

PSI Dark Matter Homework #1

Due: Thursday, March 28, 2013

1. Dark Matter Profiles:

- a) Make plots (using a plotting program of your choice) of the NFW, Iso-Core, and Einasto ($\alpha = 1.7$) galactic DM profiles using the functional forms listed in the notes. Put them on the same set of (logarithmic) axes.
- b) We will see that the cosmic ray signal strength from given point in the sky from DM annihilation in our galaxy is proportional to $\langle\sigma v\rangle n_\chi^2$, where n_χ is the local DM number density. Assuming a constant cross section everywhere, make plots of the contributions to the signal strengths (as a function of the distance from the galactic center) for the NFW and Einasto profiles relative to the Iso-Core profile.

2. Boltzmannology:

- a) The collision term for the Boltzmann equation for n_χ we derived in class for the $\chi\chi \rightarrow f\bar{f}$ process was

$$- \int \frac{d^3p_1}{(2\pi)^3} \frac{d^3p_2}{(2\pi)^3} g_\chi^2 \sigma v (f_1 f_2 - f_{1eq} f_{2eq}), \quad (1)$$

where “1” and “2” refer to the two χ particles. Assuming $f(E, t) = \xi(t) f_{eq}(E)$, with $E = \sqrt{p^2 + m^2}$, re-express this quantity in terms of n_χ , $n_{\chi eq}$, and $\langle\sigma v\rangle$.

- b) Do the integral over p^0 in the Lorentz-invariant quantity $\int d^4p \delta(p^2 - m^2) \theta(p^0)$. What does this imply for how the quantity $(d\Pi) = d^3p / (2\pi)^3 2E$ changes under Lorentz transformations?
- c) Derive for yourself the rewriting of the freeze-out equation (Eq. (21) in **notes-2**) in terms of $Y_\chi = n_\chi / s$ as a function of $x = m/T$ (Eq. (25) of **notes-2**).

3. Cross Sections

- a) Consider the interaction

$$-\mathcal{L} \supset \lambda \phi \bar{f} \chi + (\text{h.c.}), \quad (2)$$

where χ is a DM fermion, f is a SM fermion of charge $Q = -1$, and ϕ is a complex scalar also of charge $Q = -1$. This interaction, together with its Hermitian conjugate, can mediate the annihilation $\chi\bar{\chi} \rightarrow f\bar{f}$. Compute the squared matrix element for this process, averaged over both initial and final spins. Write your answer in terms of dot products involving the two incoming momenta p_1 and p_2 , and the outgoing momenta p_3 and p_4 .

Hint: consult your notes from the tutorial.

- b) Evaluate this squared matrix element in the centre-of-mass frame, but do not assume that the incoming particles are non-relativistic. You may assume that the fermion is massless.

- c) Turn this matrix element into $4E_1E_2\sigma v$ by integrating over the final-state phase space ($\int d\Pi_3 d\Pi_4 (2\pi)^4 \delta^{(4)}(\dots)$) and multiplying by the appropriate factors.
- d) Assume the collision happens with both χ and $\bar{\chi}$ highly non-relativistic, and expand the result in powers of $u = p/E$, and keep terms in the expansion up to u^2 .
- e) Use this to compute the thermal average $\langle\sigma v\rangle$ following the approximate prescription described in class.