

# Future Long Baseline Neutrino Experiment in Japan

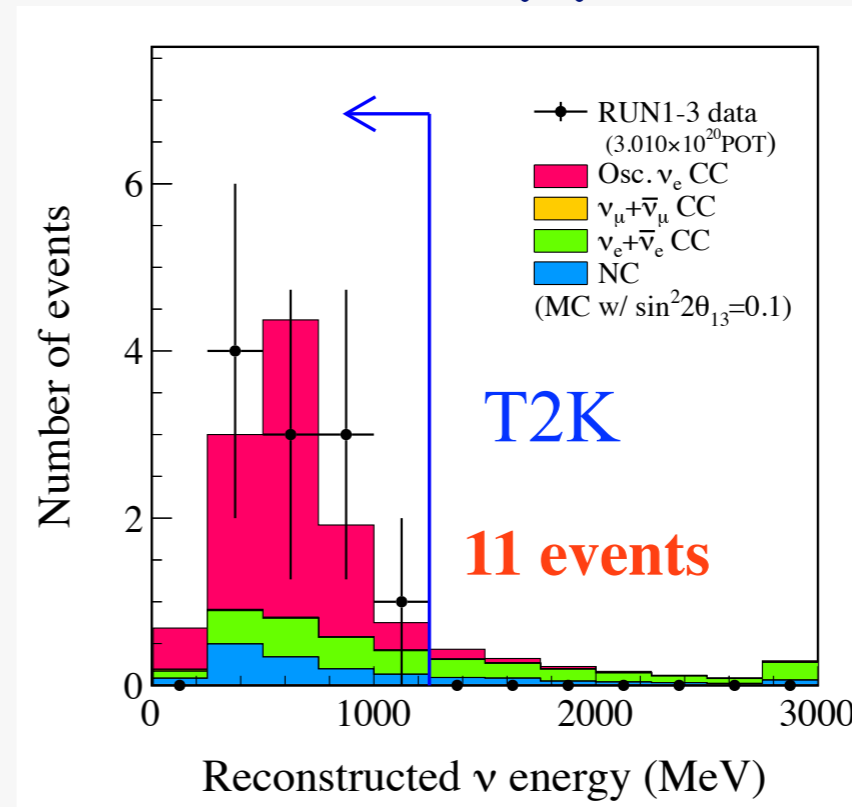
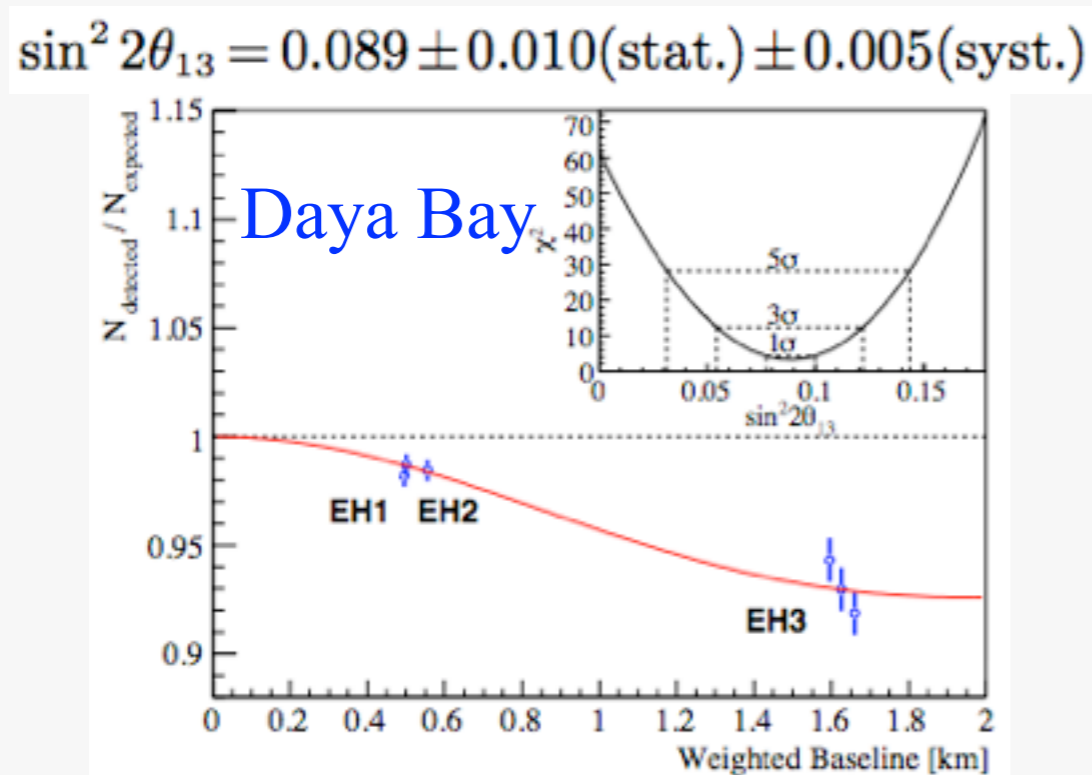
Akira Konaka  
(TRIUMF)

# Contents

- Introduction of T2K and HyperK experiments
- Motivation of future long baseline neutrino exp'ts
  - What sensitivities do we aim at?
- T2K/HyperK: Physics reaches
  - Accelerator based long baseline experiment
  - Atmospheric neutrinos (very long baseline)
  - Proton decay search
- Disclaimer
  - Official future T2K sensitivity is still in preparation
  - “Potential improvements” are my own thoughts

# $\theta_{13}$ is non-zero!

- 2012: non-zero  $\theta_{13}$  established
  - Reactor  $\bar{\nu}_e$  disappearance exp'ts:  $8\sigma$  discovery
  - T2K: 11 evts  $\rightarrow$   $3.2\sigma$  evidence in  $\nu_e$  appearance mode

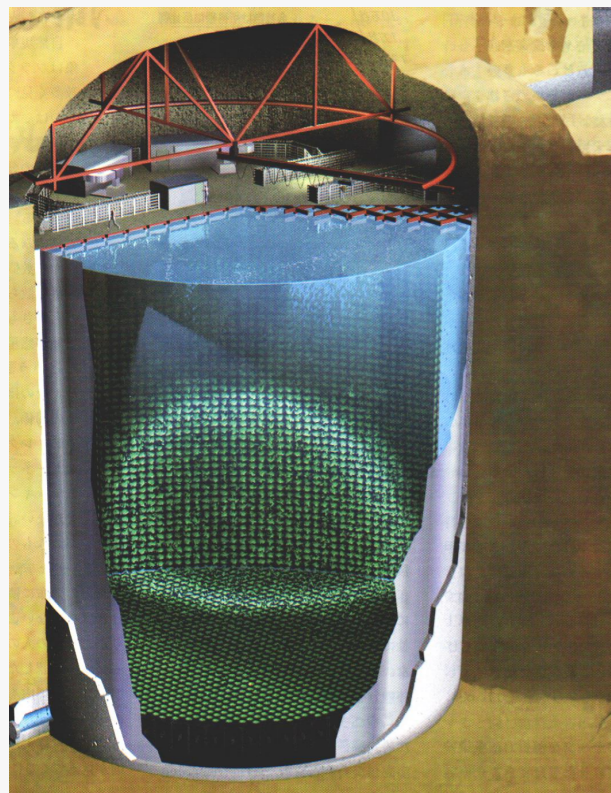


- CP measurement is feasible @ long baseline  $\nu$

$$A_{CP} = \frac{P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}{P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}$$

(reactor  $\nu$  is not sensitive to  $\delta_{CP}$ )

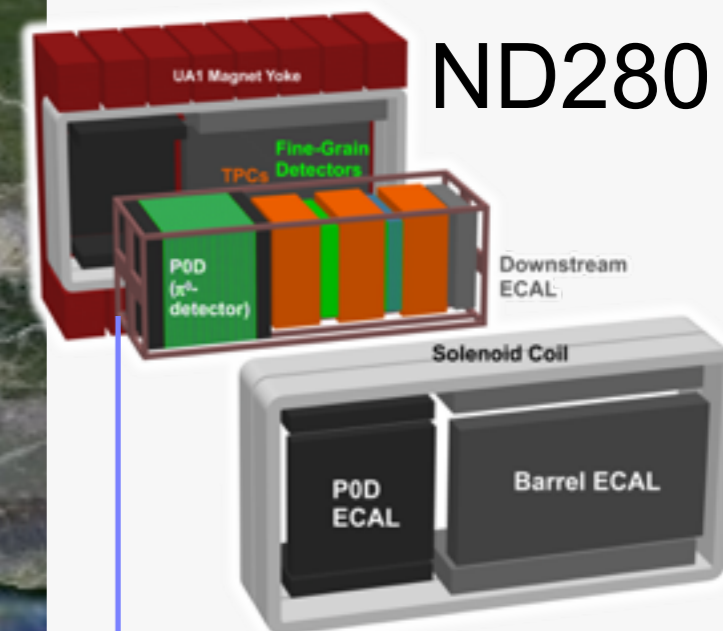
# Future of T2K and SK



Super-Kamiokande



295km



ND280

- Long baseline neutrino oscillation experiment from Tokai to Kamioka.
- $\nu_{\mu} \rightarrow \nu_e$  appearance to measure  $\theta_{13}$ , which leads to CP violation studies.



JPARC

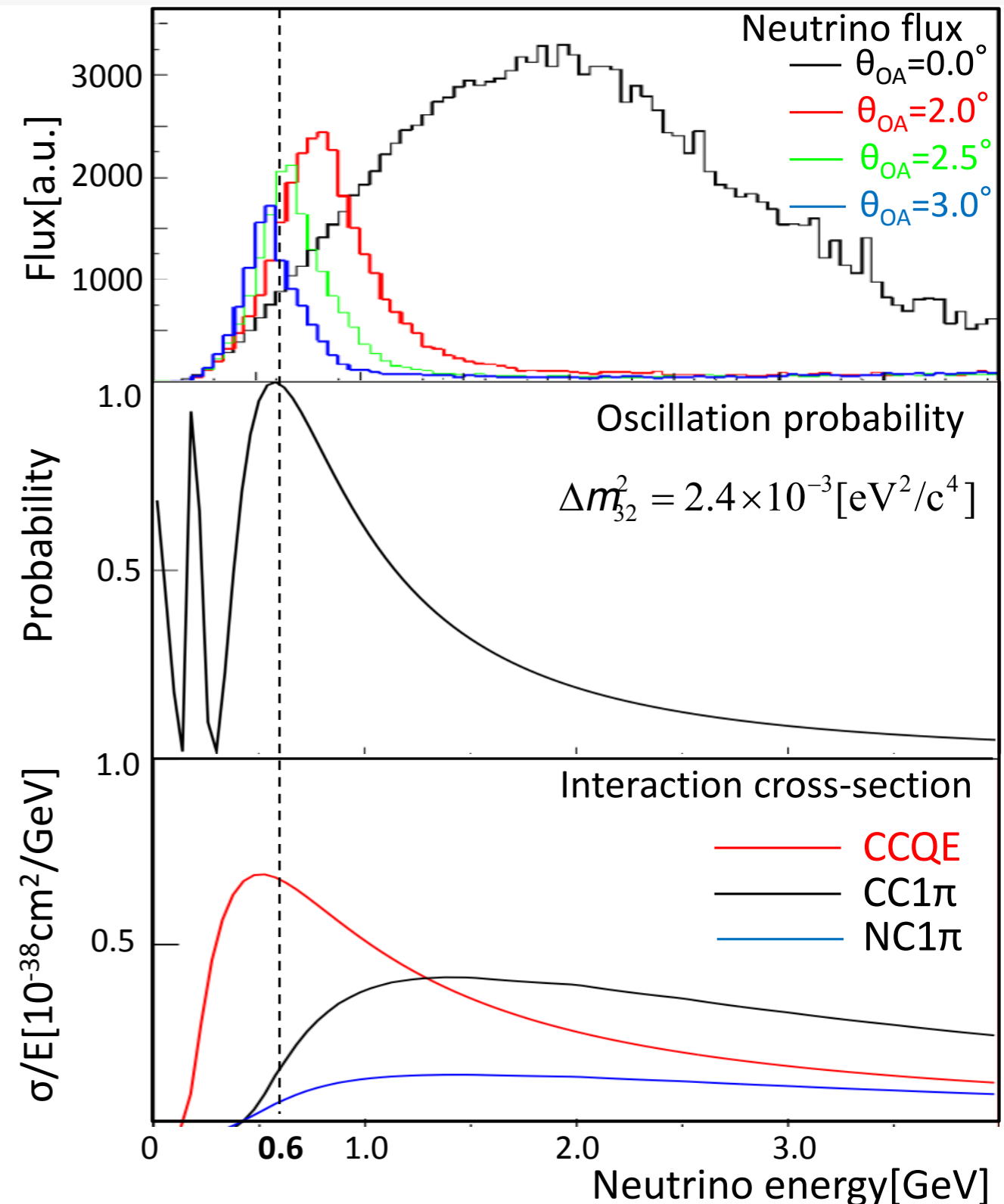
# Approach of the T2K experiment

- Narrow band beam tuned at the oscillation maximum
  - Off-axis  $\nu$  beam (2.5 deg.)
  - Maximize  $\nu$  oscillation

