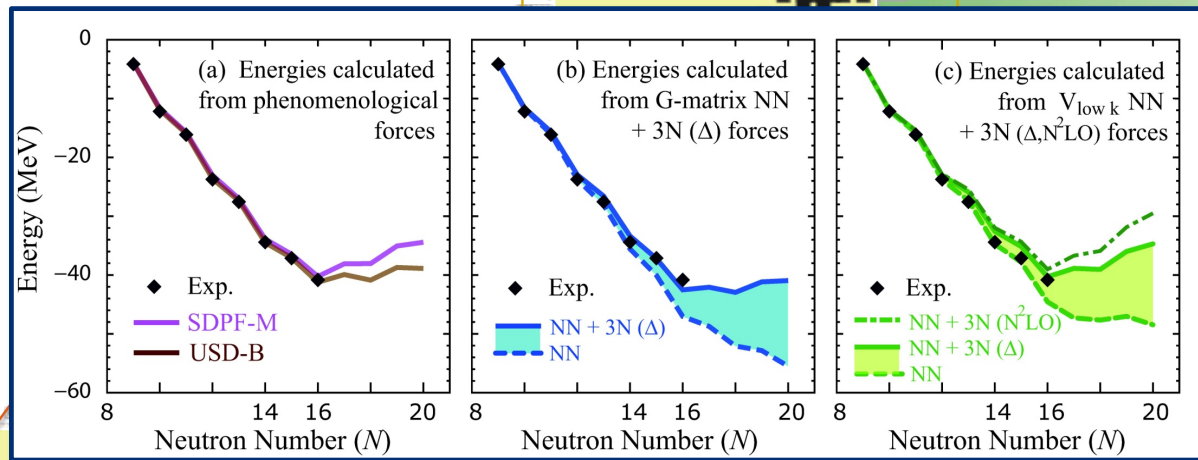
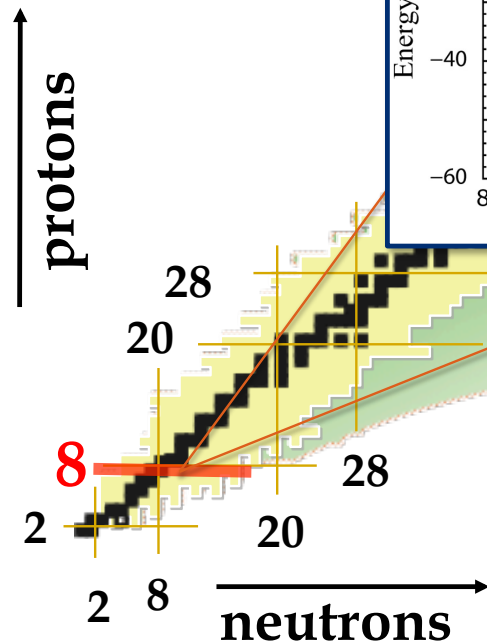
What are the limits of nuclear existence?

How do magic numbers form and evolve?

Otsuka, Suzuki, Holt, Schwenk, Akaishi, PRL (2010)

82

Heaviest oxygen isotope



Drip Lines and Magic Numbers: 3N Forces in Medium-Mass Nuclei

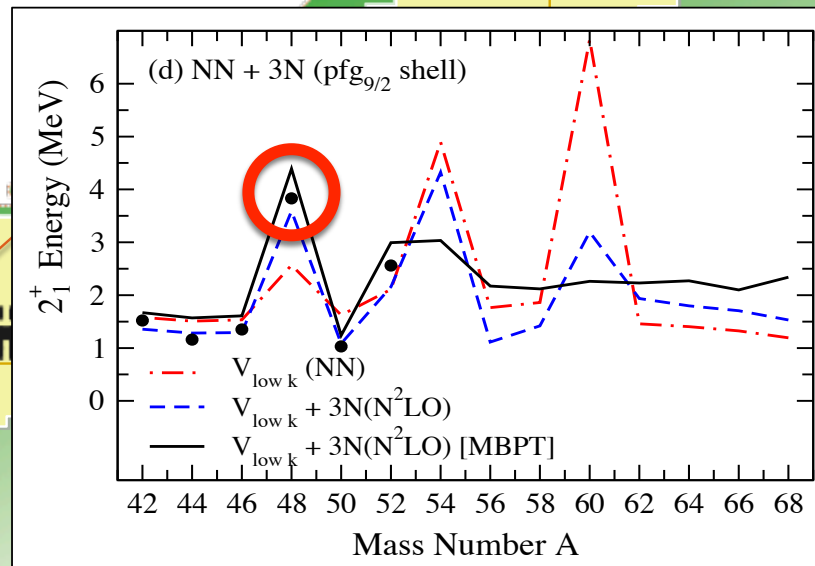
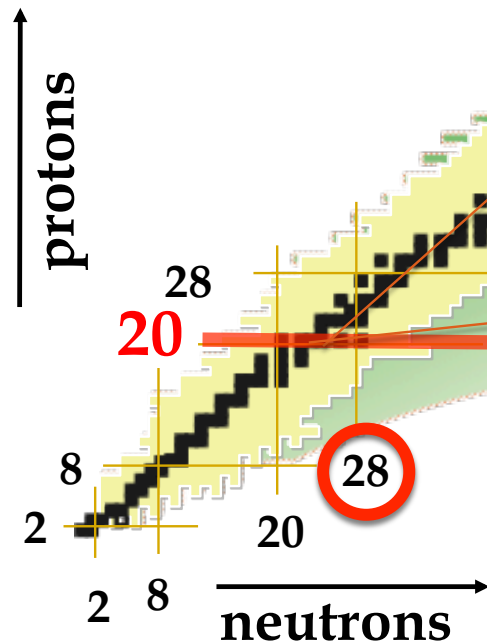
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N=28 magic number in calcium

Holt, Otsuka, Schwek,
Suzuki, arXiv:1009.5984

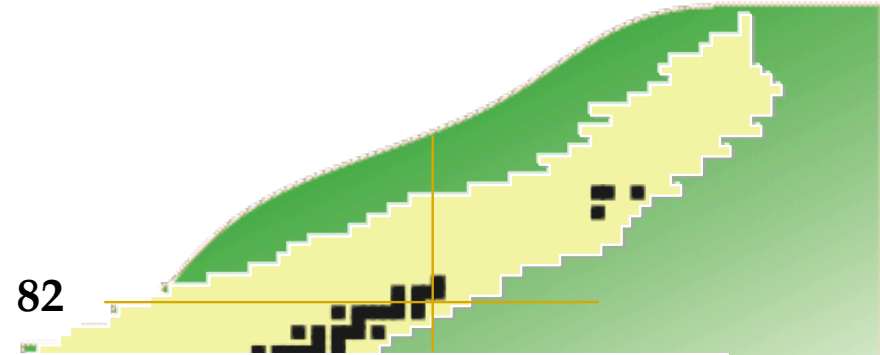


Drip Lines and Magic Numbers: 3N Forces in Medium-Mass Nuclei

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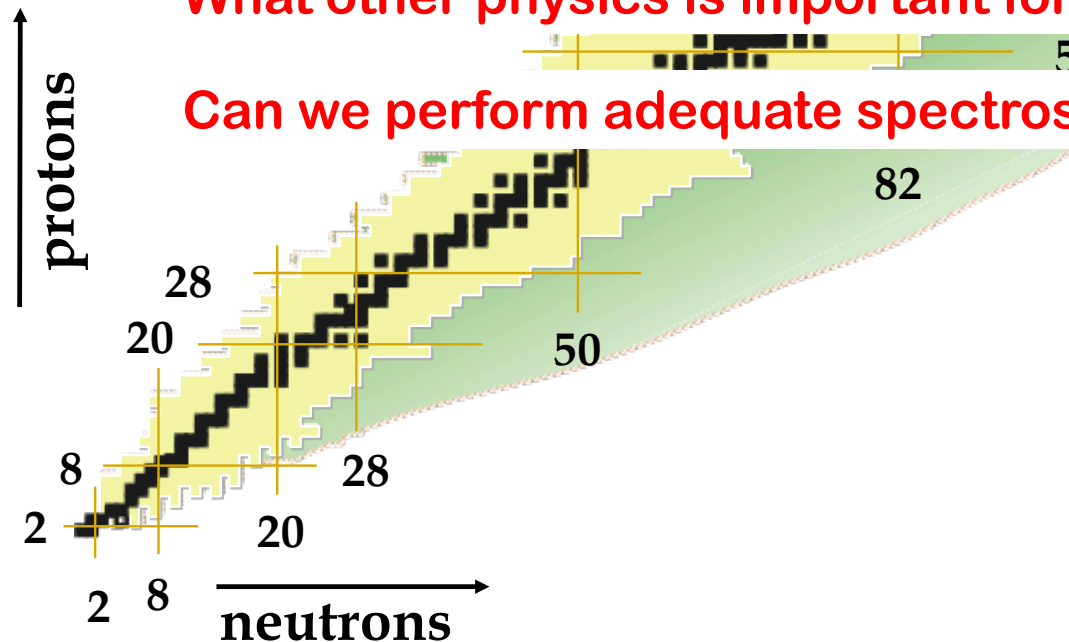
How do magic numbers form and evolve?



**Advances in microscopic theory:
Calculate all theoretical inputs consistently with NN+3N**

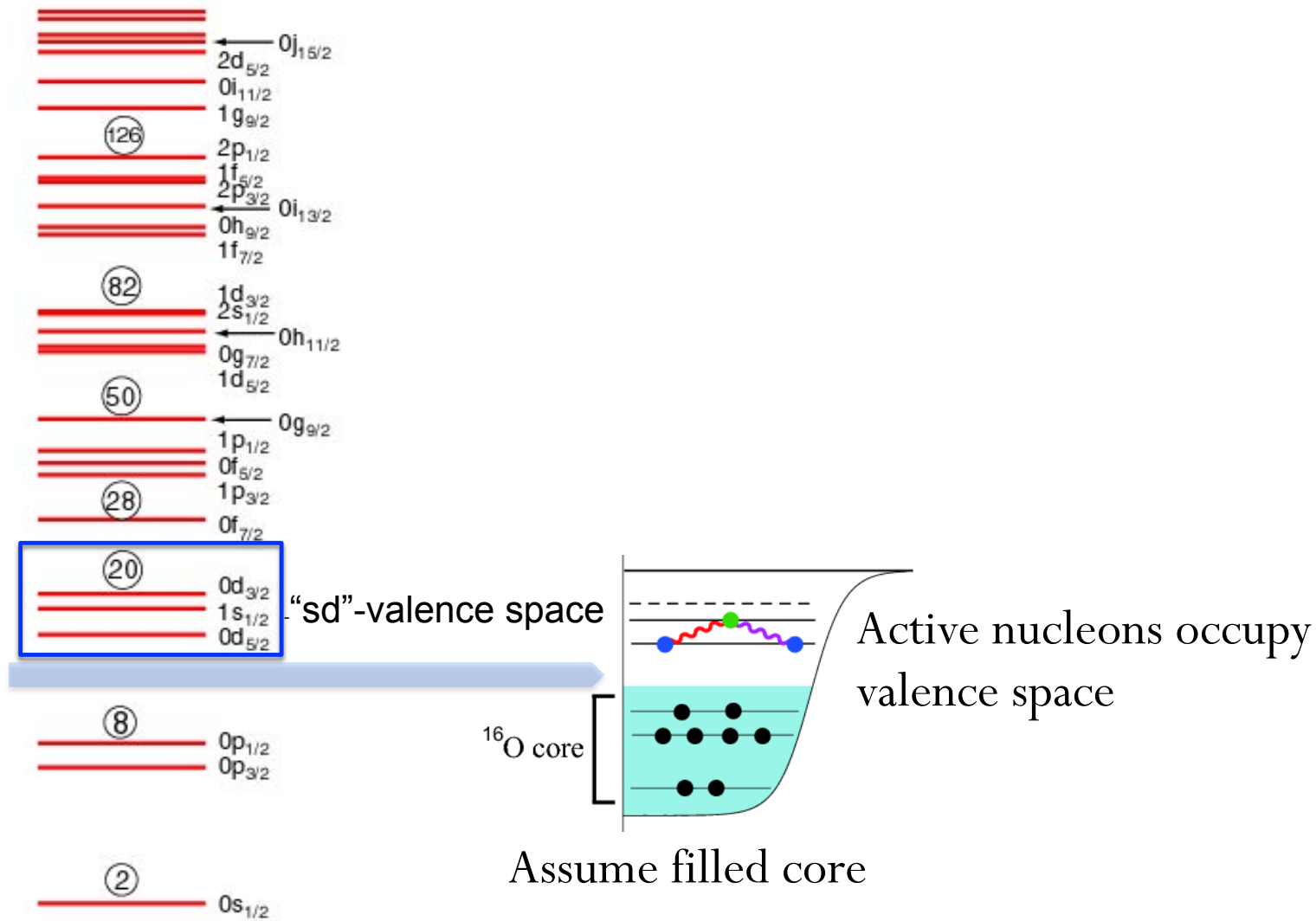
What other physics is important for neutron-rich systems?

Can we perform adequate spectroscopic calculations?



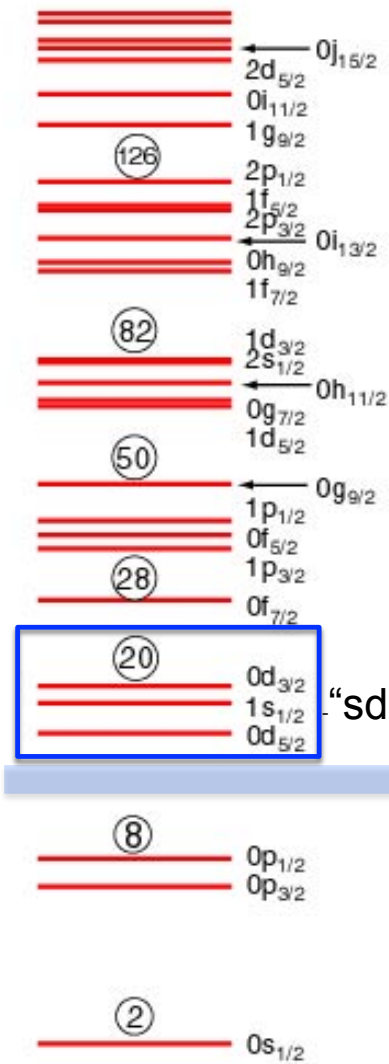
Solving the Nuclear Many-Body Problem

Nuclei understood as many-body system starting from closed shell, add nucleons



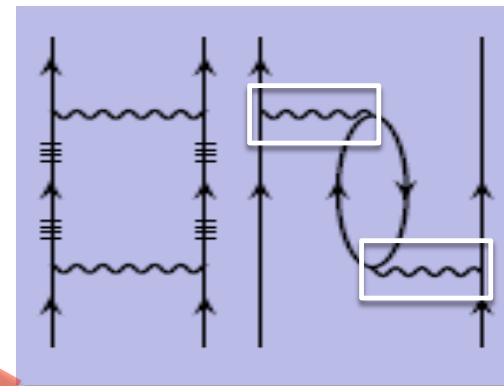
Solving the Nuclear Many-Body Problem

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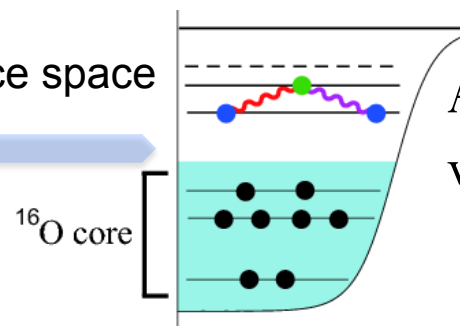


Strong interactions with core generate **effective interaction** between valence nucleons

Hjorth-Jensen, Kuo, Osnes (1995)



"sd"-valence space



Active nucleons occupy valence space

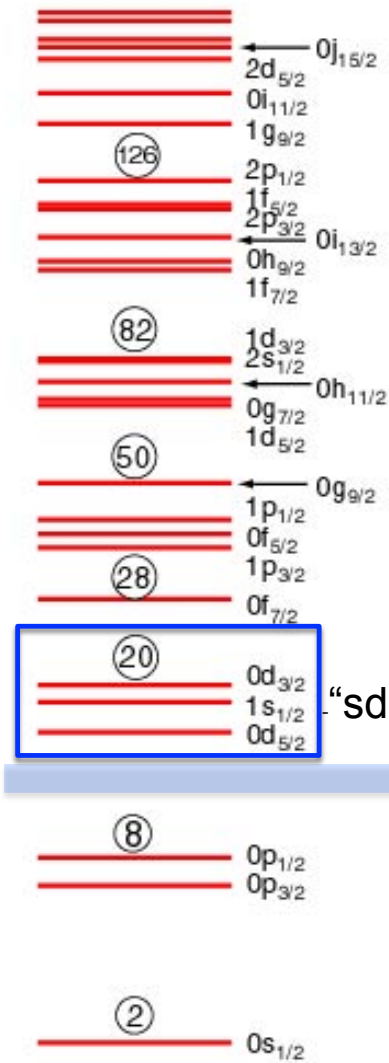
Assume filled core

Solving the Nuclear Many-Body Problem

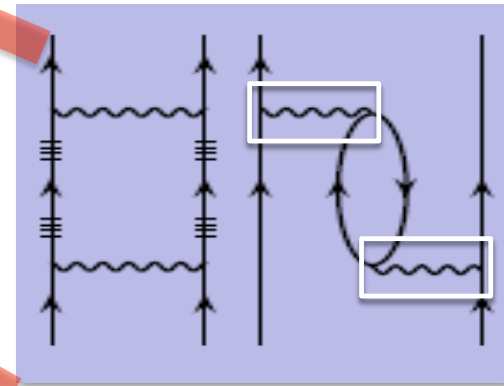
Nuclei understood as many-body system starting from closed shell, add nucleons
3rd-order MBPT: intermediate states to 13 major shells

Effective two-body interaction
Single-particle energies (SPEs)

Hjorth-Jensen, Kuo, Osnes (1995)

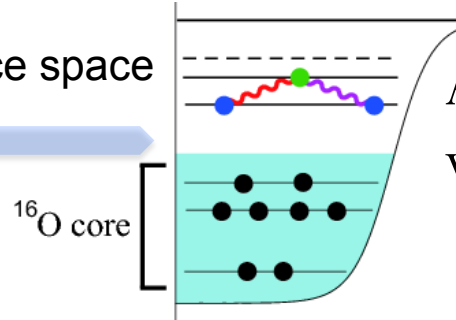


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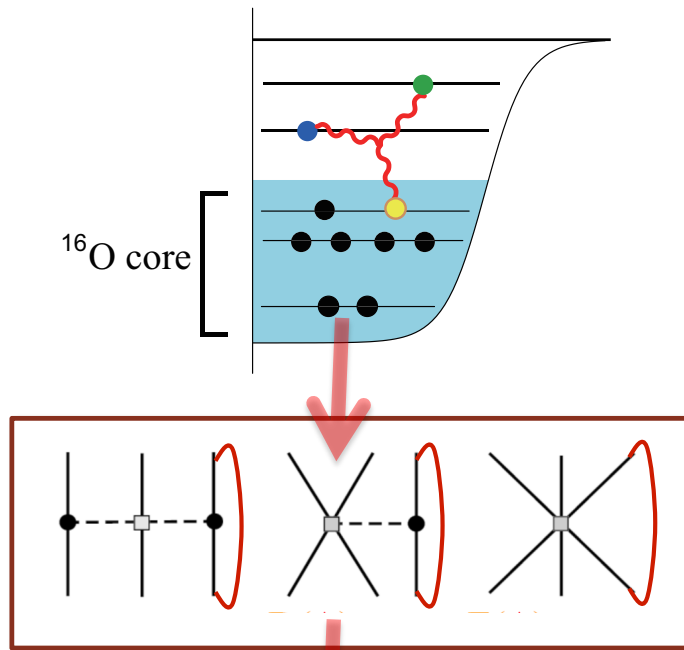


Assume filled core

3N Forces for Valence-Shell Theories

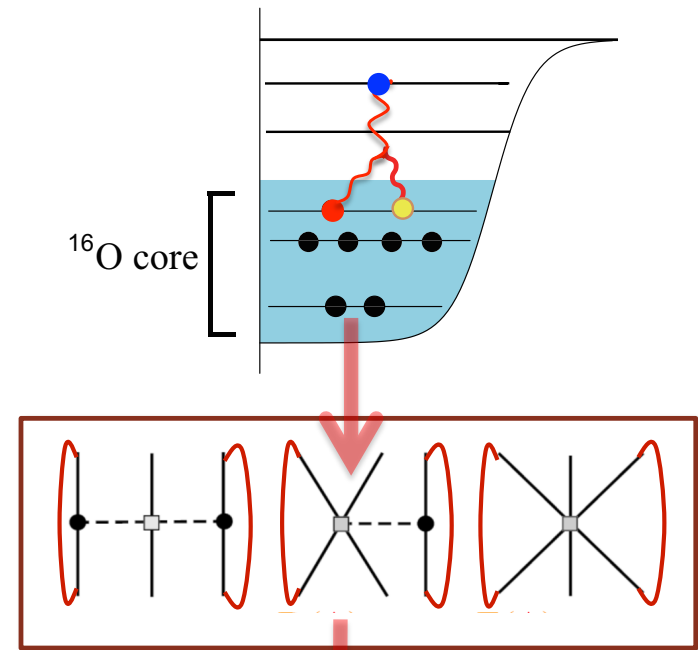
Normal-ordered 3N: contribution to valence neutron interactions

Effective two-body



$$\langle ab | V_{3N,\text{eff}} | a'b' \rangle = \sum_{\alpha=\text{core}} \langle \alpha ab | V_{3N} | \alpha a'b' \rangle$$

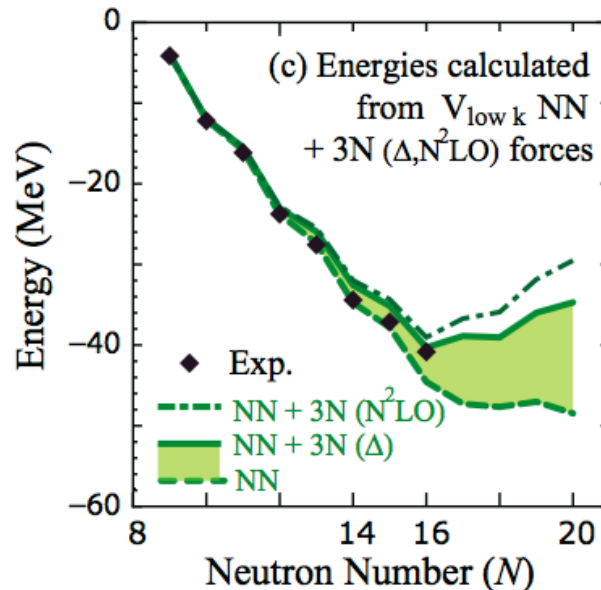
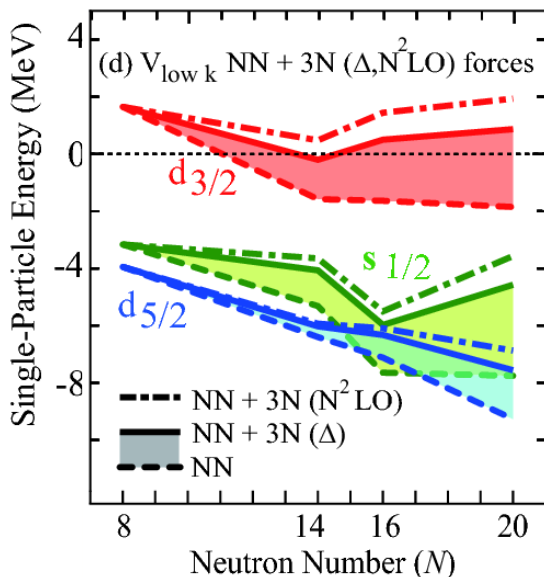
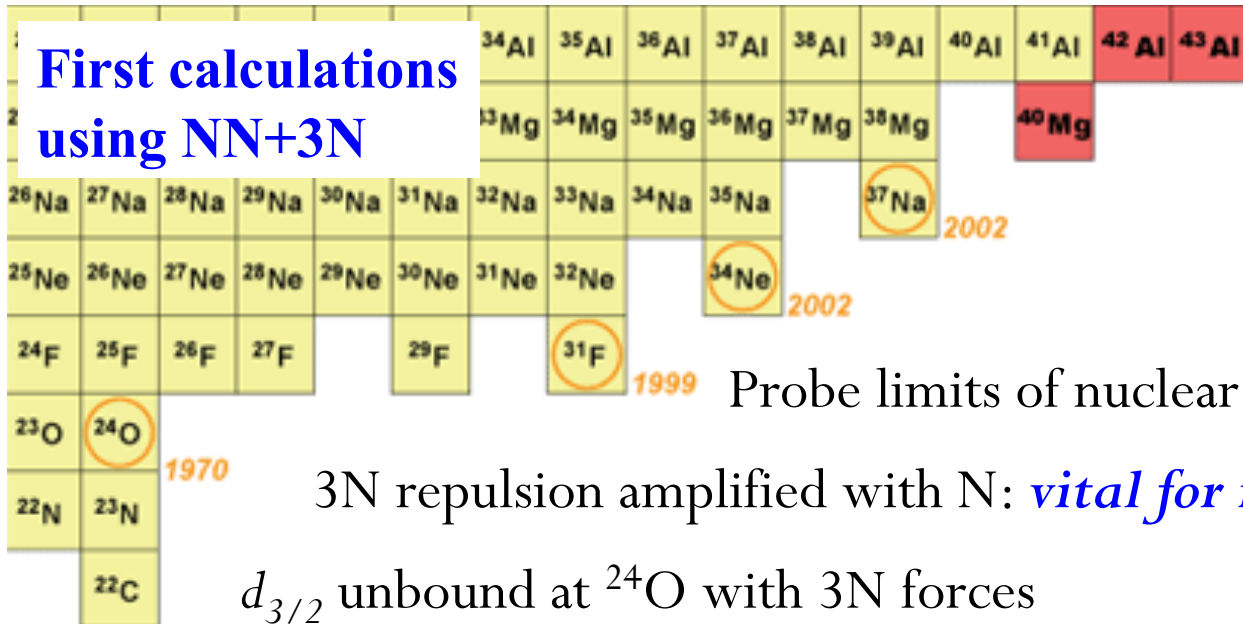
Effective one-body



$$\langle a | V_{3N,\text{eff}} | a' \rangle = \frac{1}{2} \sum_{\alpha\beta=\text{core}} \langle \alpha\beta a | V_{3N} | \alpha\beta a' \rangle$$

Combine with microscopic NN: eliminate empirical adjustments

Oxygen Anomaly



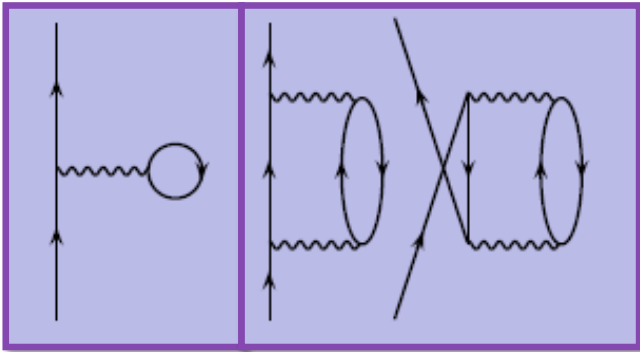
Isotopes unbound beyond ^{24}O

First microscopic explanation of oxygen anomaly

Otsuka, Suzuki, Holt, Schwenk, Akaishi, PRL (2010)