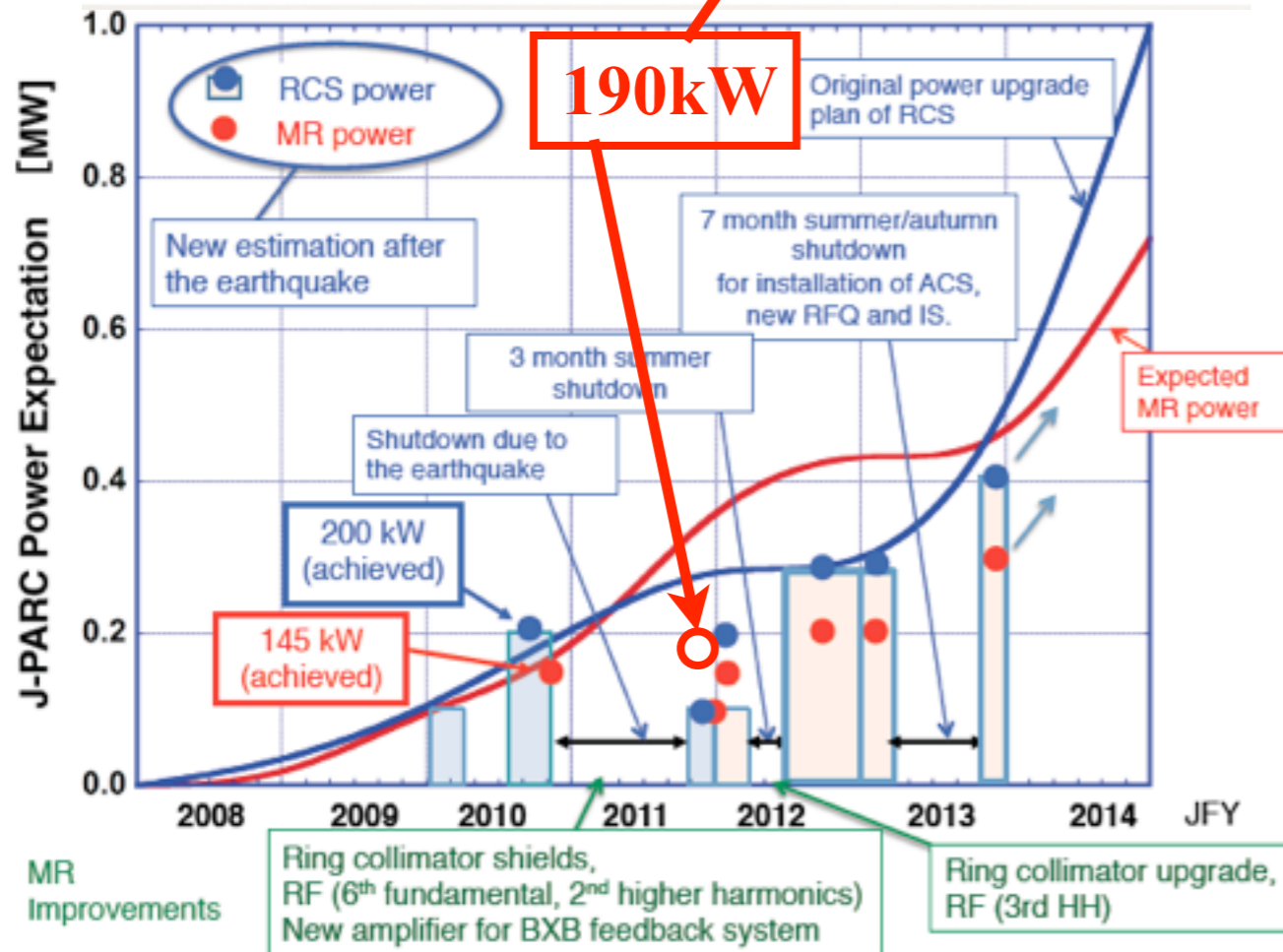
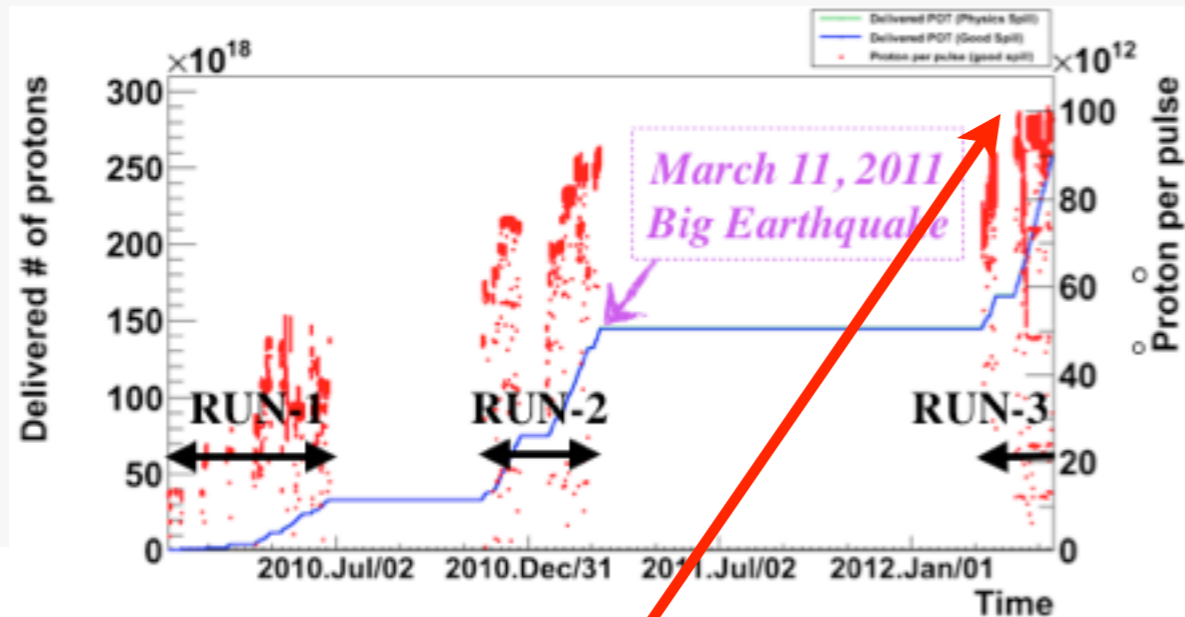
- Sub-GeV  $\nu$  beam (0.5-1GeV)
  - CCQE( $\nu_{\mu}n \rightarrow \mu p$ ) dominates
  - $E_{\nu}$  reconst. by  $\mu$  momentum

$$E_{\nu} = \frac{2E_{\ell}m_N - m_{\ell}^2}{2(m_N - E_{\ell} + P_{\ell}\cos\theta_{\ell})}$$

- Works well for water Cerenkov (Super-K)



# J-PARC beam status



- Earthquake Recovery
  - In less than a year
  - 190kW operation in 2012
    - 145kW before earthquake
    - exceed post-earthquake plan
- Plan towards full 750kW
  - Collimator upgrade
  - Linac upgrade
  - Main ring RF upgrade

# The medium-term plan of the MR-FX until 2017

We adopt the high repetition rate scheme to achieve the design beam intensity, 750 kW.  
Rep. rate will be increased from ~ 0.4 Hz to ~1 Hz by replacing magnet PS's and RF cavities.

JFY	2011	2012	2013	2014	2015	2016	2017
			Li. upgrade				
FX power [kW]	150	200	300	400			750
Cycle time of main magnet PS New magnet PS for high rep.	3.04 s	2.56 s	2.4 s				1.3 s
Present RF system New high gradient rf system	Install. #7,8	Install. #9					
Ring collimators	Additional shields	Add.collimators and shields (2kW)	Add.collimators (3.5kW)				
Injection system FX system	New injection kicker	Kicker PS improvement, Septum 2 manufacture /test					
		LF septum, PS for HF septa manufacture /test					

T.Koseki@HyperK meeting

# Potential MW upgrade of J-PARC

For the MR, scenarios for Multi-MW output beam power for neutrino experiment are being discussed .

T.Koseki@HyperK meeting

## 1. Large aperture MR

Enlarging the physical aperture from 81 to  $> 120 \pi \text{mm.mrad}$

A new synchrotron in the MR tunnel

## 2. Second booster ring for MR (emittance damping ring)

A ring with extraction energy  $\sim 8 \text{ GeV}$ , between the RCS and the MR

## 3. New rapid cycling synchrotron using 3 GeV RCS beam

- as a proton driver for neutrino beam production
- as an injector of MR, which is exclusively operated for the SX users

## 4. New rapid cycling synchrotron using the upgraded $\sim 1 \text{ GeV}$ linac beam

- as a proton driver for neutrino beam production
- as an injector of MR, which is exclusively operated for the SX users

....