One-Body 3N: Single Particle Energies

NN-only microscopic SPEs yield poor results – rely on empirical adjustments

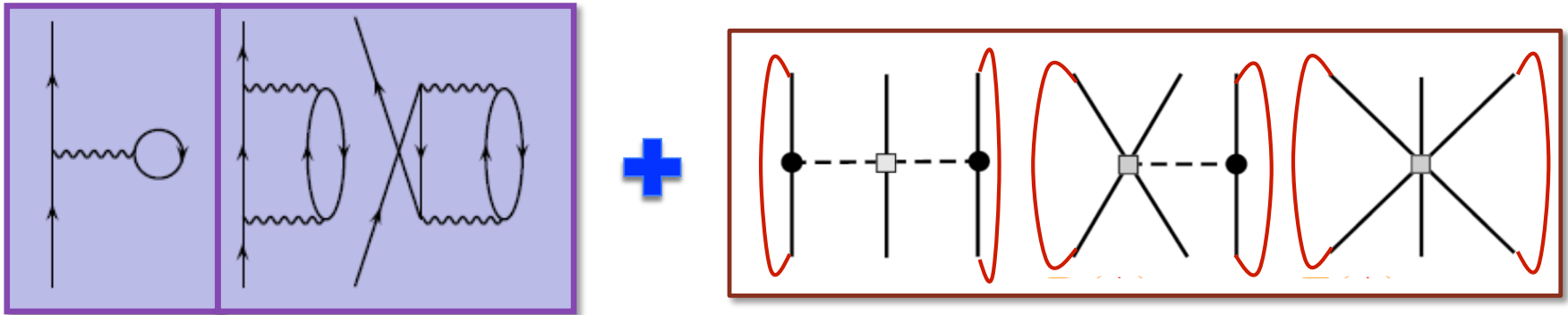


sd-shell: SPEs much too bound, unreasonable splitting

Orbit	“Exp”	USD b	$T+V_{NN}$
$d_{5/2}$	-4.14	-3.93	-5.43
$s_{1/2}$	-3.27	-3.21	-5.32
$d_{3/2}$	0.944	2.11	-0.97

One-Body 3N: Single Particle Energies

NN-only microscopic SPEs yield poor results – rely on empirical adjustments



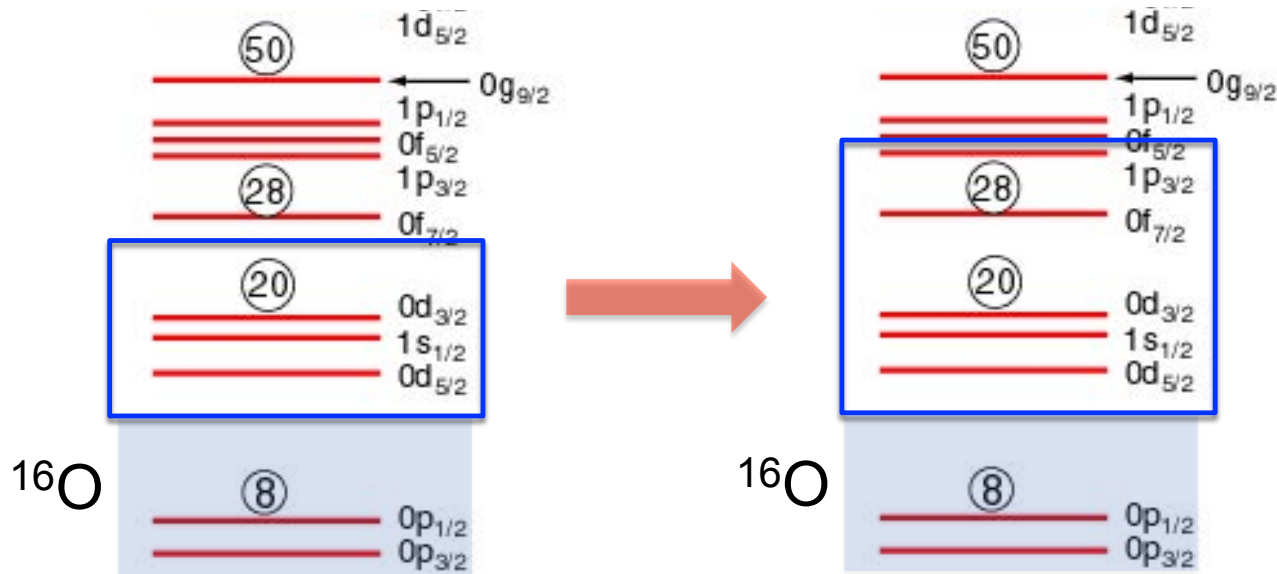
sd-shell: SPEs much too bound, unreasonable splitting

3N forces: additional repulsion – reasonable values!

Orbit	USD b	$T+V_{NN}+V_{3N}$
$d_{5/2}$	-3.93	-3.82
$s_{1/2}$	-3.21	-2.14
$d_{3/2}$	2.11	2.01

One-Body 3N: Single Particle Energies

Effects of correlations beyond one major oscillator shell:



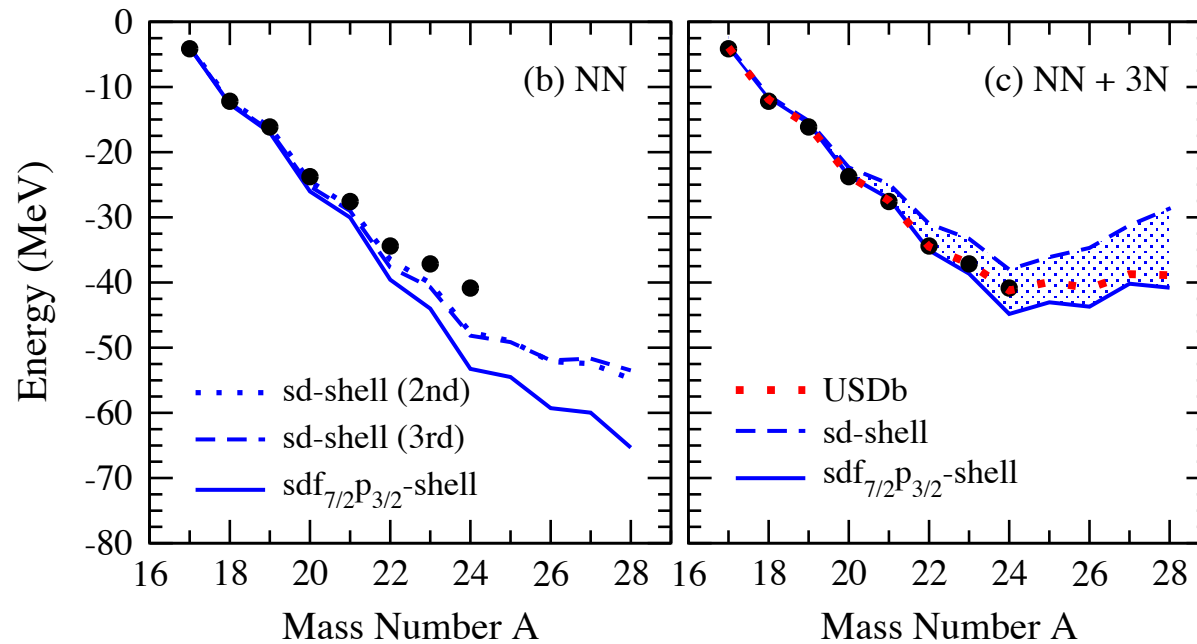
Orbit	USD ^b	$T+V_{NN}+V_{3N}$	SDPF-M	$T+V_{NN}+V_{3N}$
$d_{5/2}$	-3.93	-3.82	-3.95	-3.75
$s_{1/2}$	-3.21	-2.14	-3.16	-2.10
$d_{3/2}$	2.11	2.01	1.65	2.13
$f_{7/2}$			3.10	2.96
$p_{3/2}$			3.10	4.82

Fully microscopic framework and extended valence space

Fully-Microscopic Calculations

Interaction and self-consistent SPEs from NN+3N

Empirical SPEs for NN-only



Holt, Schwenk, arXiv:1108.2680

NN-only: dripline at ^{28}O

NN+3N: dripline at ^{24}O

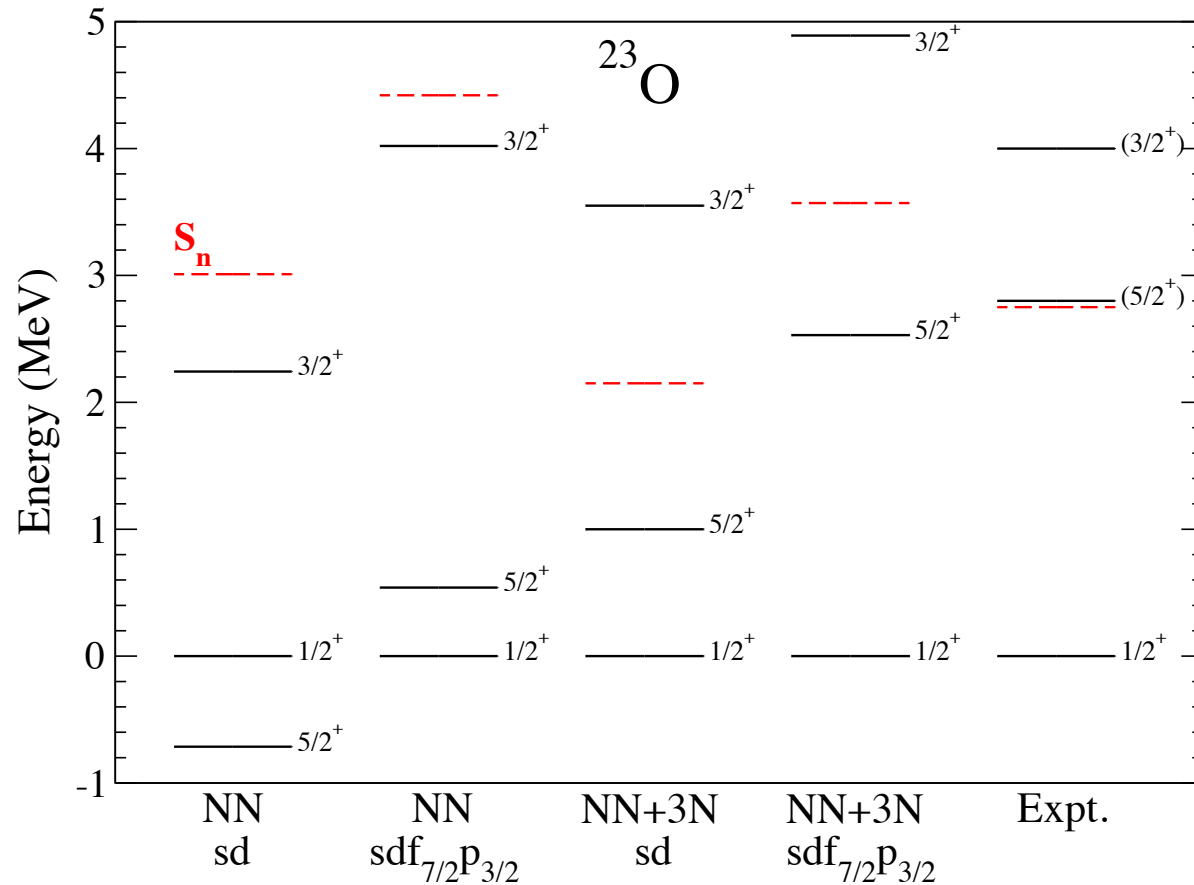
sd-shell results underbound; improved in $sdf_{7/2} p_{3/2}$

Continuum: $\sim 300\text{keV}$ more binding beyond ^{24}O (from CC)

Impact on Spectra: ^{23}O

Neutron-rich oxygen spectra with NN+3N

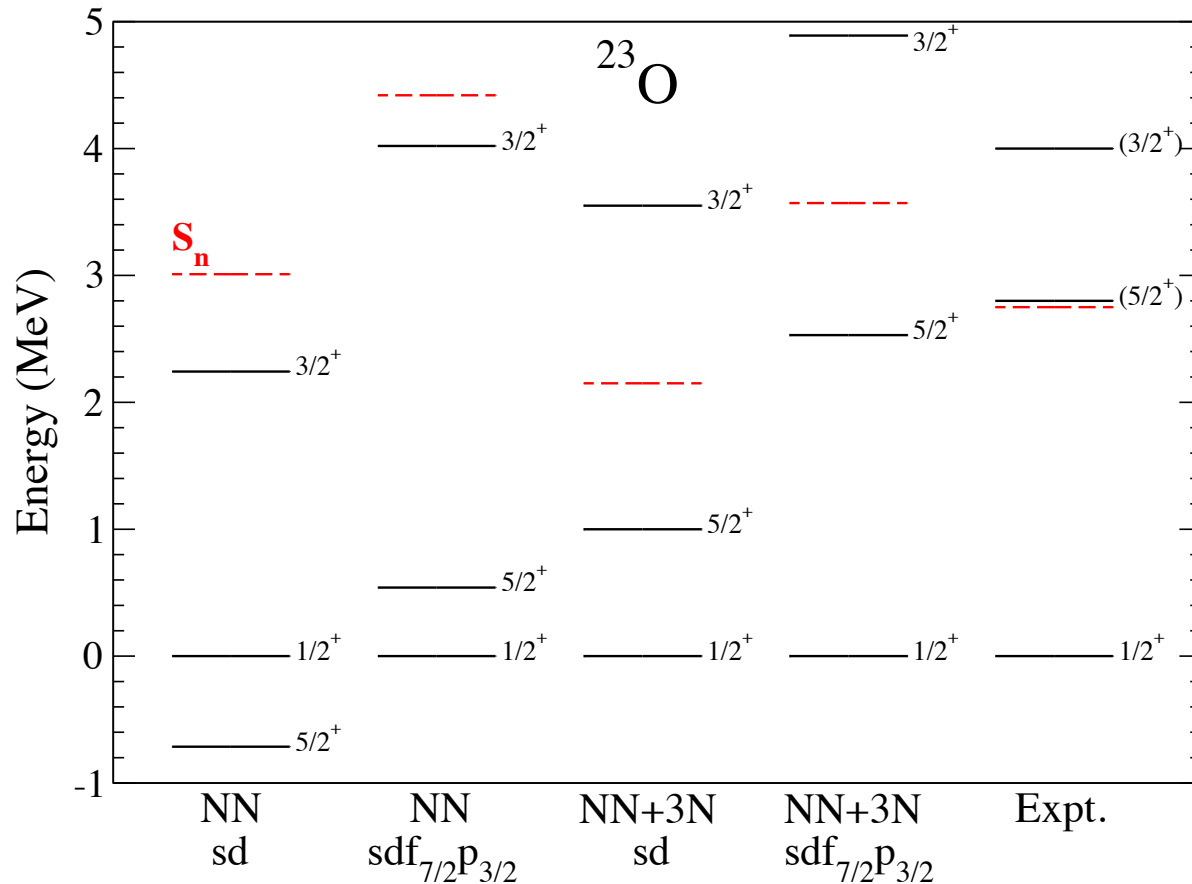
$5/2^+$, $3/2^+$ indicate position of $d_{5/2}$ and $d_{3/2}$ orbits



Impact on Spectra: ^{23}O

Neutron-rich oxygen spectra with NN+3N

$5/2^+$, $3/2^+$ indicate position of $d_{5/2}$ and $d_{3/2}$ orbits



sd-shell NN-only

Wrong ground state!

$5/2^+$ too low

$3/2^+$ bound

Microscopic NN+3N

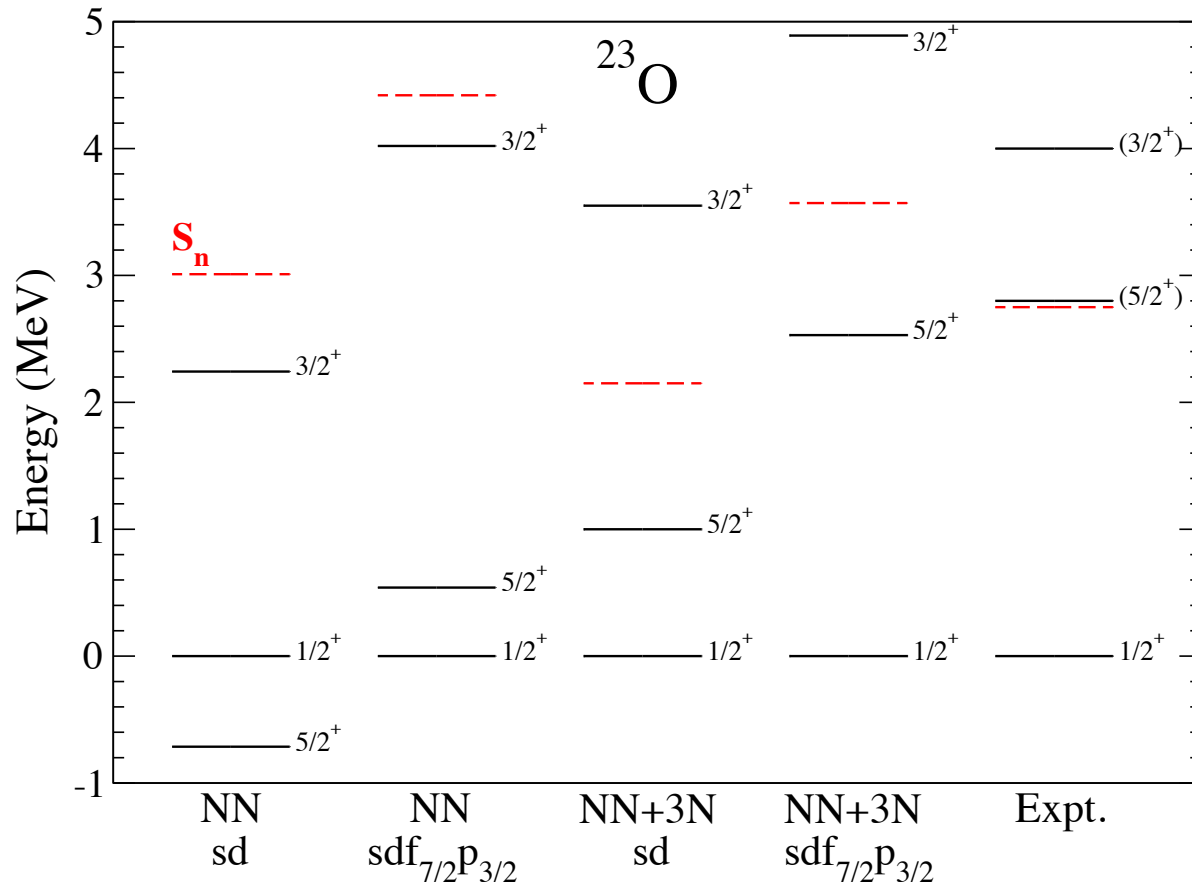
Great improvements in extended valence space!

Holt, Schwenk, arXiv:1108.2680

Impact on Spectra: ^{23}O

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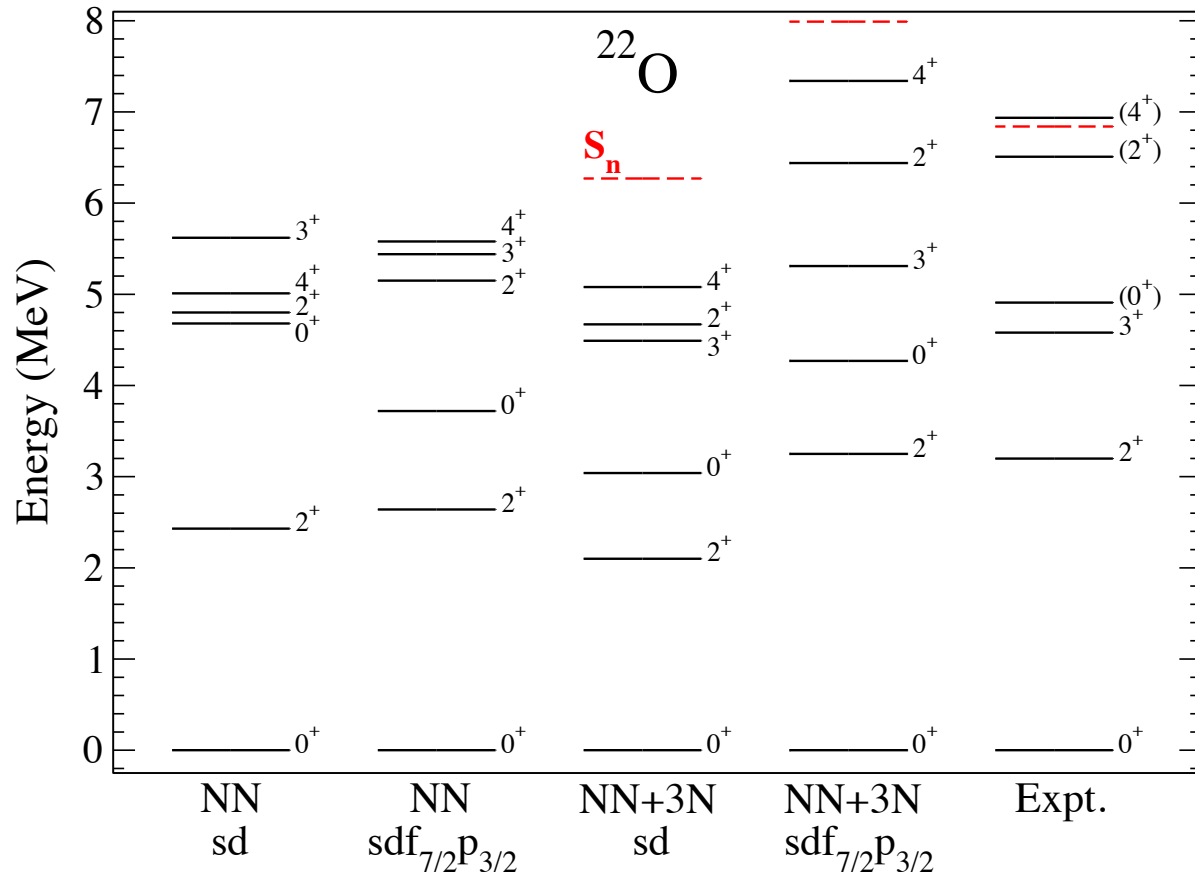
Holt, Schwenk, arXiv:1108.2680

Continuum effectively lowers $3/2^+$ - vital for $^{24-28}\text{O}$

Impact on Spectra: ^{22}O

Neutron-rich oxygen spectra with NN+3N

^{22}O : $N=14$ new magic number – not reproduced with NN



NN-only

2⁺ 500 keV too low

Spectrum too compressed

Microscopic NN+3N

Extended space essential

Reproduces $N=14$ magic number

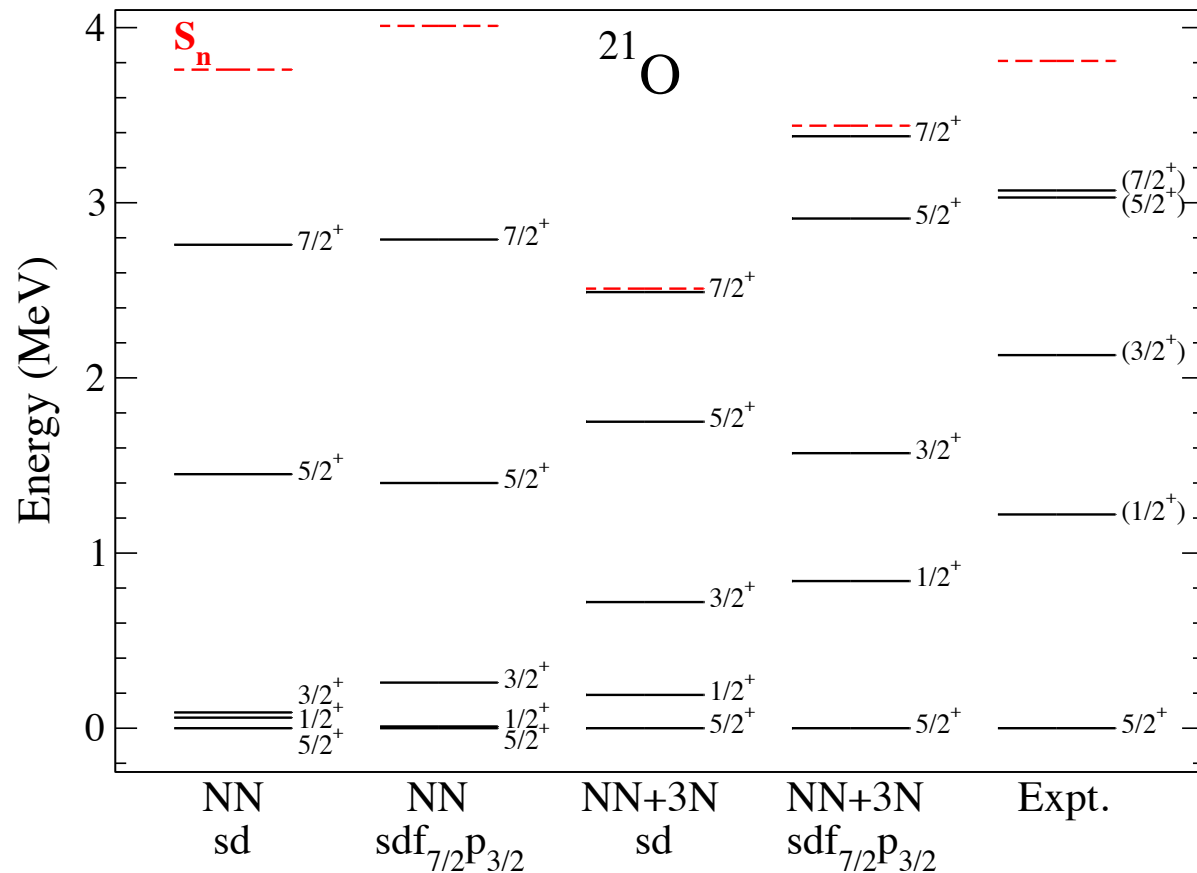
Holt, Schwenk, arXiv:1108.2680

Contributions from 3N and extended valence orbitals important

Impact on Spectra: ^{21}O

Neutron-rich oxygen spectra with NN+3N

Spectrum sensitive to $s_{1/2}$ shell closure



NN-only

Low-lying states

too compressed

$7/2^+ - 5/2^+$ too wide

Microscopic NN+3N

Improvement in sd

Extended orbits essential

Improved spacing in all levels

Holt, Schwenk, arXiv:1108.2680

3N improvements largely due to higher calculated $s_{1/2}$ orbital

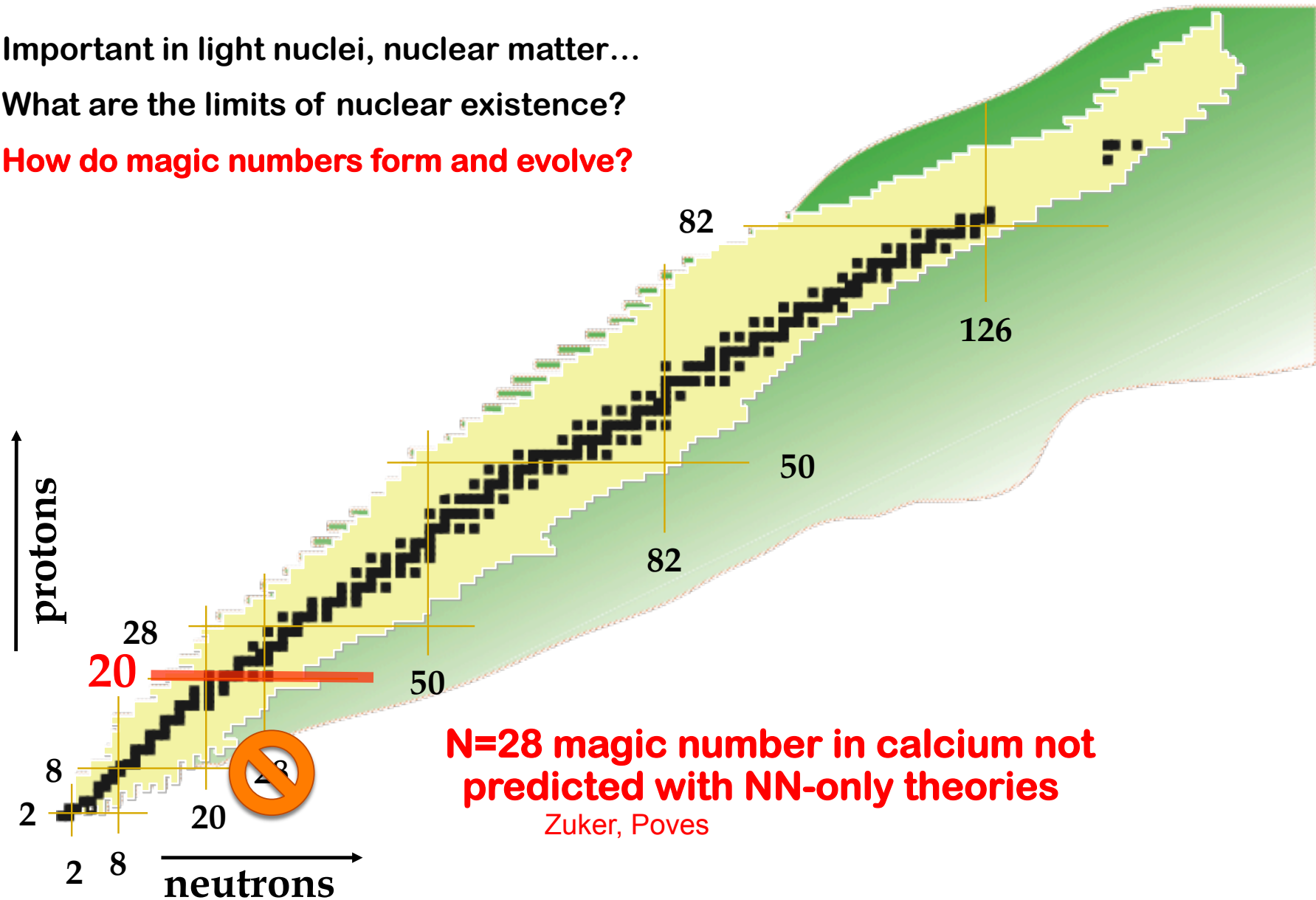
Need proper treatment of non-degenerate spaces