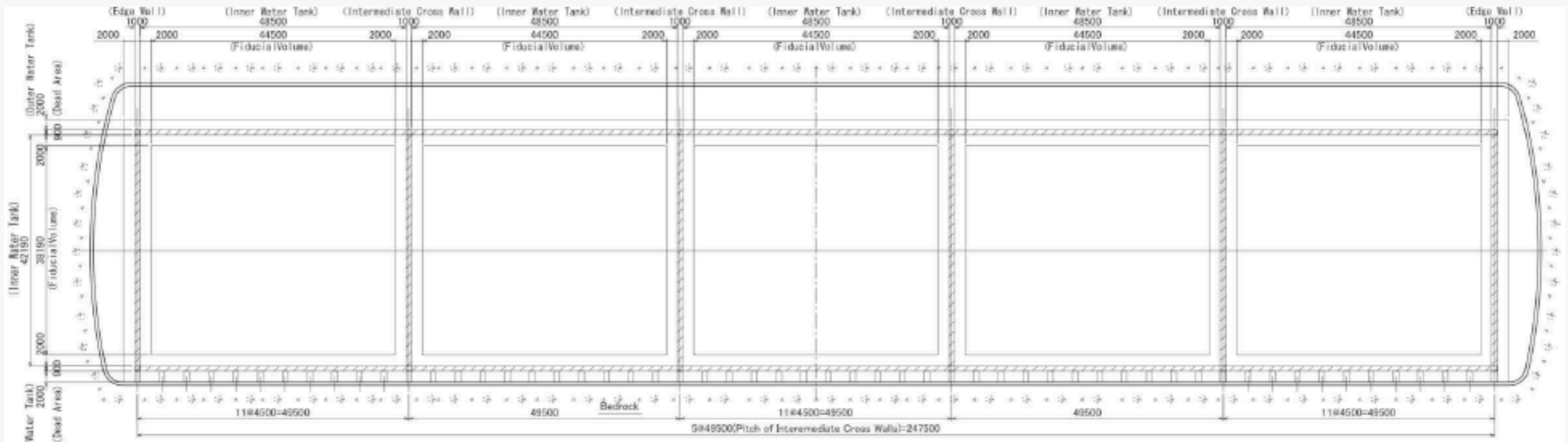
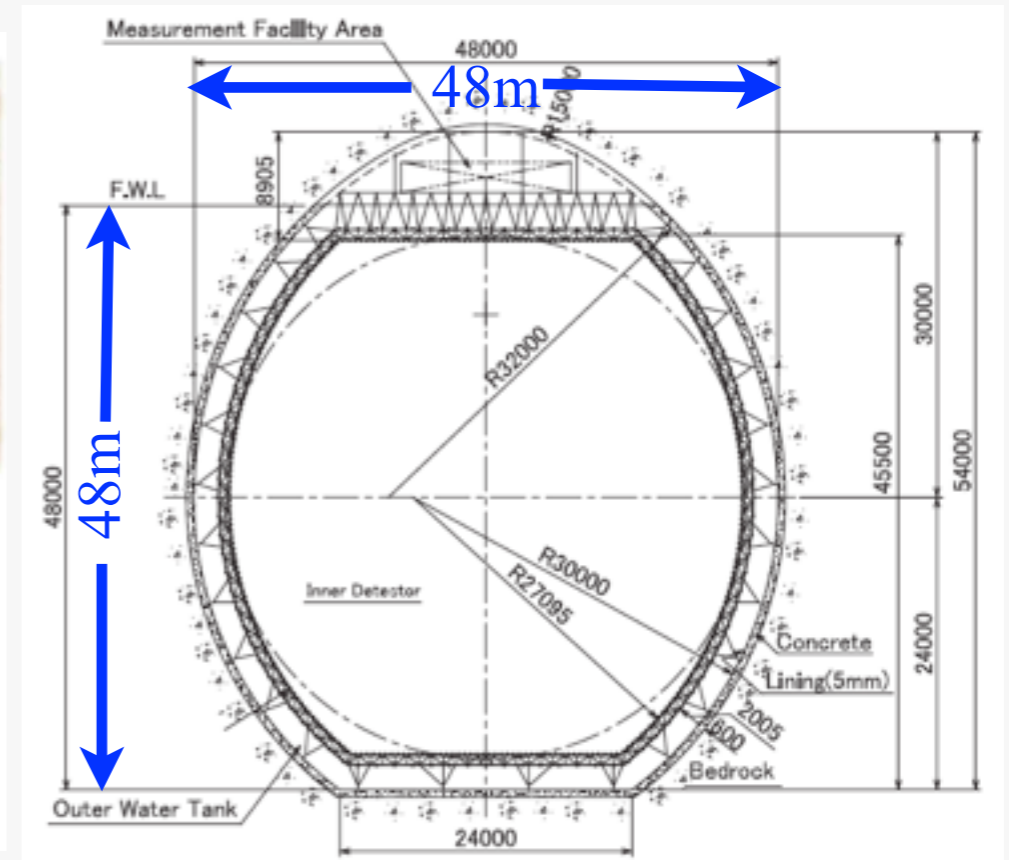
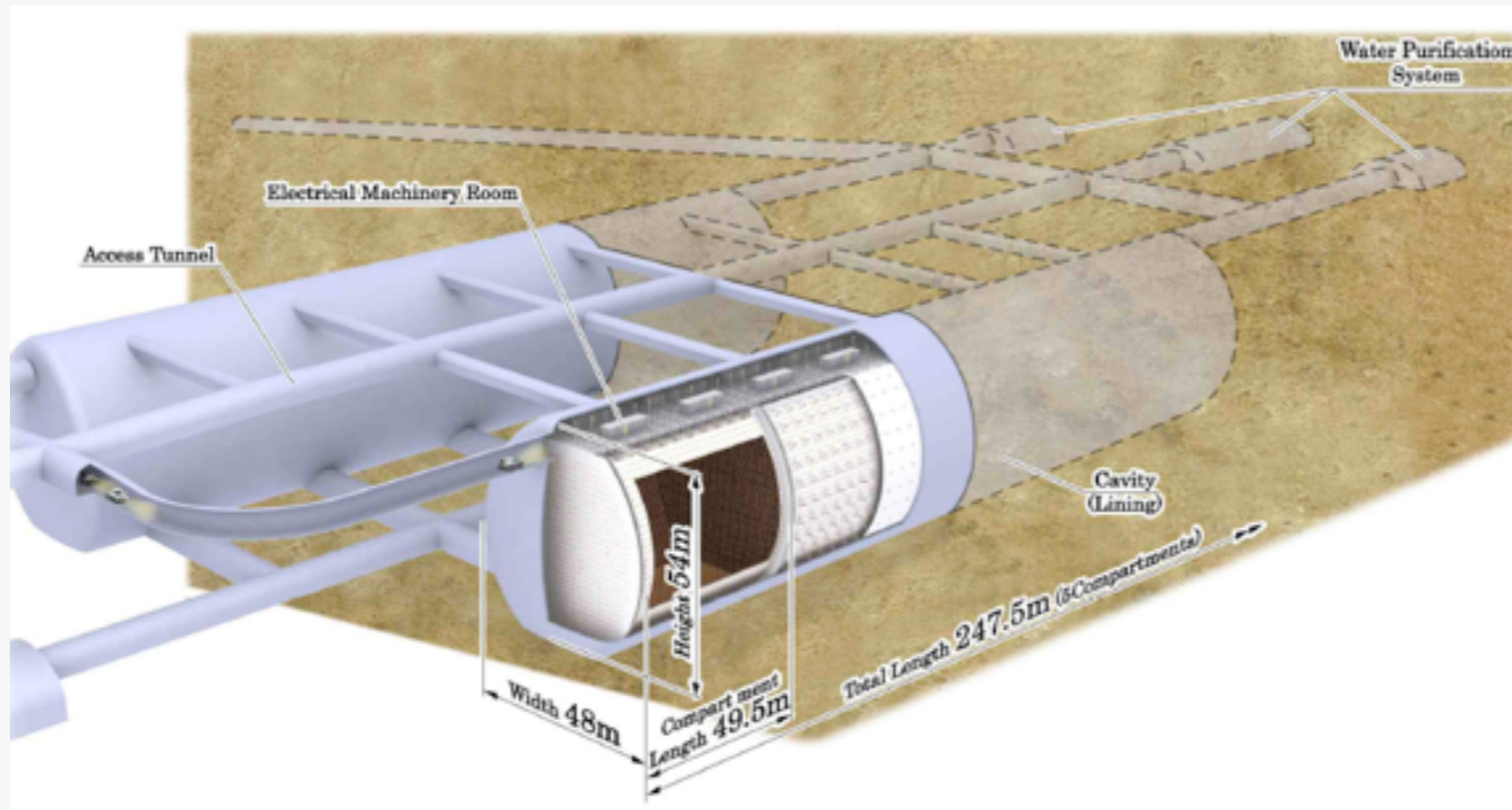
The detailed scheme of the future MW proton driver is discussed in the next five years and prepare to submit budget proposal to the government in 2018 or later.



# J-PARC upgrade status

- Effort to achieve 750kW is on going
  - R&D to double the repetition rate has started:
    - high gradient RF
    - magnet power supply for high repetition rate
  - R&D to double the #protons/bunch to be done
- Scenario to go beyond 750kW beam power is being discussed, but no concrete scenario or timeline yet.

# Hyper-Kamiokande project



# Comparison with SuperK

	Hyper-K	Super-K
Total volume	990kton	50kton
inner volume	740kton	32kton
fiducial volume	560kton	22.5kton
PMT's (20-inch)	99,000	11,146
photocathode coverage	20%	40%
Overburden (water eq.)	1,750m	2,700m
Off-axis angle	2.5 degree	2.5 degree
Baseline	295km	295km

- 25 times larger fiducial volume than SuperK



# PMNS texture

- Large mixing angles:
  - Indicating different origin for leptons from quarks
  - Lead us to understand the origin of family? GUT?
- Textures
  - Anarchy: random
  - Bi-maximal:  $\theta_{12}=\theta_{13}=45^\circ$
  - Tri-bimaximal: fits the data reasonably well

$$\begin{aligned}\sin^2 \theta_{12} &= \frac{1}{3} \\ \sin^2 \theta_{23} &= \frac{1}{2} \\ \sin^2 \theta_{13} &= 0\end{aligned}$$

$$\begin{aligned}|\nu_3\rangle &= \frac{1}{\sqrt{2}}(-|\nu_\mu\rangle + |\nu_\tau\rangle) \\ |\nu_2\rangle &= \frac{1}{\sqrt{3}}(|\nu_e\rangle + |\nu_\mu\rangle + |\nu_\tau\rangle) \\ |\nu_1\rangle &= \frac{1}{\sqrt{6}}(2|\nu_e\rangle - |\nu_\mu\rangle - |\nu_\tau\rangle)\end{aligned}$$

# Follow the History of CKM?

- Cabibbo quark mixing:  $\lambda = \cos\theta_c = 0.224$
- GIM mechanism: unitarity  $\rightarrow$  no FCNC, predict charm
- Kobayashi-Maskawa:  $\delta_{CP}$ , predict 3 families (t, b)
- Precision studies (Wolfenstein parametrization)

$$V = \begin{pmatrix} 1 - \frac{1}{2}\lambda^2 & \lambda & A\lambda^3(\rho - i\eta) \\ -\lambda & 1 - \frac{1}{2}\lambda^2 & A\lambda^2 \\ A\lambda^3(1 - \rho - i\eta) & -A\lambda^2 & 1 \end{pmatrix}$$

CKM expanded by  $\lambda$

$$\lambda = 0.2243 \pm 0.0016$$

$$A = 0.82$$

$$\rho = 0.20, \eta = 0.33$$

- Tri-bimaximal in lepton mixing? Expand in  $\sin\theta_{13}$ ?
- Precision down to  $\sin^2\theta_{13} = 0.024$  ( $\sim 2\%$ ) level.
- Discovery of the underlying principle? GUT?

# Global fit (2012)

- Global fit results are getting very interesting:
  - consistent with tri-bimaximal but some hints
  - $\sin^2\theta_{13} > 0$ ,  $\sin^2\theta_{23} < 1/2?$ ,  $\sin^2\theta_{12} < 1/3?$ ,  $\delta_{cp}?$

parameter	best fit	1 $\sigma$ range	2 $\sigma$ range	3 $\sigma$ range	tri-bimaximal
$\sin^2\theta_{12}$ (NH/IH)	0.307	0.291-0.325	0.275-0.342	0.259-0.359	0.333
$\sin^2\theta_{13}$ (NH)	0.0241	0.0216-0.0266	0.0193-0.0290	0.0169-0.0313	0.0
$\sin^2\theta_{13}$ (IH)	0.0244	0.0219-0.0267	0.0194-0.0291	0.0171-0.0315	0.0
$\sin^2\theta_{23}$ (NH)	0.386	0.365-0.410	0.348-0.448	0.331-0.637	0.5
$\sin^2\theta_{23}$ (IH)	0.392	0.370-0.431	0.353-0.641	0.335-0.663	0.5
$\delta$ (NH)	1.08 $\pi$	(0.77-1.36) $\pi$	-	-	-
$\delta$ (IH)	1.09 $\pi$	(0.83-1.47) $\pi$	-	-	-

**Fogli et. al. ArXiv:1205.5254v3**

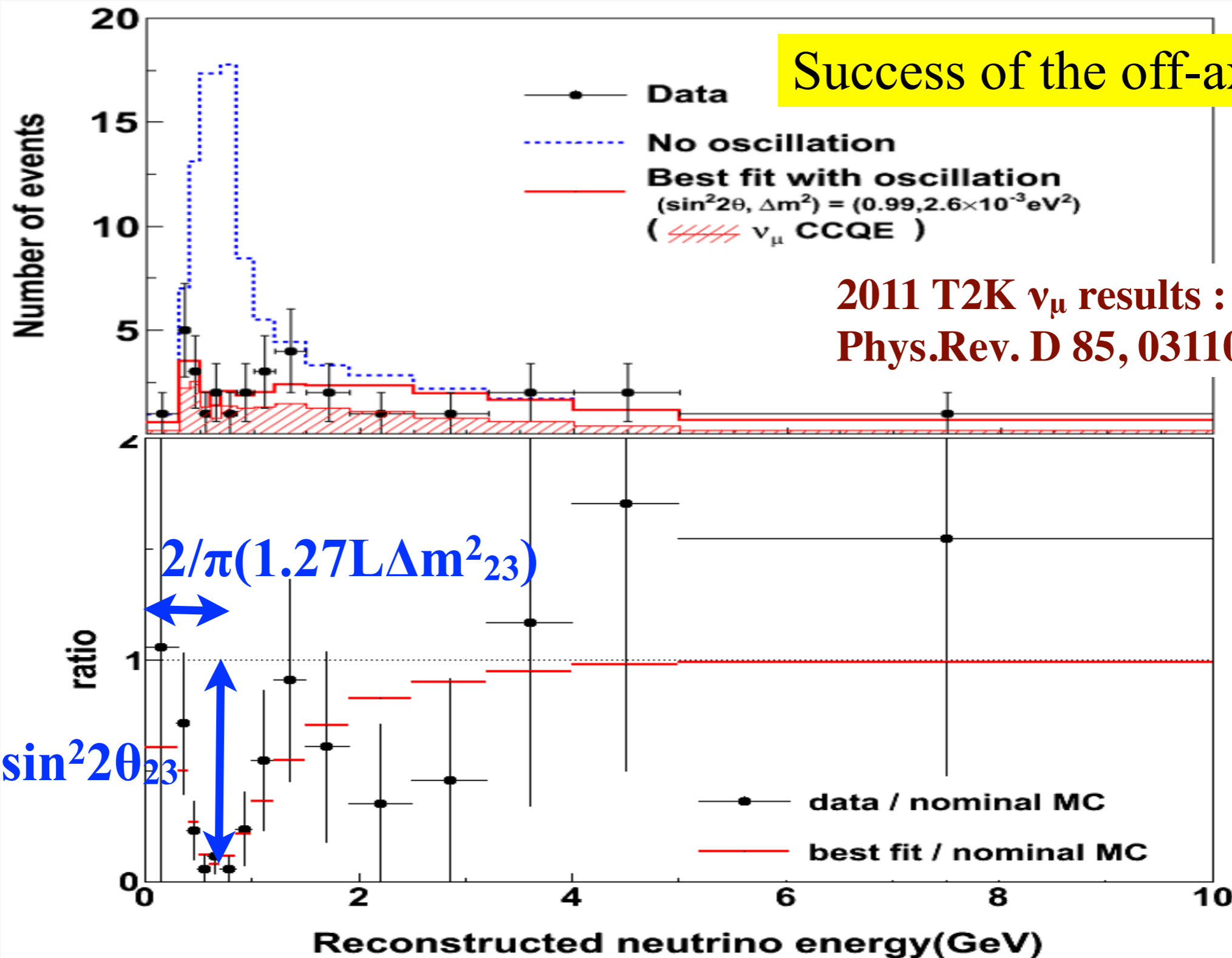
# Future precision measurements

- $\theta_{12}$ : solar neutrinos
  - $\theta_{12}$ : SNO (CC and NC solar),  $\Delta m^2_{12}$ : Kamland (reactor)
  - Future: Daya Bay2 (60km detector), pp/pep solar  $\nu$ 's
- $\theta_{13}$ : Reactor neutrinos
  - Improvements expected by Daya Bay and others.
- $\theta_{23}$ : **Long baseline neutrinos**
- $\delta_{cp}$ : **Long baseline neutrinos**
- Degeneracy: mass hierarchy,  $\theta_{23}$  octant
  - Very long baseline (accelerator, **atm. neutrinos**)



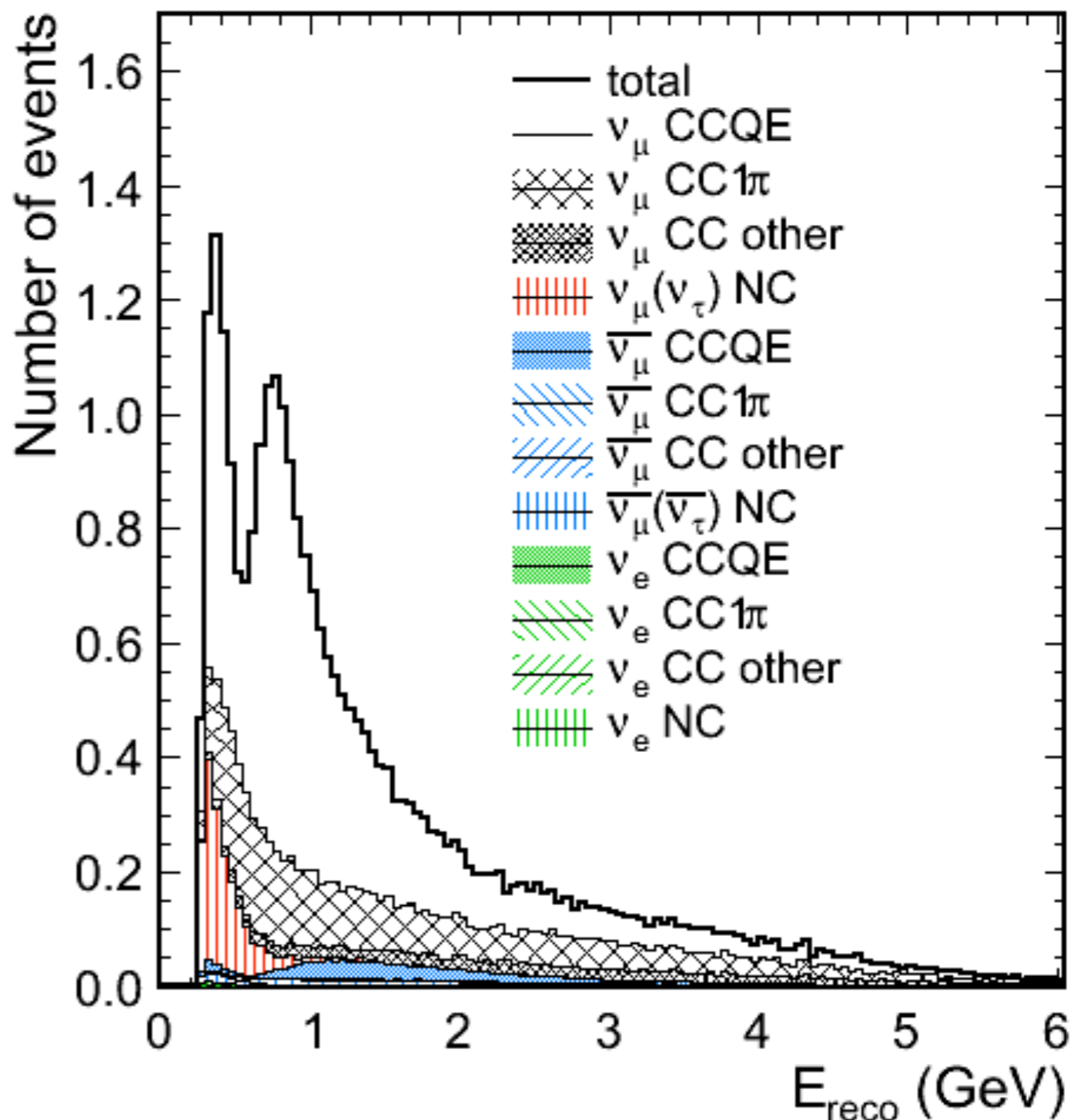
$\theta_{23}$

# T2K $\nu_\mu$ disappearance

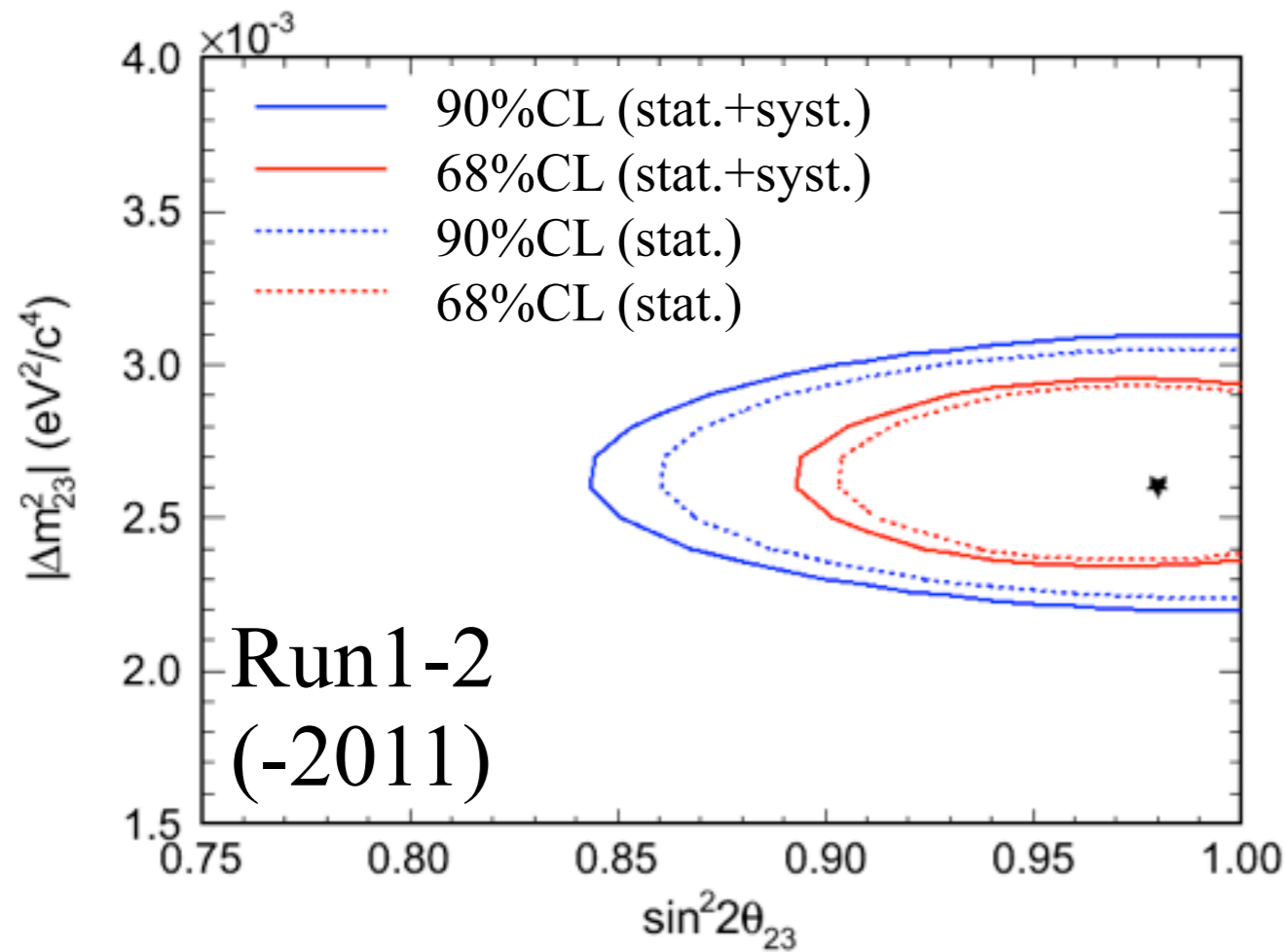


# Backgrounds

- Main backgrounds at the dip:
  - CC1 $\pi$
  - NC1 $\pi$
  - CCQE resolution tail



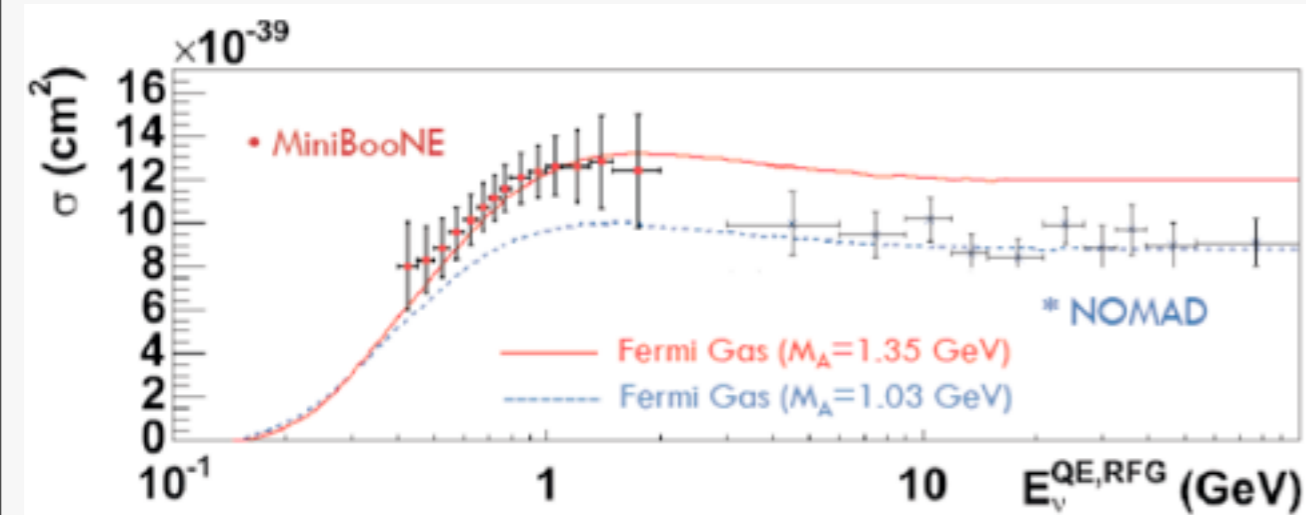
# Systematics in $\nu_\mu$ disap.