Shell Formation/Evolution in Calcium Isotopes

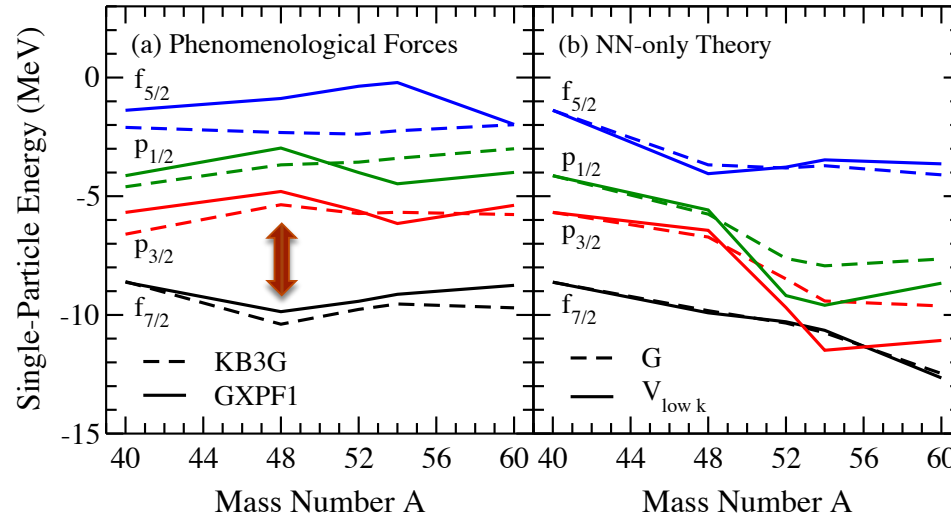
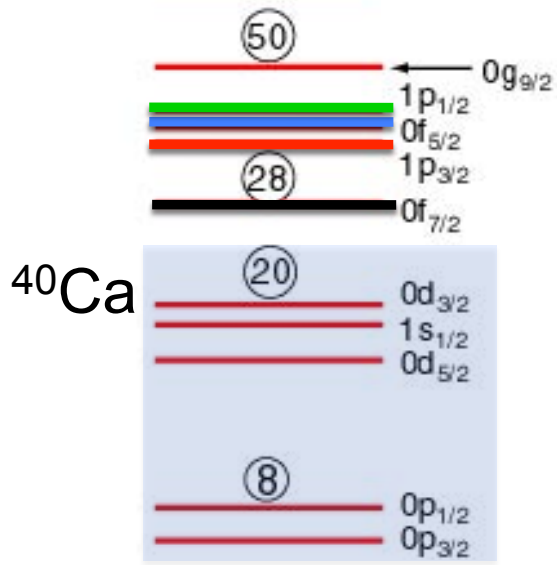
Important in light nuclei, nuclear matter...

What are the limits of nuclear existence?

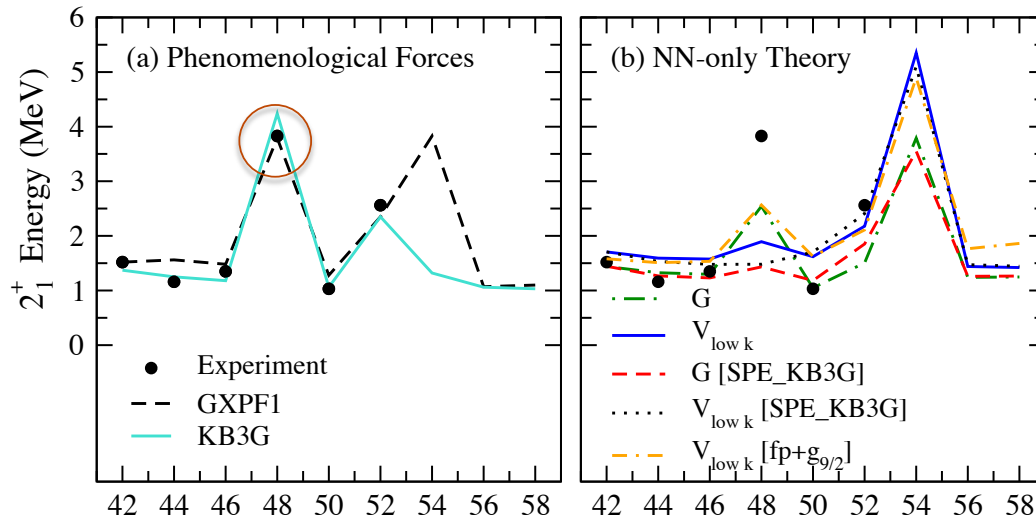
How do magic numbers form and evolve?



Calcium Isotope Physics: Magic Numbers



GXPF1: Honma, Otsuka, Brown, Mizusaki (2004)
 KB3G: Poves, Sanchez-Solano, Caurier, Nowacki (2001)



Phenomenological Forces

Large gap at ^{48}Ca
 Discrepancy at $N=34$

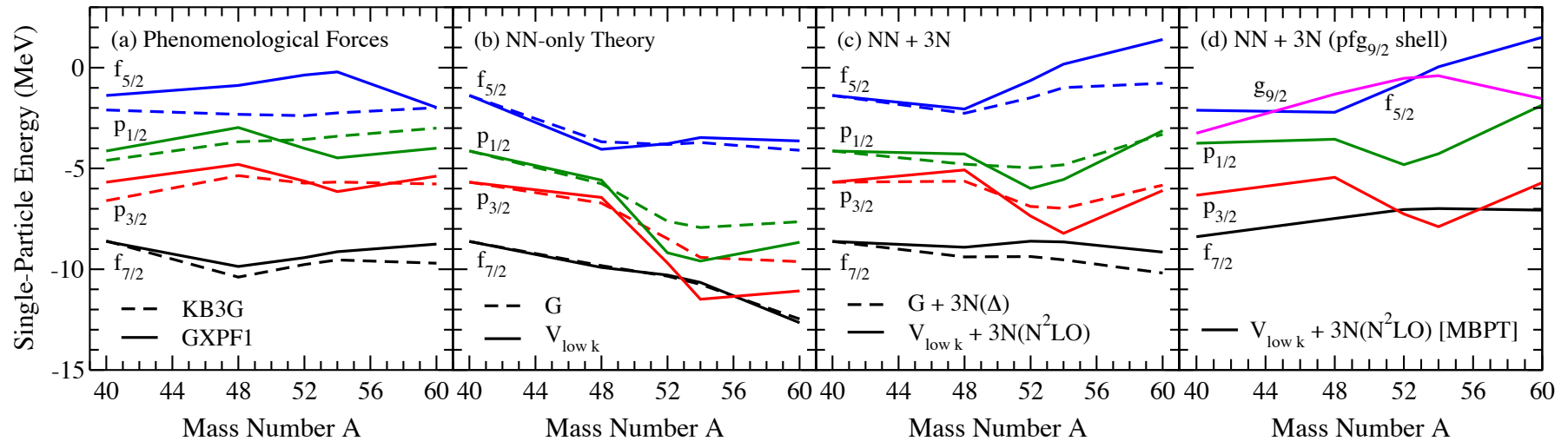
Microscopic NN Theory

Small gap at ^{48}Ca

$N=28$: first standard magic number not reproduced in microscopic NN theories

Evolution of Shell Structure

SPE evolution with 3N forces in pf and $pf g_{9/2}$ spaces:



Holt, Otsuka, Schwenk, Suzuki arXiv:1009.5984

NN+3N pf -shell:

Trend across: improved binding energies

Increased gap at ^{48}Ca : enhanced closed-shell features

Include $g_{9/2}$ orbit, calculated SPEs

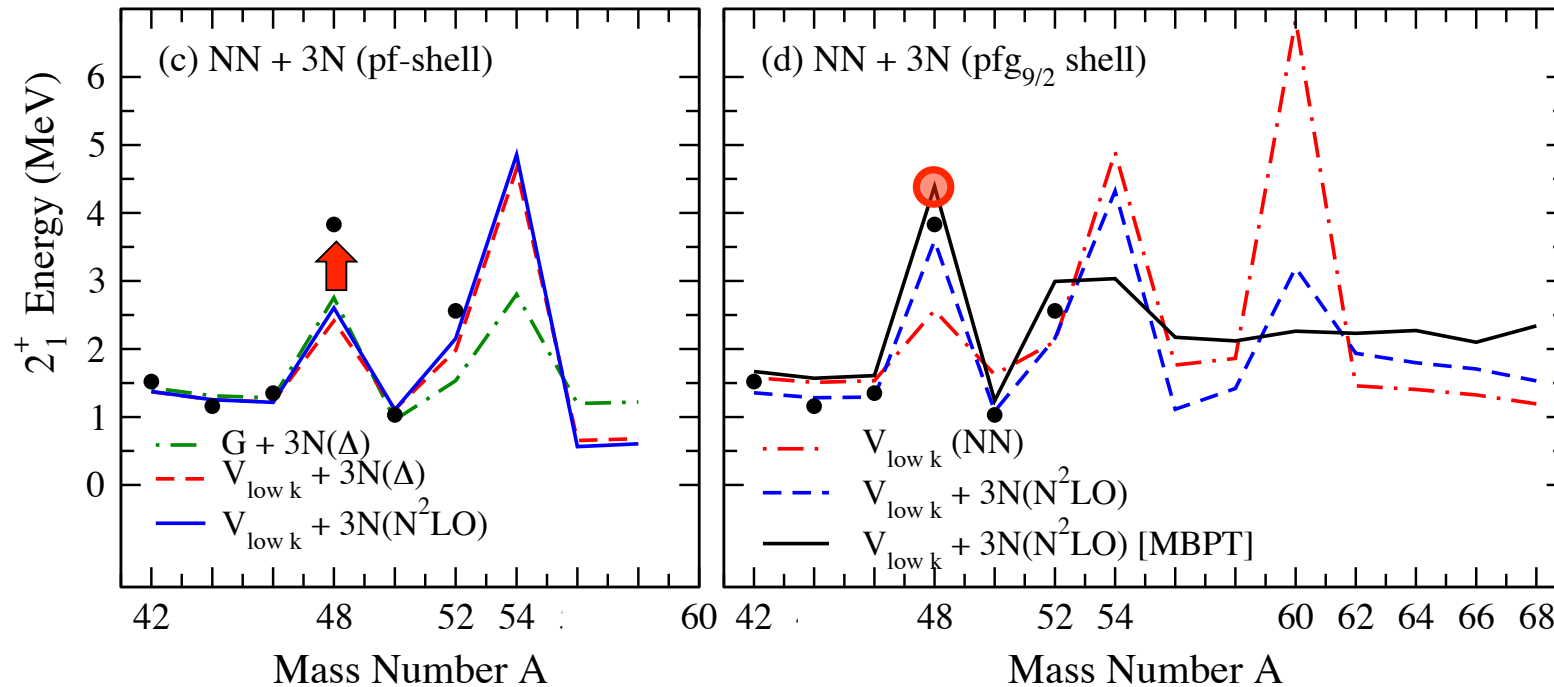
More modest N=34 gap

Different behavior of ESPEs

Small gap can give large 2^+ energy Duguet, Hagen, arXiv:1110.2468

N=28 Magic Number in Calcium

First excited 2^+ energies in calcium isotopes with NN+3N



Holt, Otsuka, Schwenk, Suzuki arXiv:1009.5984

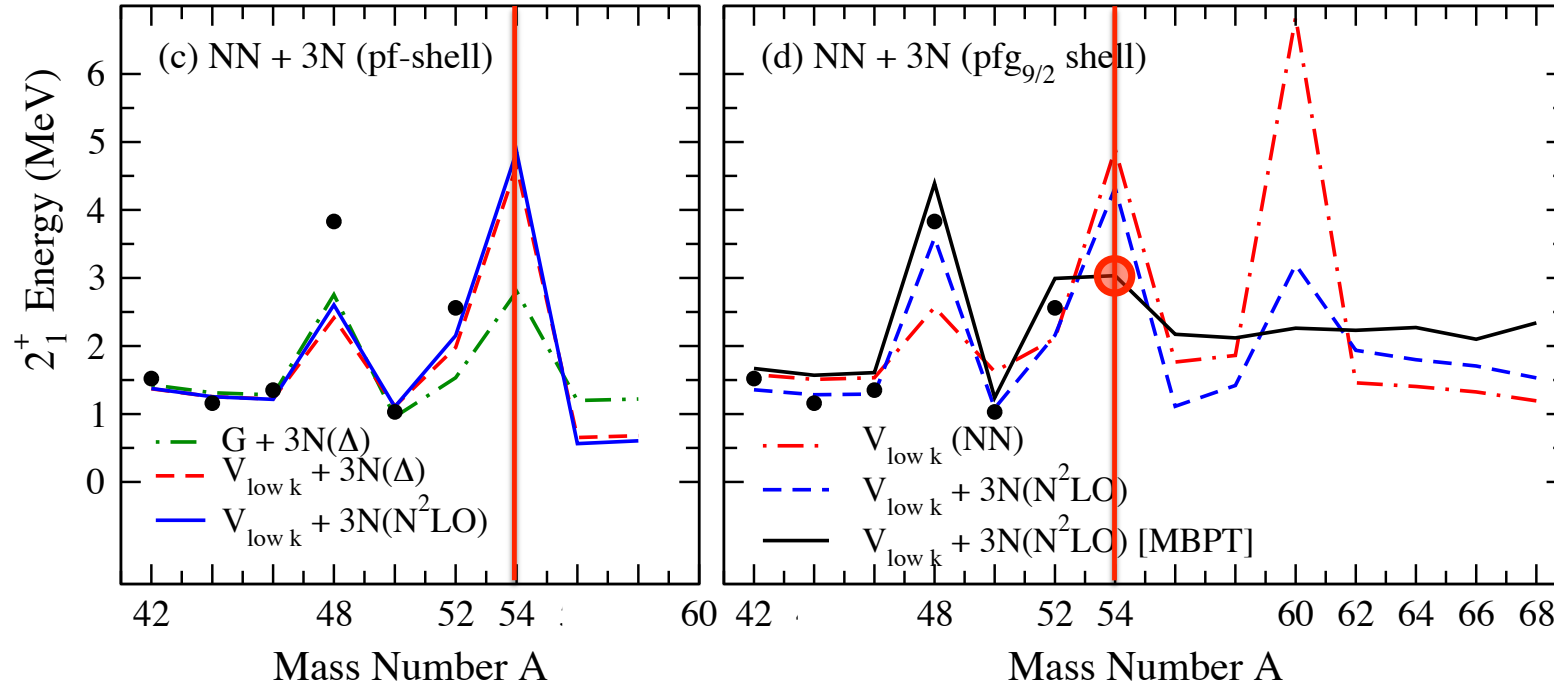
pf-shell: robust but modest improvement in 2^+ energies, below experiment