- Run1-3 (-2012): x2 stat.
- Run1-4 (-2014): x4 stat. expected
- Systematics started to impact
  - NC $\pi$ , CC1 $\pi$ (CCnonQE), FSI
  - ⇒  $\pi$  detection is the key

Error source	$\delta N_{SK}^{exp} / N_{SK}^{exp}$
SuperK CCQE efficiency (ring-counting)	$\pm 3.23\%$
SuperK CCQE efficiency (other)	$\pm 1.04\%$
SuperK CCnonQE efficiency	$\pm 6.51\%$
SuperK NC efficiency	$\pm 6.96\%$
SuperK $\nu_e$ CC efficiency	$\pm 0.05\%$
ND280 efficiency	+5.59% -5.30%
ND280 normalization	$\pm 2.62\%$
Flux normalization	$\pm 4.73\%$
CCQE cross section	$\pm 2.36\%$
CC1 $\pi$ /CCQE cross section ratio	+0.52% -0.59%
CCother/CCQE cross section ratio	+4.10% -3.67%
NC/CCQE cross section ratio	+0.82% -0.84%
FSI	$\pm 5.80\%$
Total	+14.79% -14.61%

# Neutrino cross section



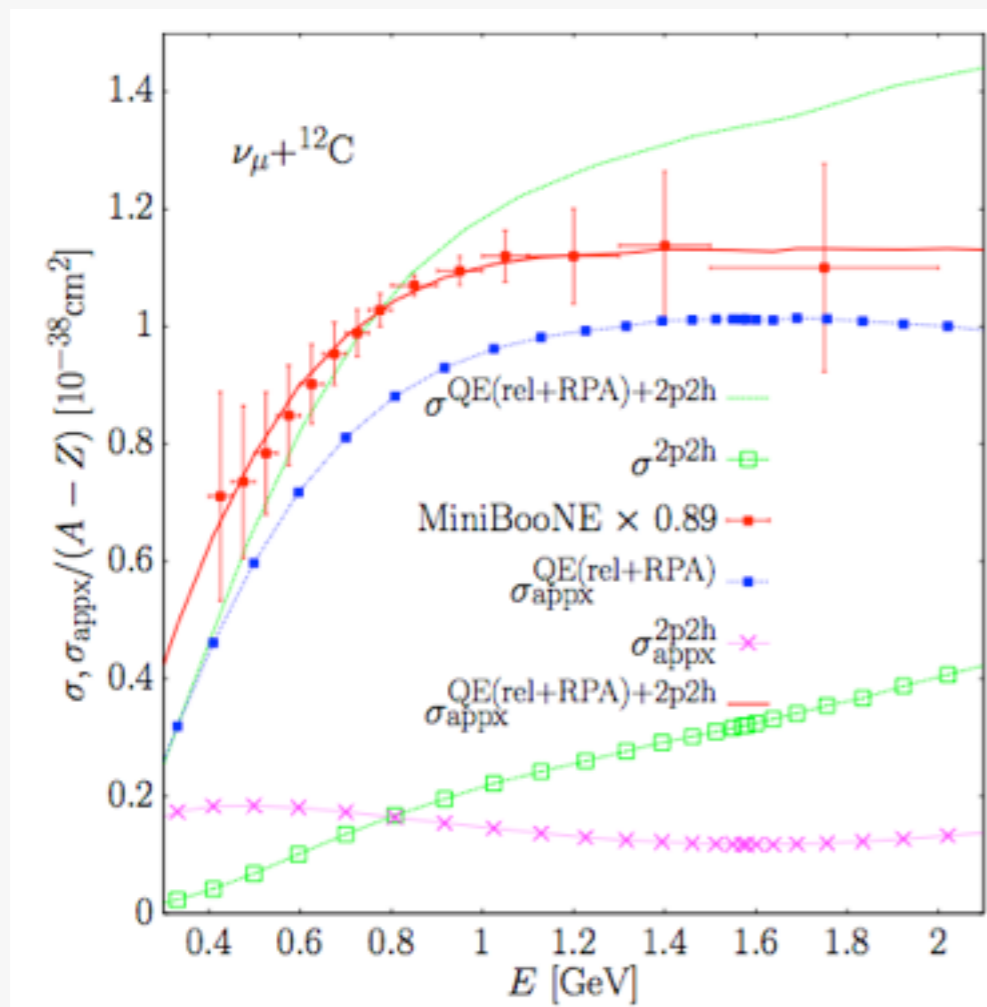
- CCQE cross section shows enhancement at  $\sim$ GeV

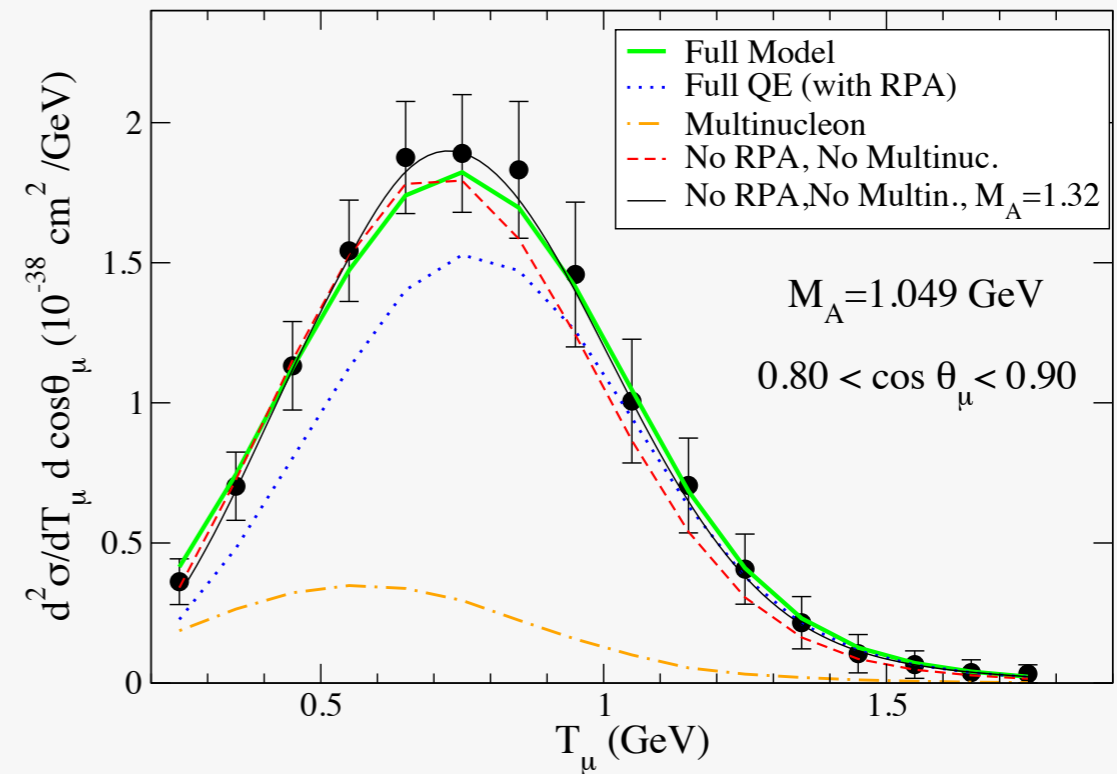
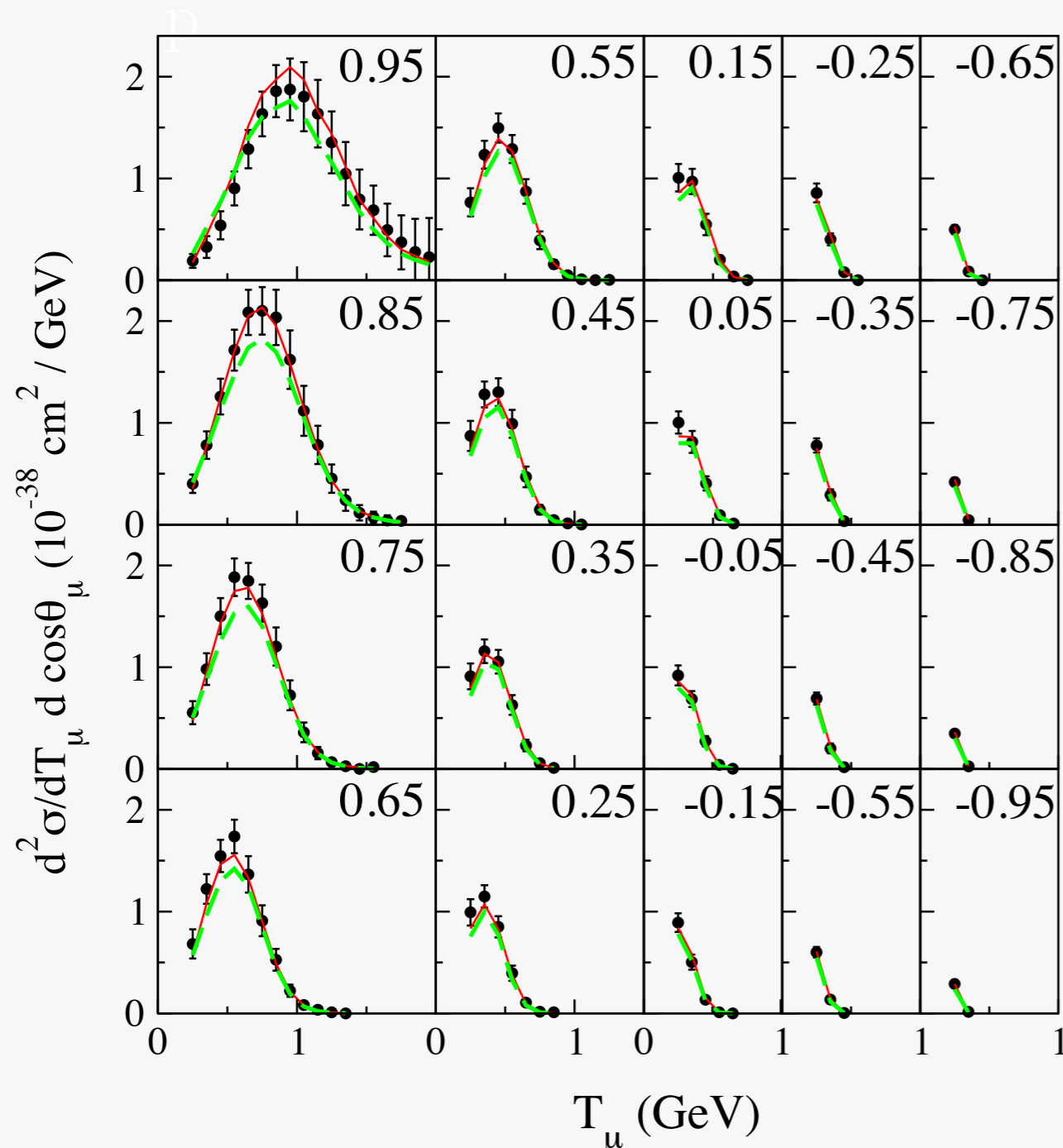
– Multi-nucleon process: 20–30% contributions?

– Reconstructed neutrino energy is not correct:

- It assumes 2 body react. with proton recoil

– Neutrino reaction model need to be developed!





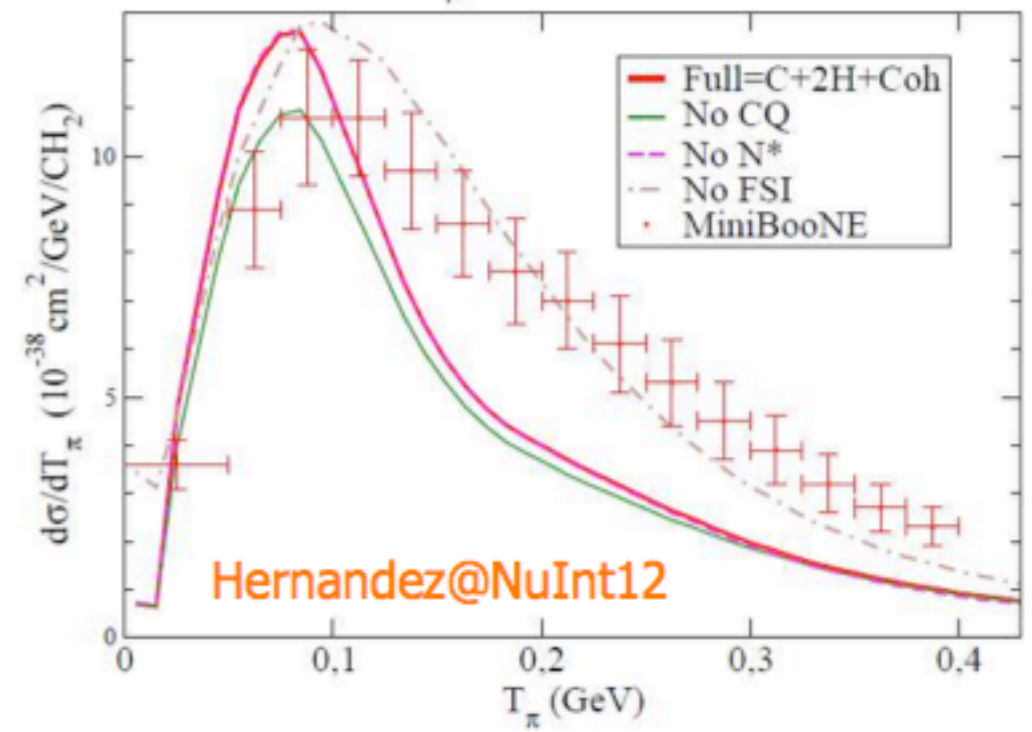
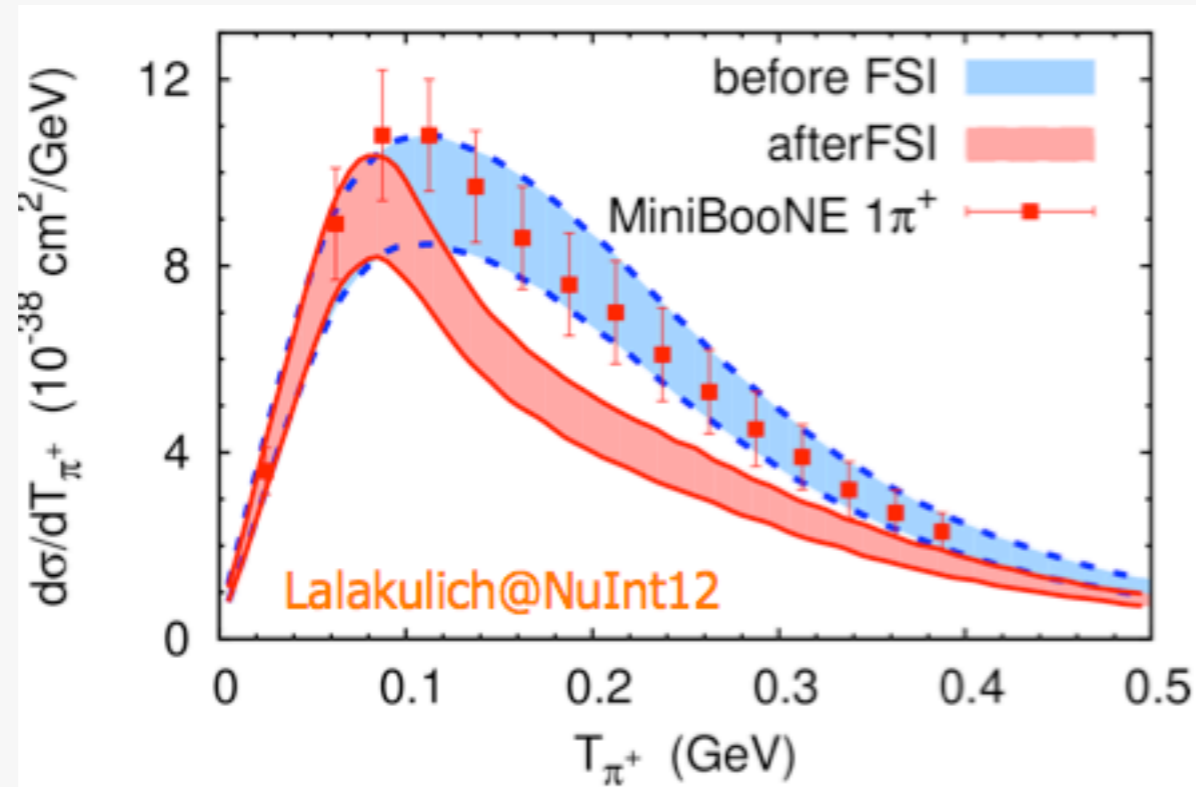
Model	Scale	$M_A$ (GeV)	$\frac{\chi^2}{\text{\#bins}}$
LFG	$0.96 \pm 0.03$	$1.32 \pm 0.03$	35/137
<b>Full</b>	<b><math>0.92 \pm 0.03</math></b>	<b><math>1.08 \pm 0.03</math></b>	<b>50/137</b>
Full $ q  > 0.4^\dagger$ GeV	$0.83 \pm 0.04$	$1.01 \pm 0.03$	30/123