pfg_{9/2}-shell: reproduce experimental 2^+ in ^{48}Ca

Both 3N and extended space essential

Evolution of Magic Numbers: N=34

First excited 2^+ energies in calcium isotopes with NN+3N



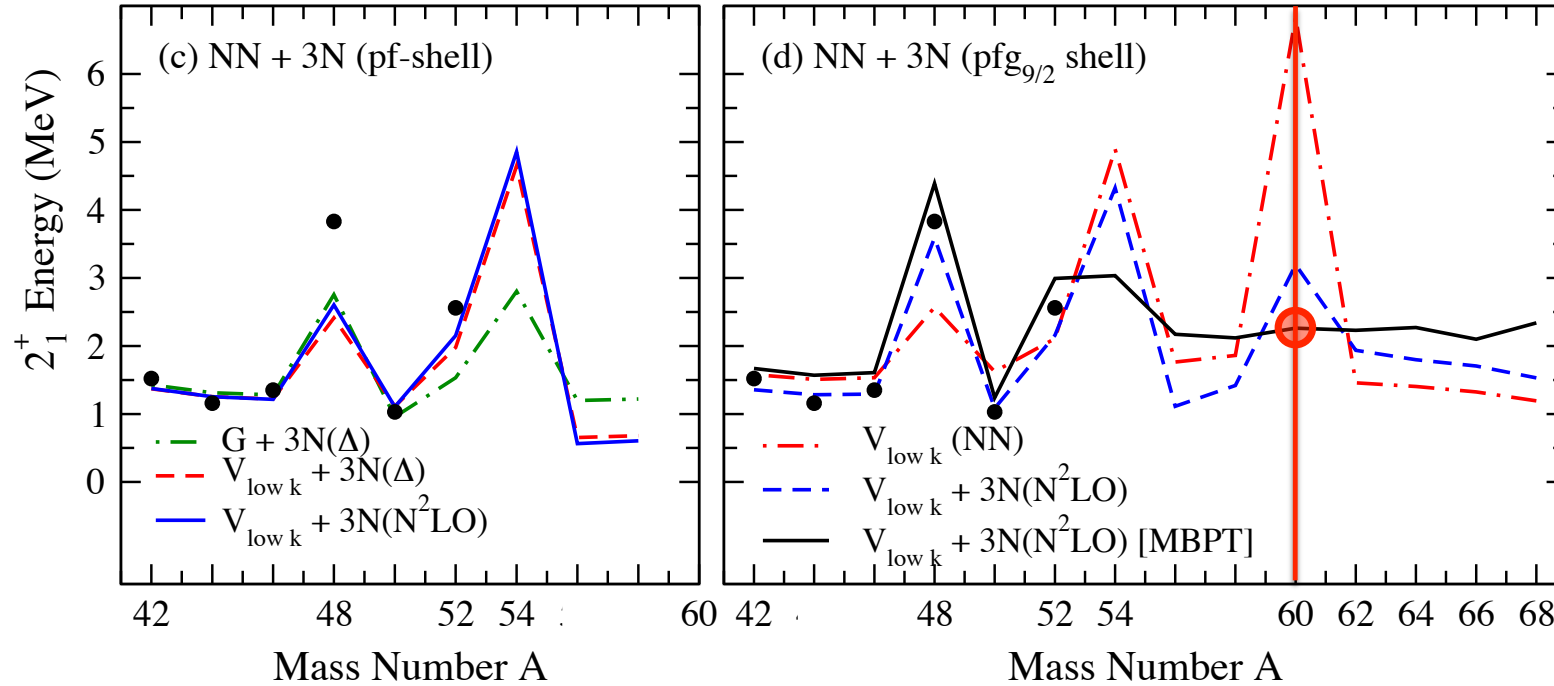
Holt, Otsuka, Schwenk, Suzuki arXiv:1009.5984

pf-shell: Very pronounced closed-shell properties

pfg_{9/2}-shell: More modest, similar to ^{52}Ca

Evolution of Magic Numbers: N=40

First excited 2^+ energies in calcium isotopes with NN+3N



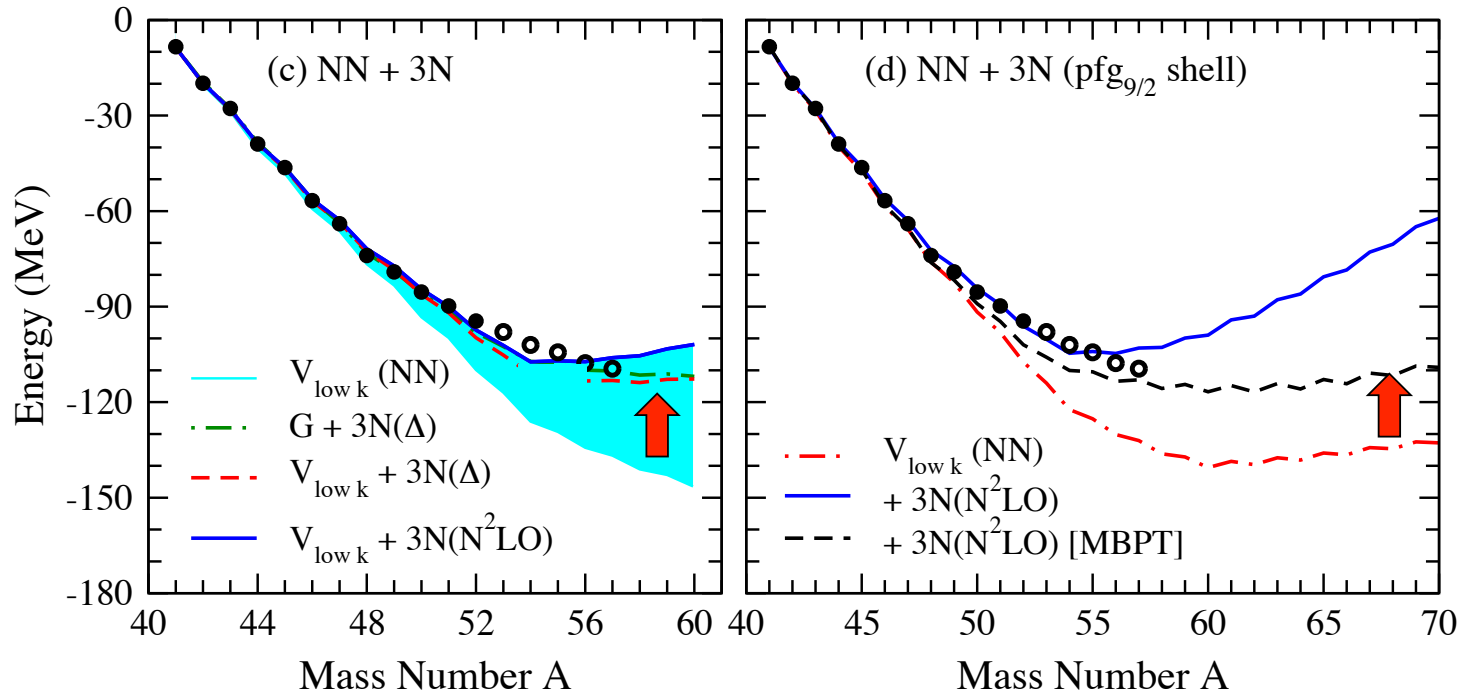
Holt, Otsuka, Schwenk, Suzuki arXiv:1009.5984

Extremely pronounced with NN-only
Dramatically quenched with NN+3N

Calcium Ground State Energies and Dripline

Ground state energies using NN+3N

NN-only: overbinds beyond $\sim {}^{46}\text{Ca}$



Holt, Otsuka, Schwenk, Suzuki arXiv:1009.5984

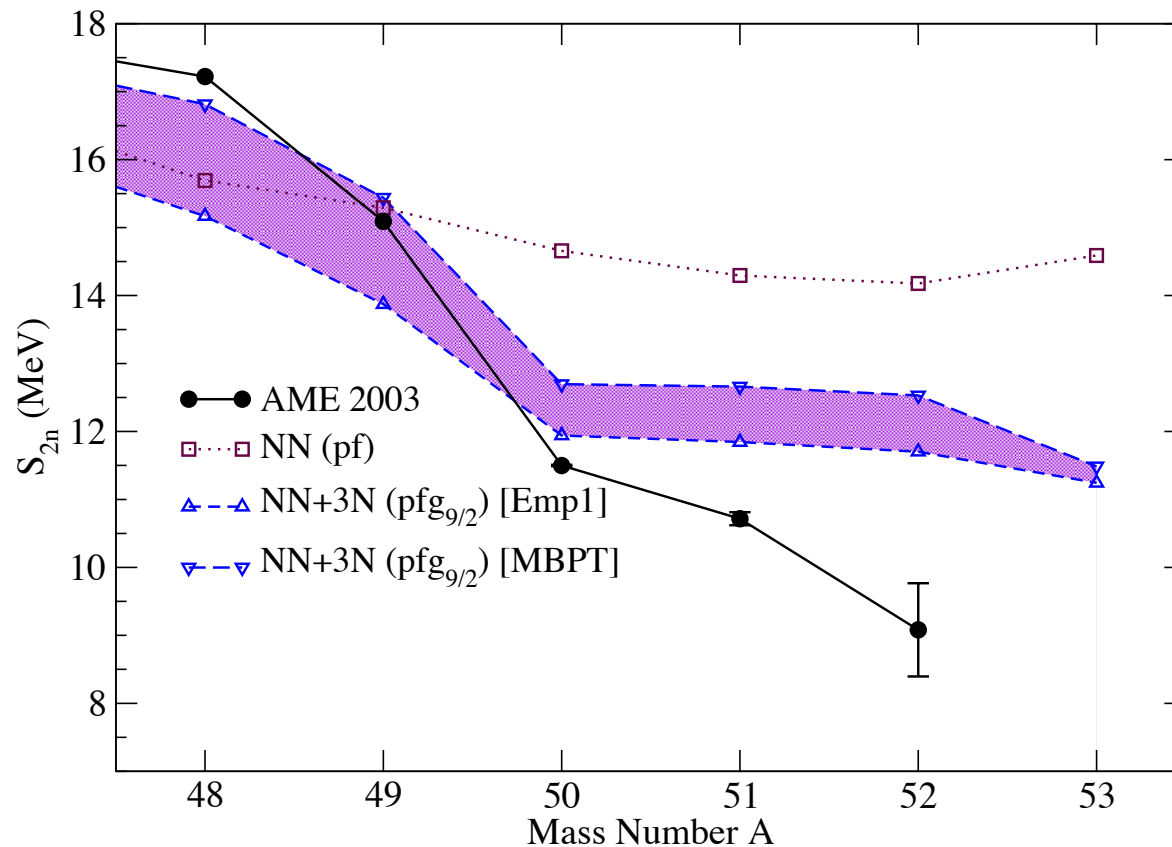
pf-shell: 3N forces correct binding energies; good experimental agreement

pfg_{9/2}-shell: calculate to ${}^{70}\text{Ca}$; modest overbinding (as in oxygen)

Predict heaviest calcium isotope $\sim {}^{58-60}\text{Ca}$

Experimental Connection: Mass of ^{52}Ca

S_{2n} energies for exotic calcium isotopes:

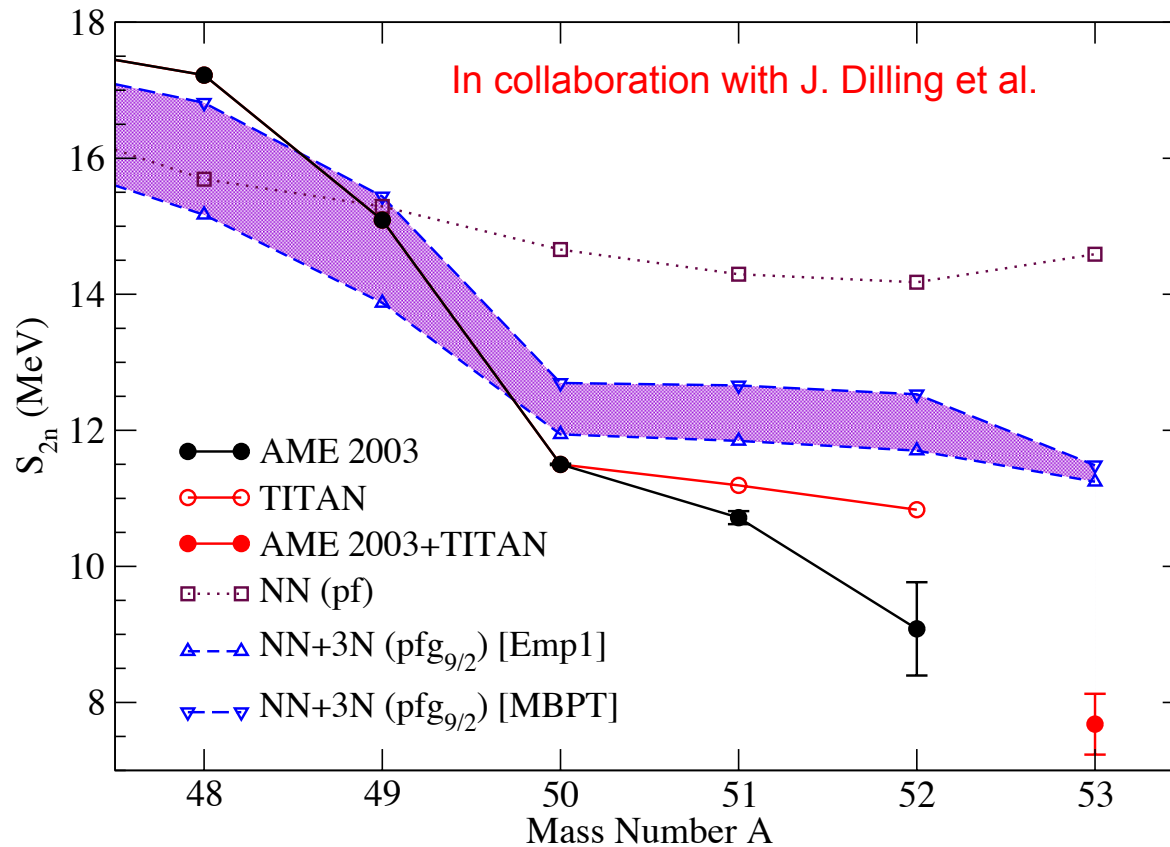


NN-only: poor agreement with experiment

NN+3N: reasonable improvement for lighter calcium, wrong behavior past ^{50}Ca

Experimental Connection: Mass of ^{52}Ca

New mass measurements of $^{51,52}\text{Ca}$ at **TITAN**: Penning trap experiment at TRIUMF



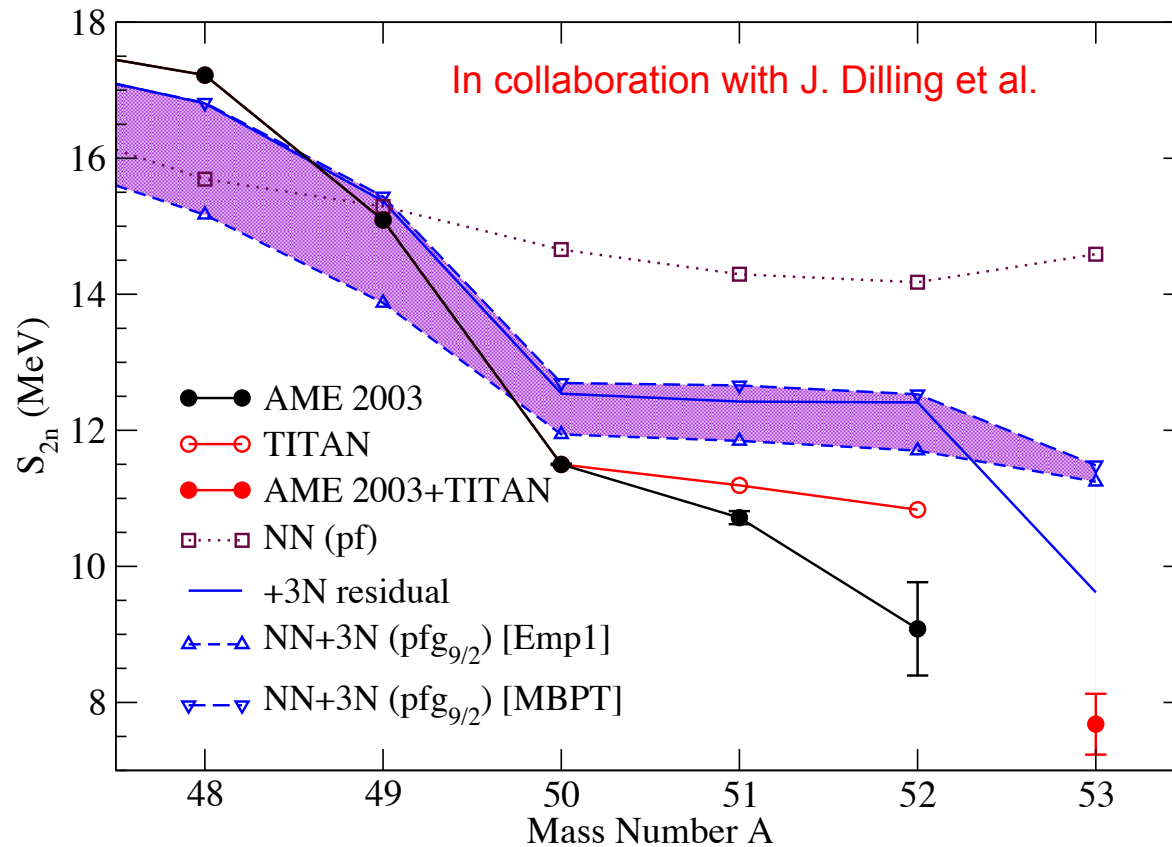
TITAN Measurement
 ^{52}Ca mass 1.75MeV
more bound than
AME2003 value!

NN+3N: Now reasonable agreement with experiment from $^{50-52}\text{Ca}$

What about at ^{53}Ca ? No sharp downward trend...

Experimental Connection: Mass of ^{52}Ca

New mass measurements of $^{51,52}\text{Ca}$ at **TITAN**: Penning trap experiment at TRIUMF



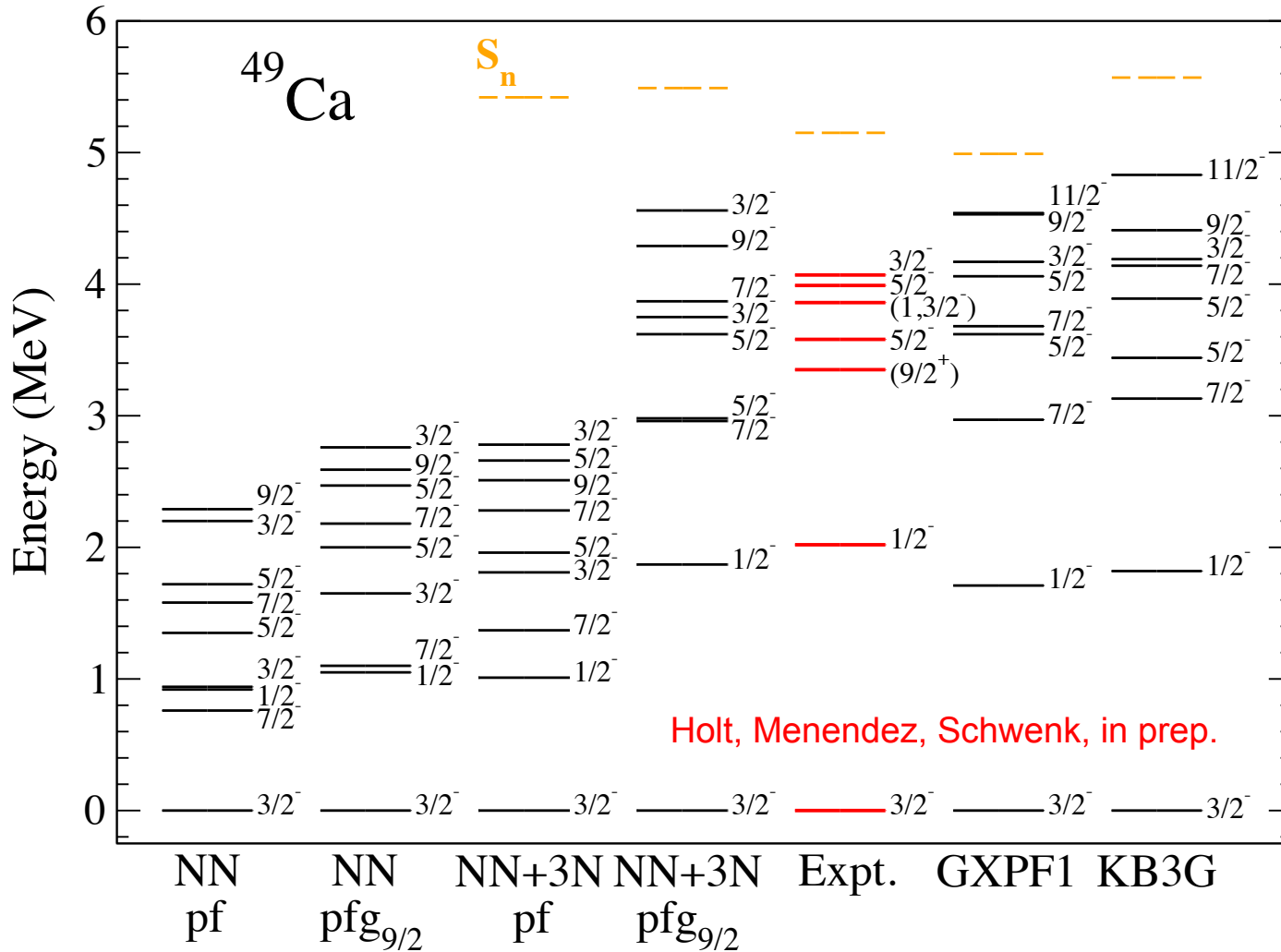
TITAN Measurement
 ^{52}Ca mass 1.75MeV
more bound than
 AME2003 value!

NN+3N: Now reasonable agreement with experiment from $^{50-52}\text{Ca}$

Adding approximated **residual 3N** important beyond ^{52}Ca : gives correct trend

Impact on Spectra: ^{49}Ca

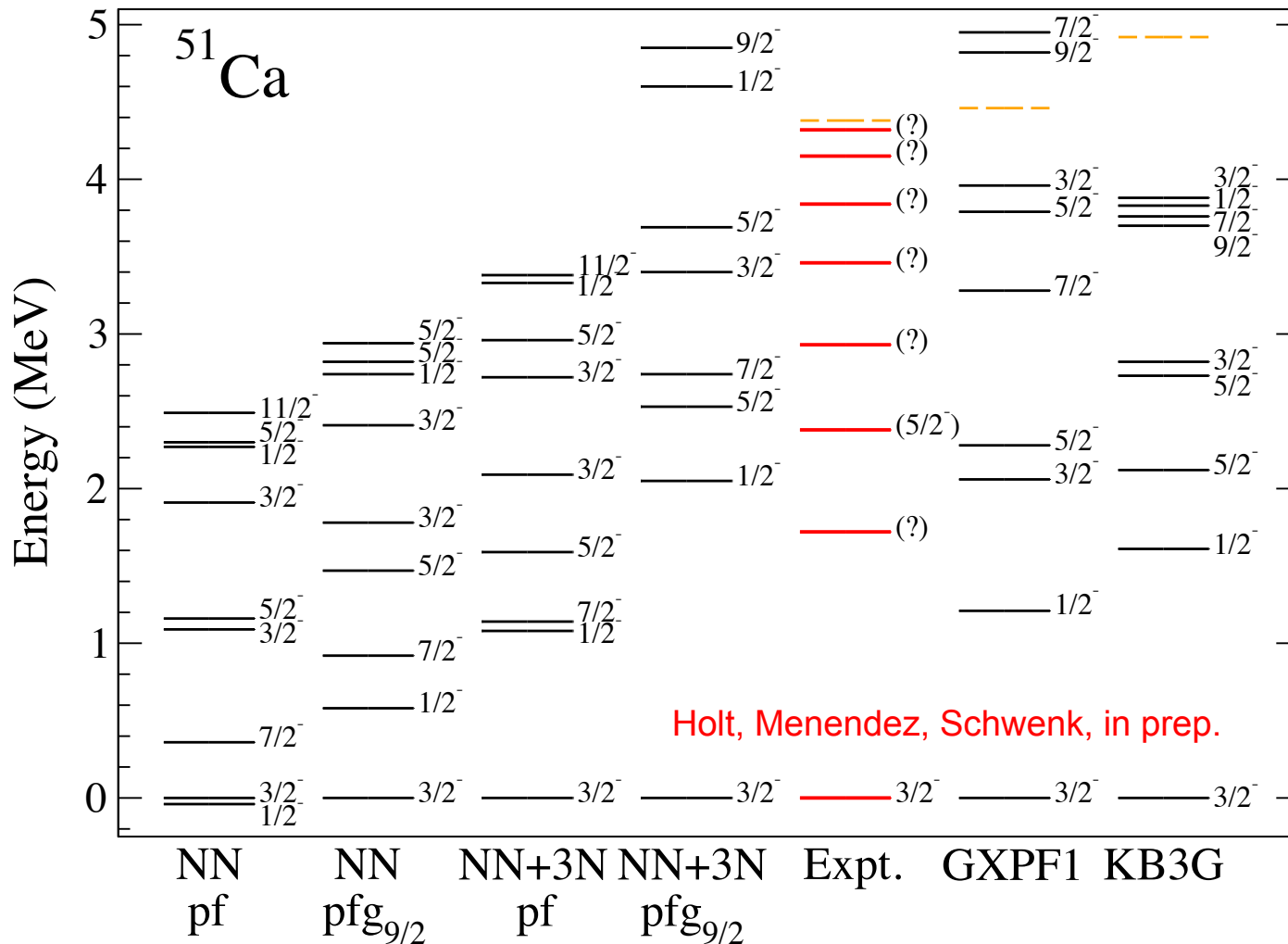
Neutron-rich calcium spectra with NN+3N



NN+3N in $pfg_{9/2}$ comparable to phenomenology (as for all lighter Ca isotopes)