$^\dagger$  : As suggested by Sobczyk et al. PRC 82, 045502

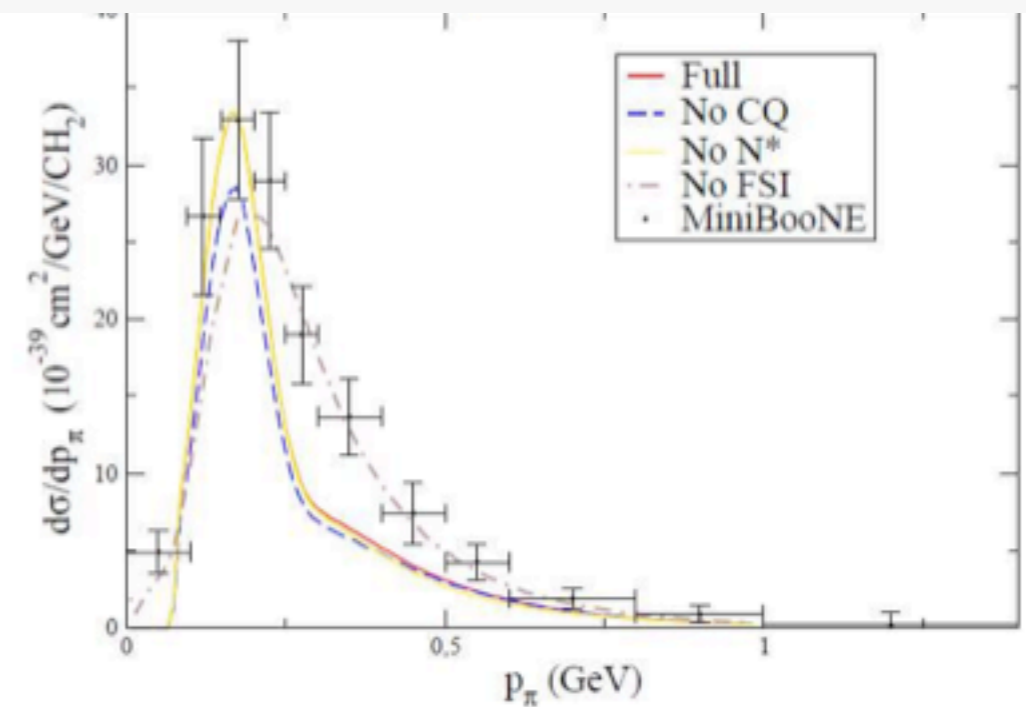
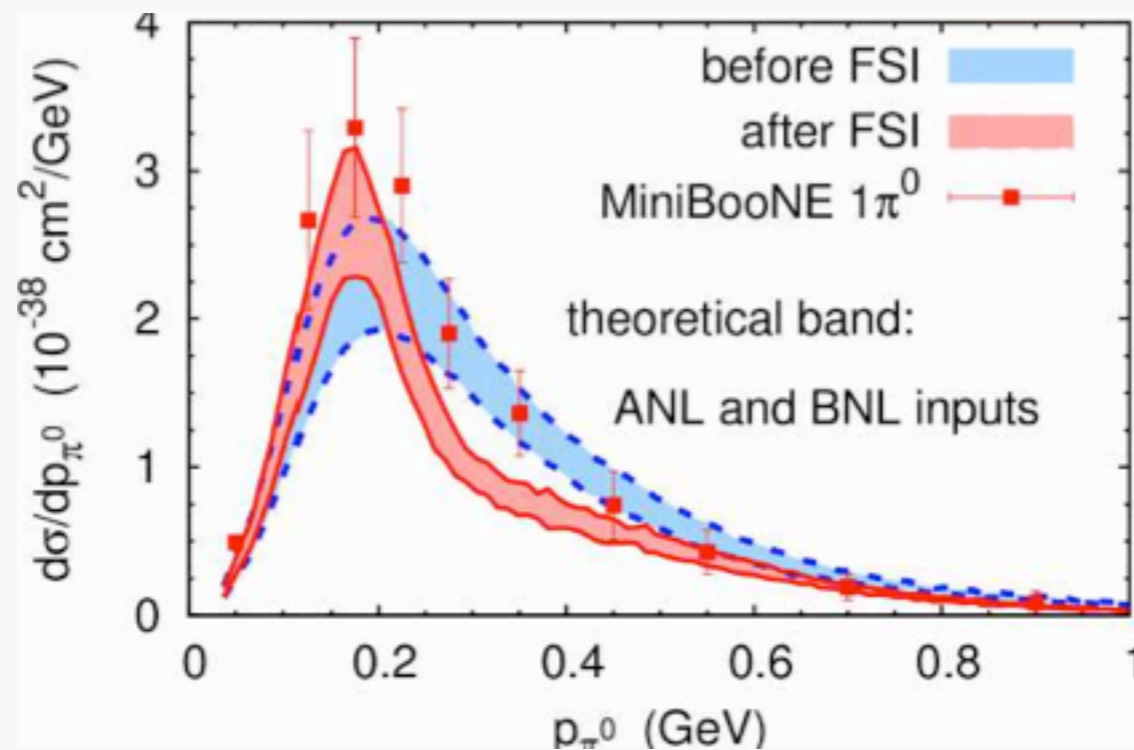
Model with 2p2h and RPA fits MiniBooNE CCQE data well when data is scaled by 0.89 (within MiniBooNE error)