Impact on Spectra: ^{51}Ca

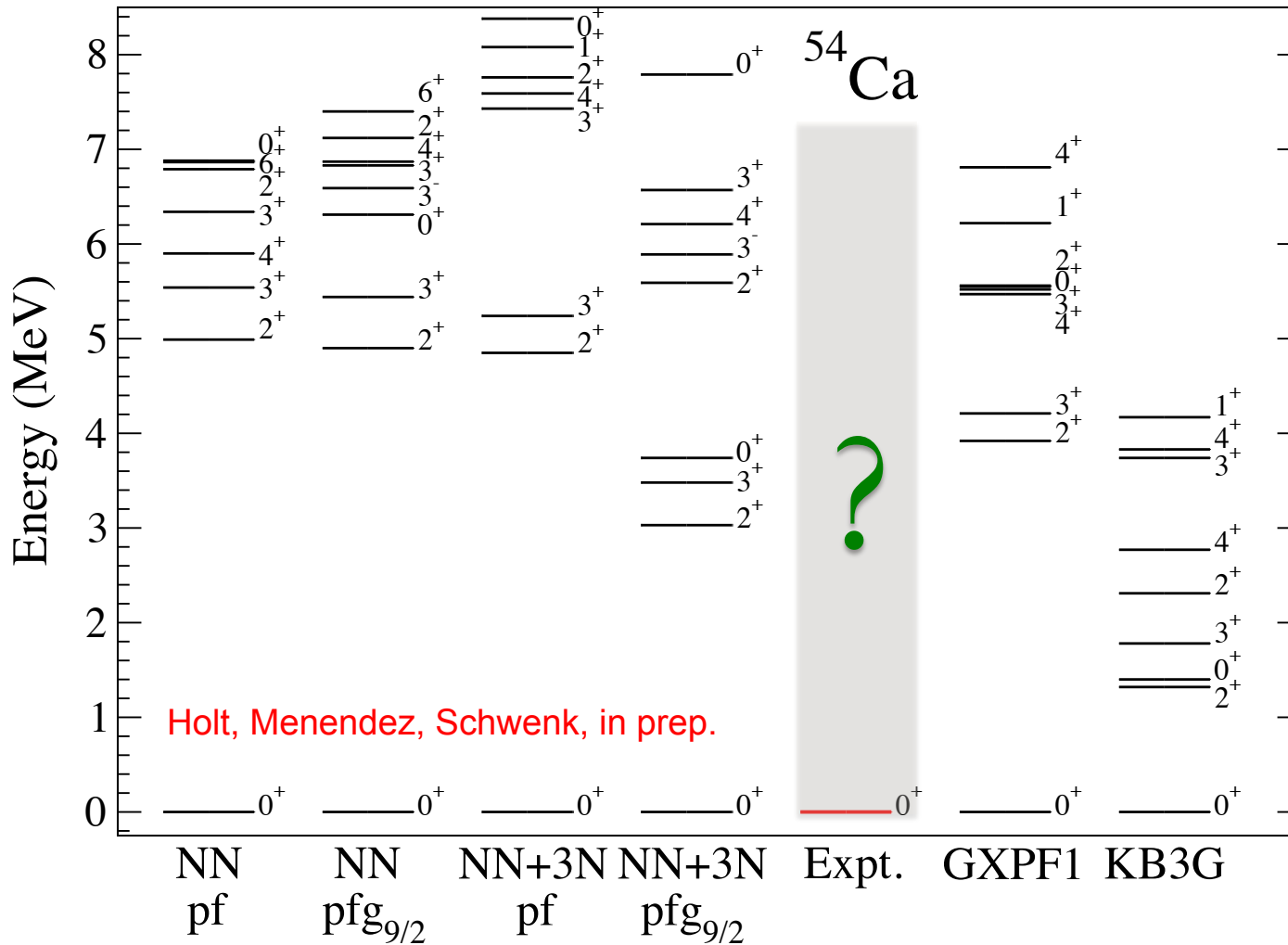
Neutron-rich calcium spectra with NN+3N



Different ordering from different models: gamma-ray spectroscopy needed!

Impact on Spectra: ^{54}Ca

Neutron-rich calcium spectra with NN+3N

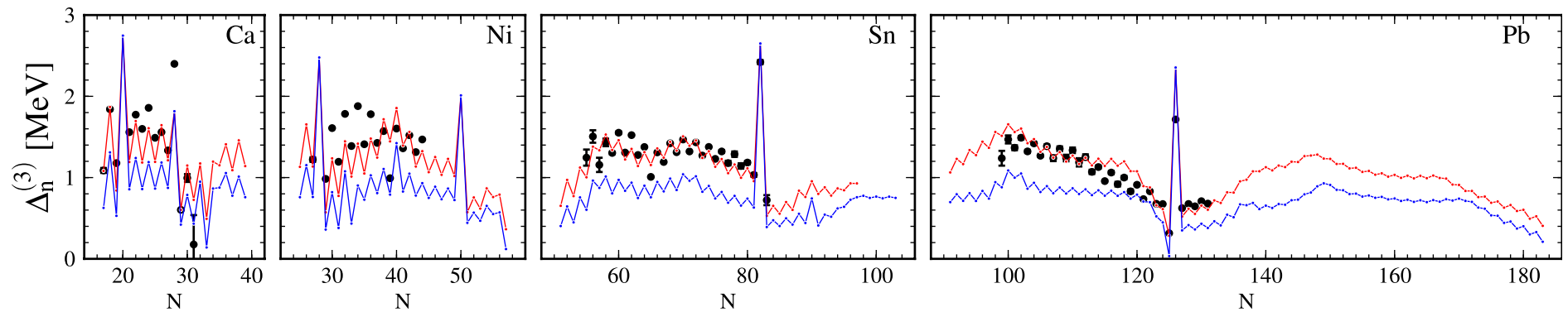


Microscopic prediction for ^{54}Ca spectrum

Pairing in EDF with 3N Forces

Calculations from Energy Density Functional with 3N forces

Lower gaps systematically:



Lesinski, Hebeler, Duguet, Schwenk, arXiv:1104.2955

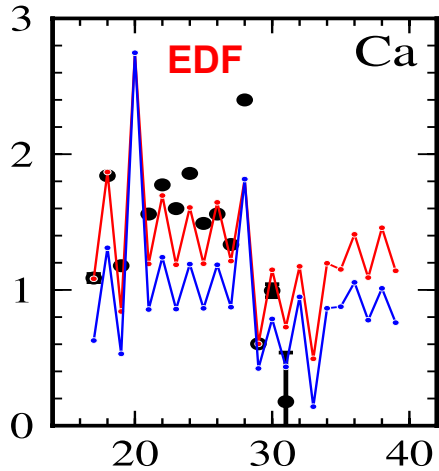
Improvement in theory – worse agreement with experiment

Due to neglected many-body effects?

Pairing in Calcium Isotopes

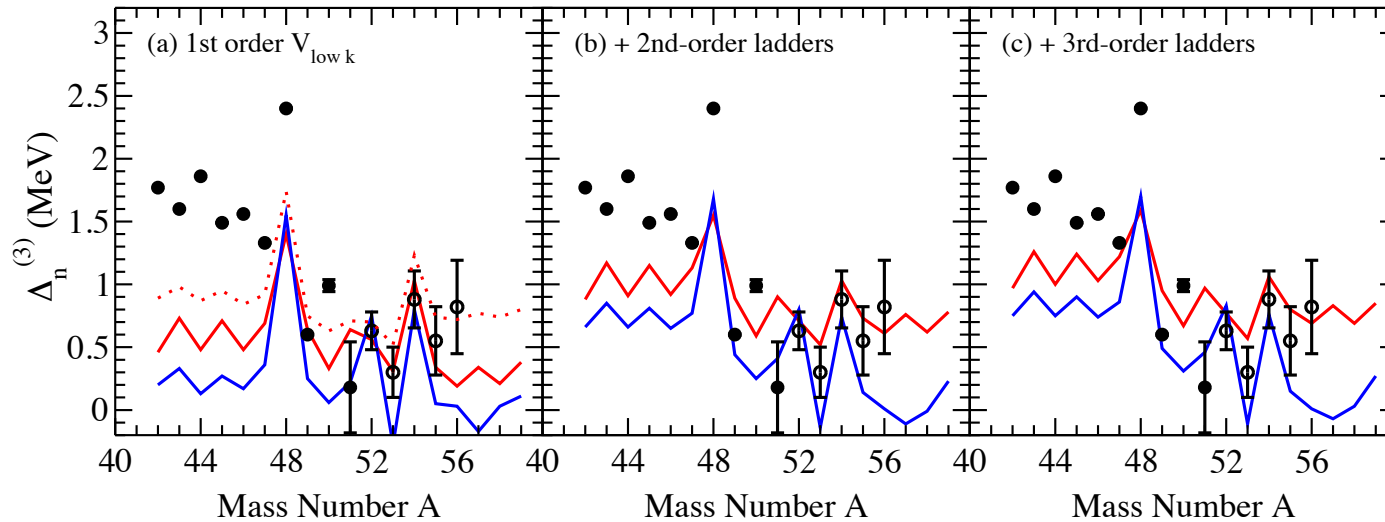
Compare with $\Delta_n^{(3)}$ calculated from microscopic NN+3N in calcium

Negative shift from 3N repulsion



Convergence in terms of order-by-order ladders

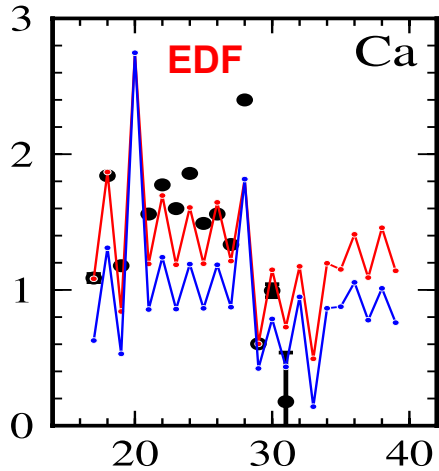
Holt, Menendez, Schwenk, in prep.



Pairing in Calcium Isotopes

Compare with $\Delta_n^{(3)}$ calculated from microscopic NN+3N in calcium

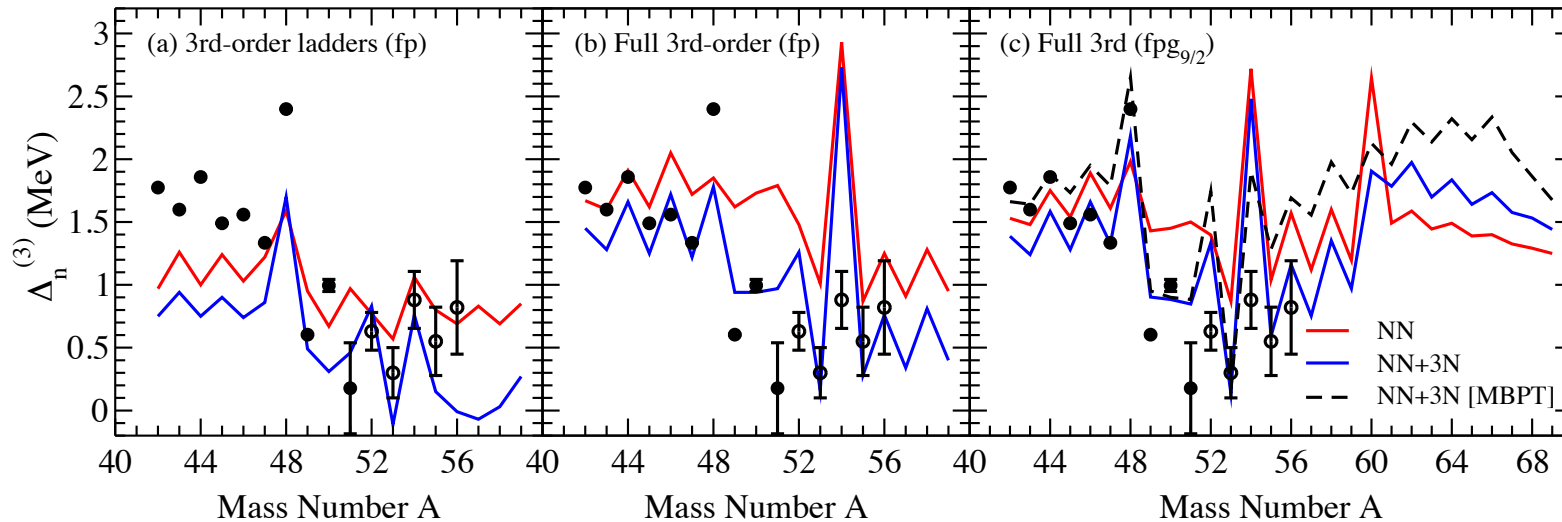
Negative shift from 3N repulsion



High order processes increase gaps - experimental agreement

Clear signature of shell closure in $^{48,54}\text{Ca}$, quenching in ^{60}Ca

Holt, Menendez, Schwenk, in prep.



Conclusion

- Exciting era for nuclear structure – **experimentally and theoretically**
- Exploring frontiers of nuclear structure theory of medium-mass nuclei with 3N forces
- Discovered robust and general repulsive 3N mechanism for $T=1$ neutron-rich nuclei

Conclusion

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- Discovered robust and general repulsive 3N mechanism for $T=1$ neutron-rich nuclei
- **Oxygen isotopes**: first fully-microscopic results with **NN+3N, extended spaces**
 - Cures NN-only failings: dripline, shell evolution, spectra in oxygen isotopes
- **Calcium isotopes** in pf - and $pf g_{9/2}$ -shells:
 - First microscopic prediction of $N=28$ magic number in ^{48}Ca
 - Shell evolution towards the dripline: modest $N=34$ closure, quenching of $N=40$
 - Dripline near ^{60}Ca
 - Spectra comparable to phenomenological models
 - NN+3N with higher many-body processes describe pairing without adjustments

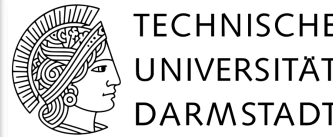
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 - NN+3N with higher many-body processes describe pairing without adjustments
- Clearly improvable upgrade path

Much to look forward to!

Acknowledgments

Collaborators



J. Menendez, A. Schwenk



T. Otsuka



T. Suzuki (Nihon U.)

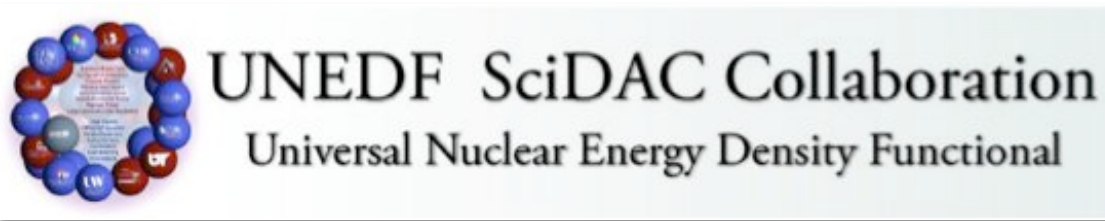


G. Hagen, T. Papenbrock



THE UNIVERSITY
of NORTH CAROLINA
at CHAPEL HILL

J. Engel



Computing support



Travel support