# CC1 $\pi$ consistent with no FSI?

CC1 $\pi^+$



CC1 $\pi^0$



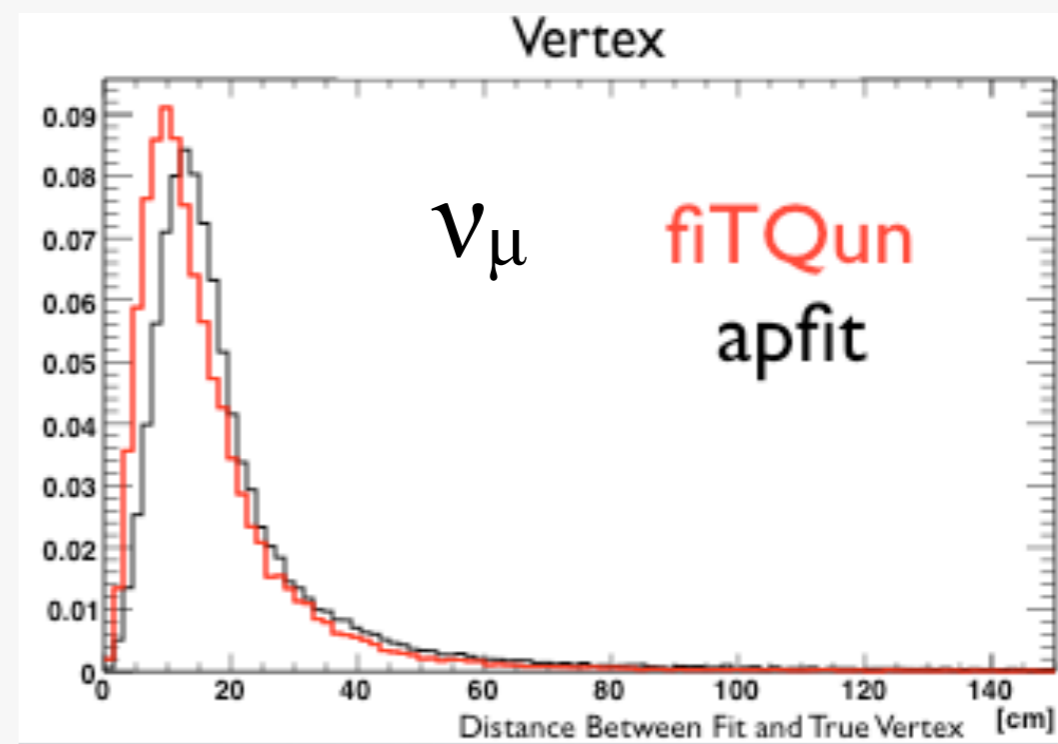
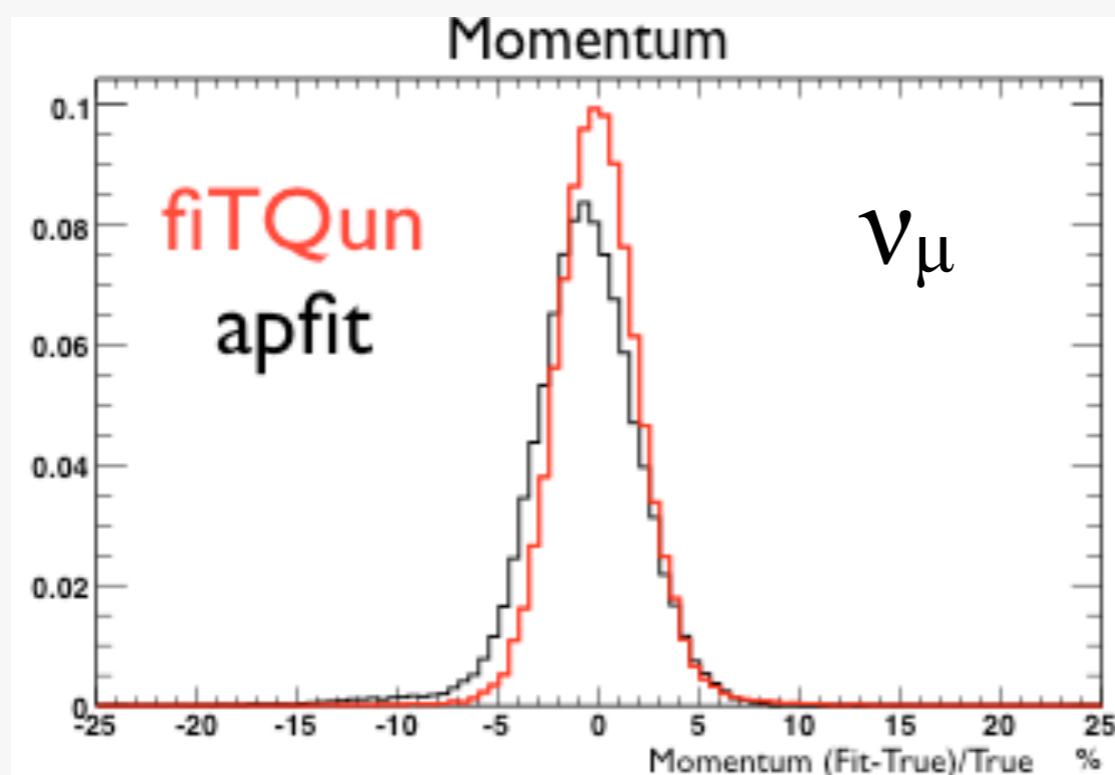
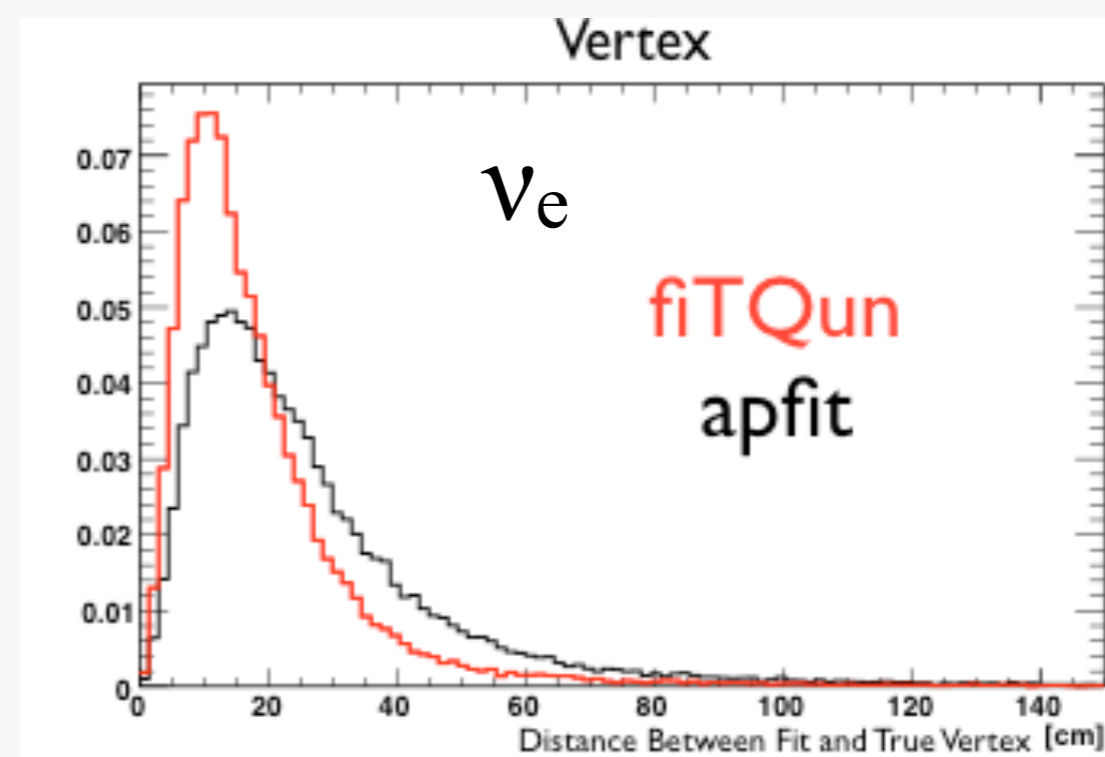
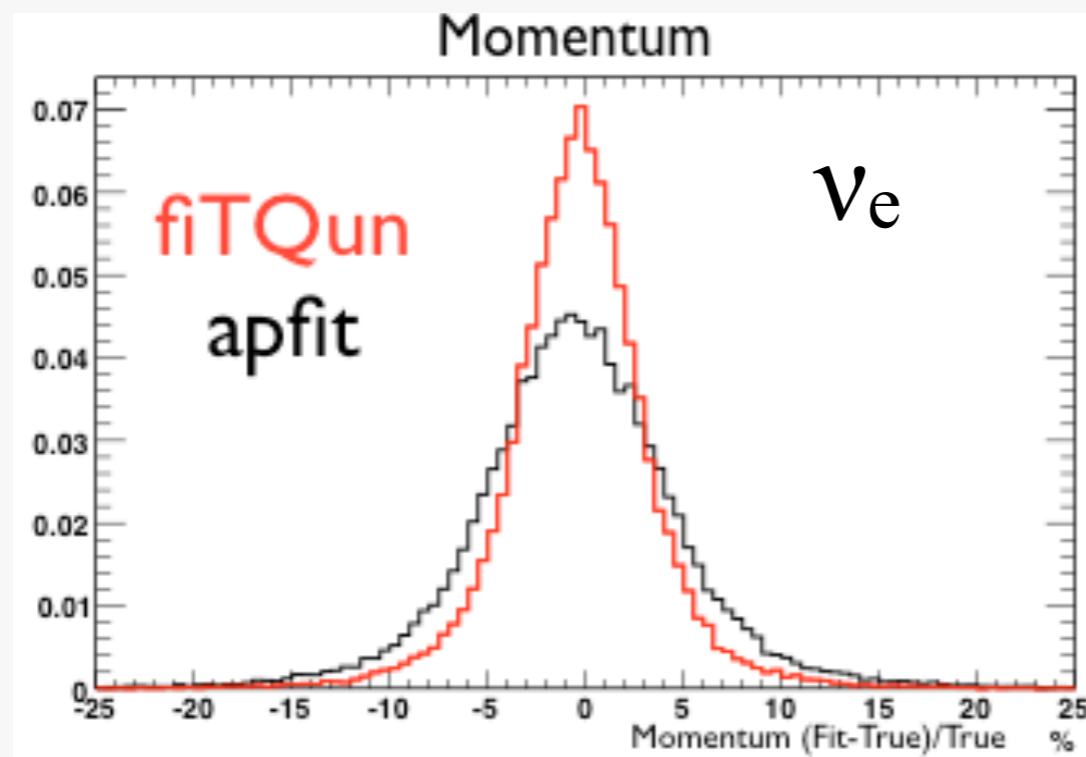
# New SK Reconstion code

- Current SK reconstruction code (APFIT)
  - Originally written for Kamiokande experiment.
  - Works well with detailed check with data.
  - Limitation in the design for limited CPU power.
- New SK reconstruction code (fitQun) M. Wilking et. al.
  - Likelihood fit of all the PMT charge and timing
    - taking advantage of all available information
    - simultaneous fit of vertex and particle kinematics
    - based on the method that worked well for MiniBooNE
  - To be used for Hyper-K event reconstruction, too.

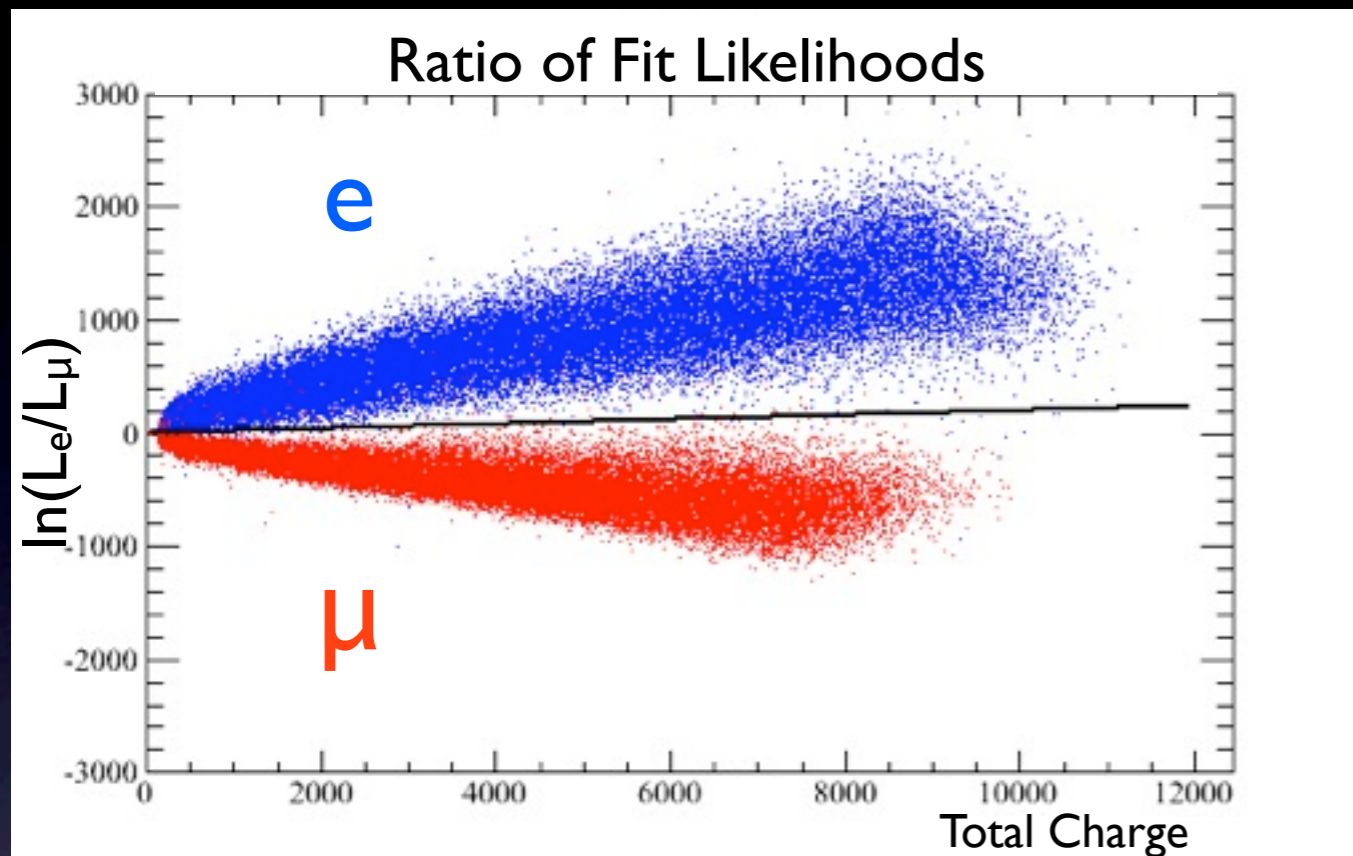


# Single ring with FitQun

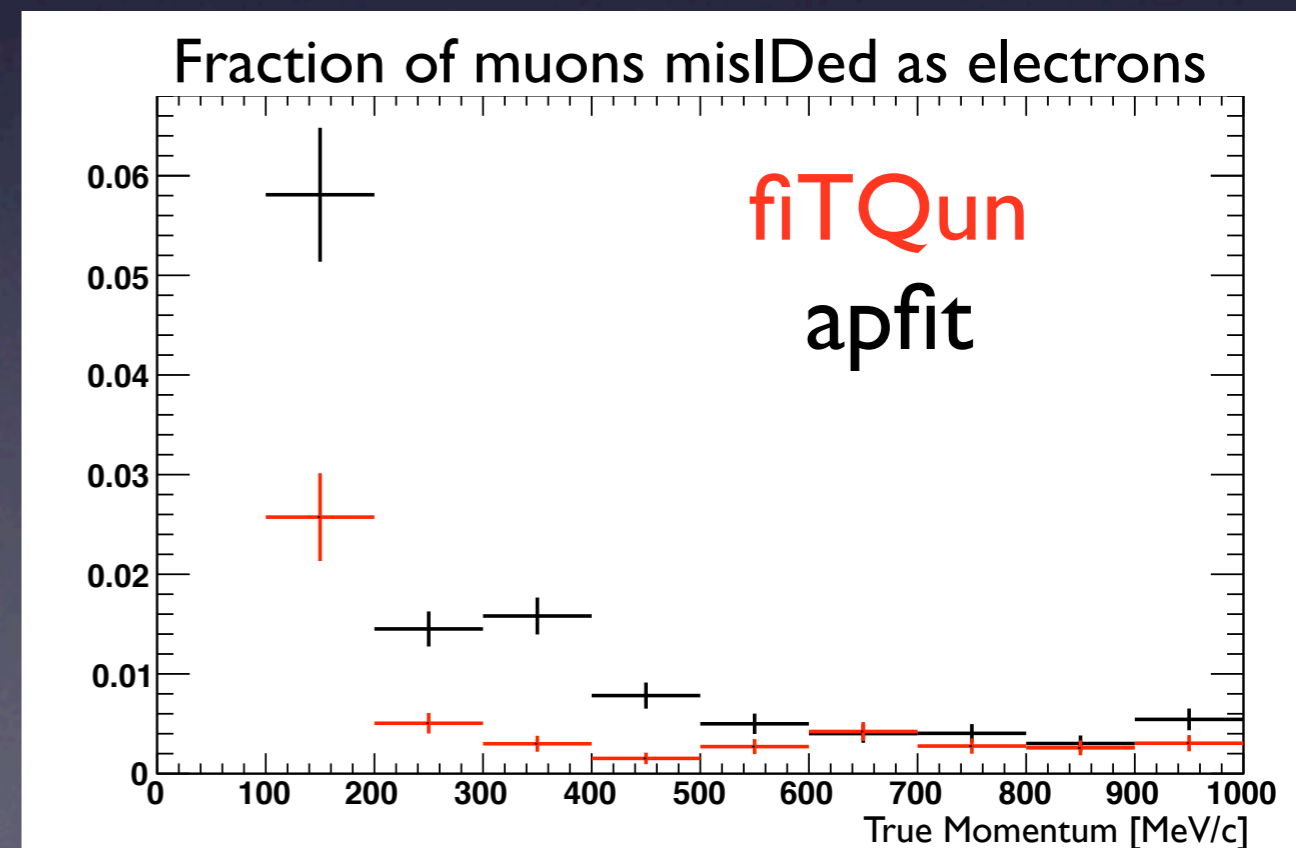
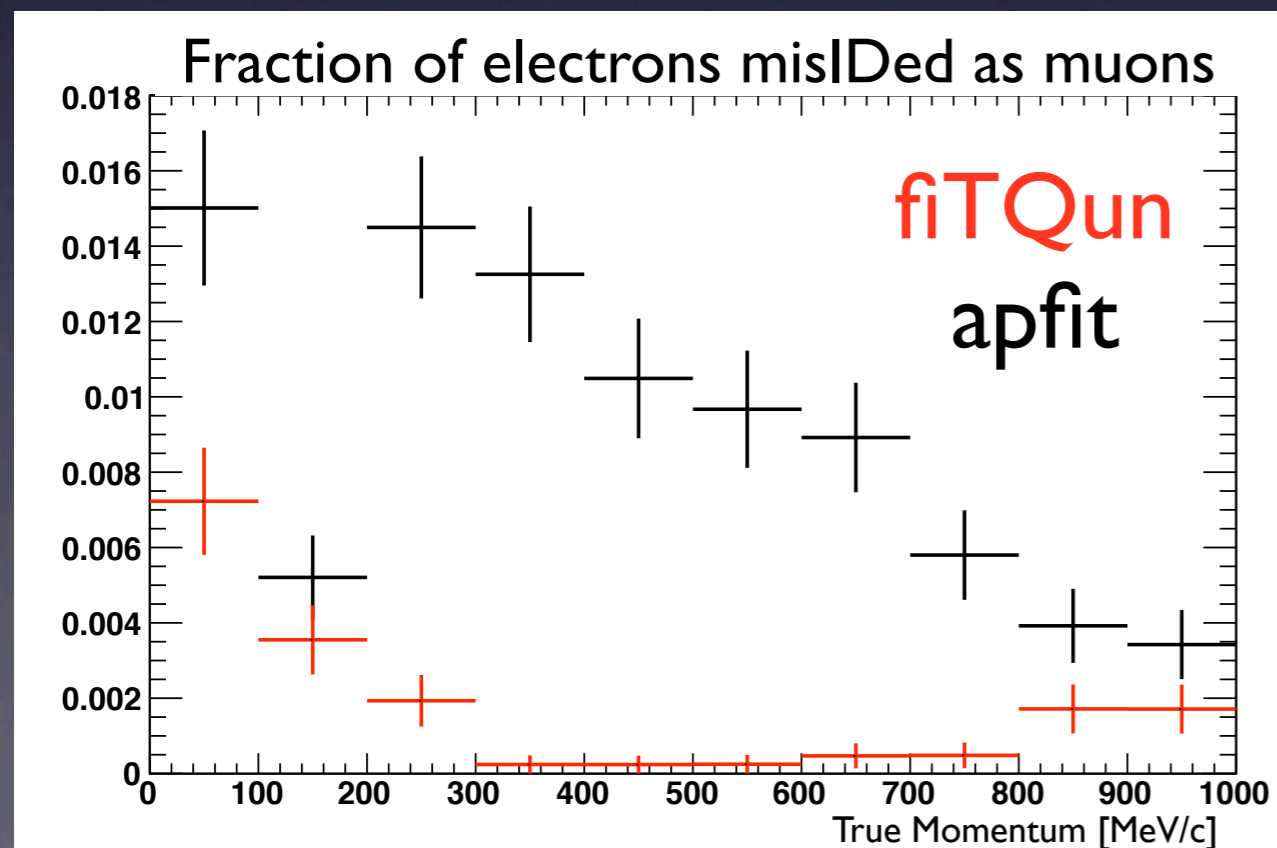
M. Wilking@HyperK meeting



# Single Track Particle ID

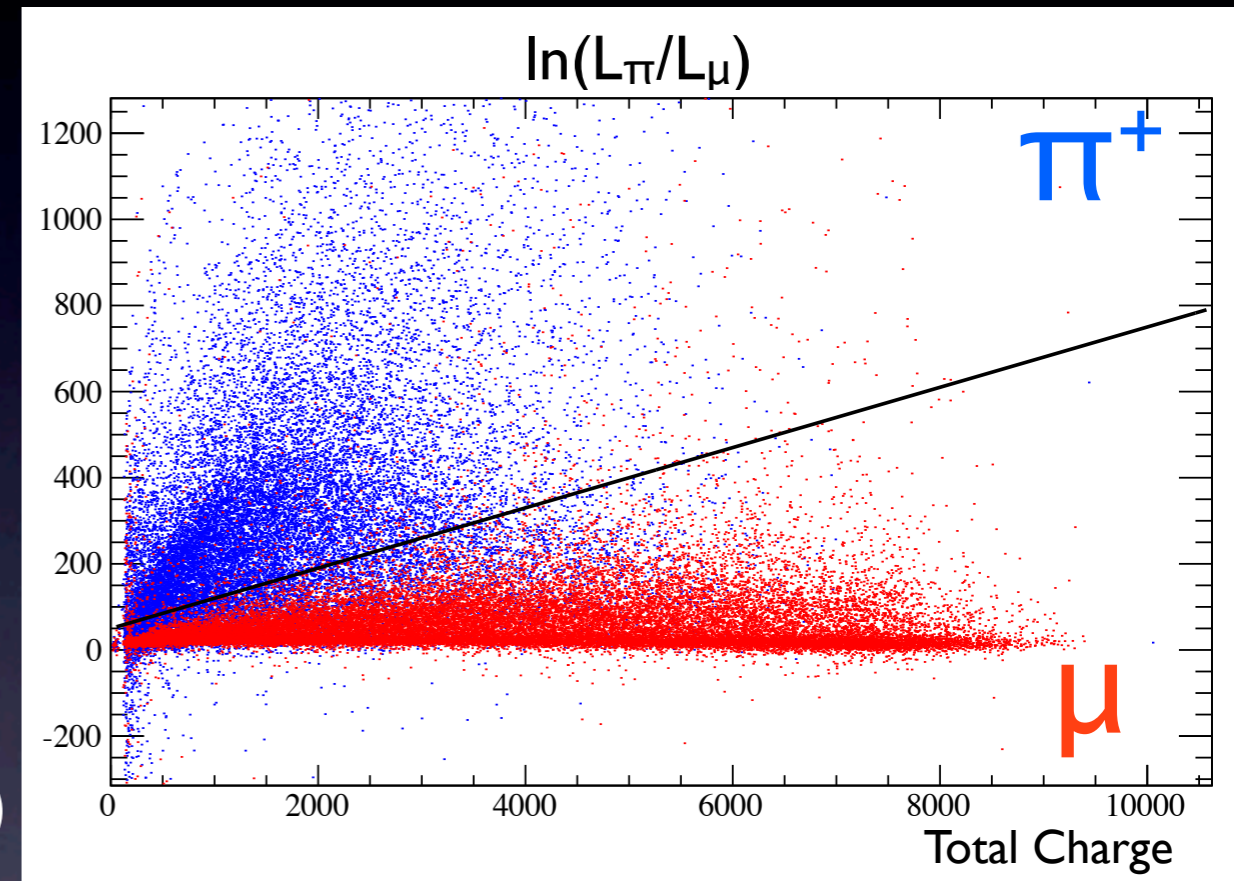


- Simple line cut can be used to separate muons and electrons
- Significantly improved particle ID



# $\pi^+$ Fitter

- Pions and muons propagate and produce Cherenkov light in a very similar manner (similar masses)
- The main difference is due to hadronic interactions
- Ring pattern observed is a “kinked” pion trajectory
- This is the first demonstration of pion/muon separation at SK (in MC)
- Allows for  $CC\pi^+ E_\nu$  reconstruction



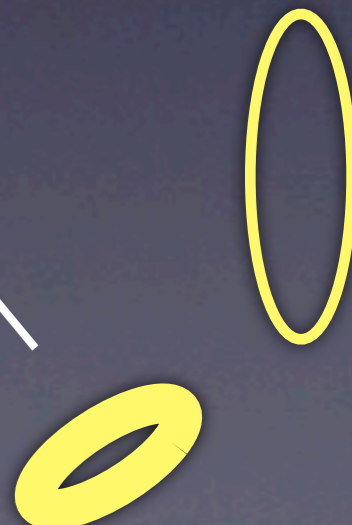
electron tracks



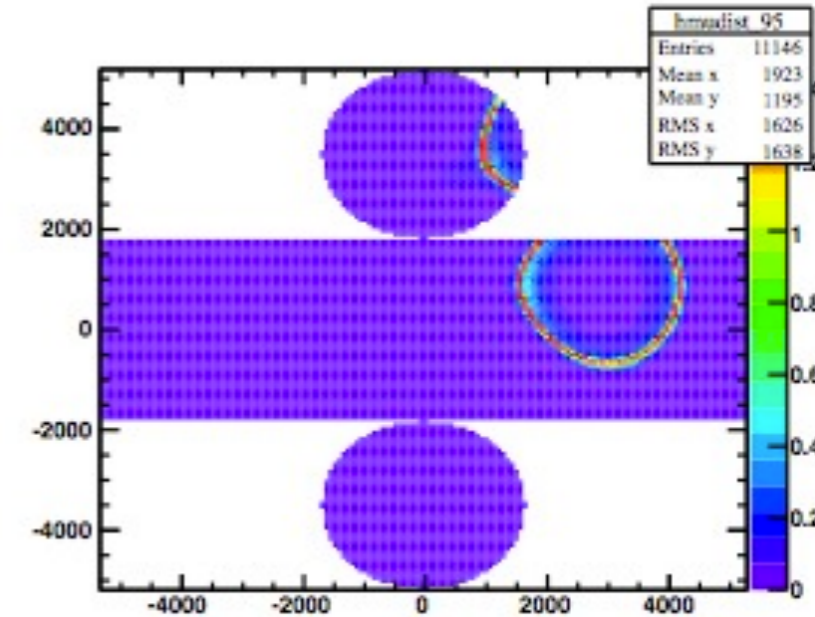
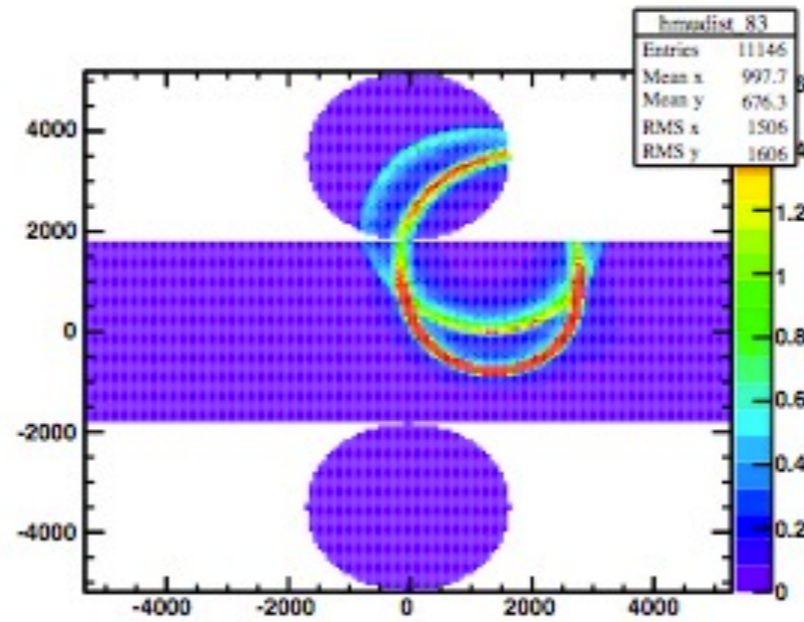
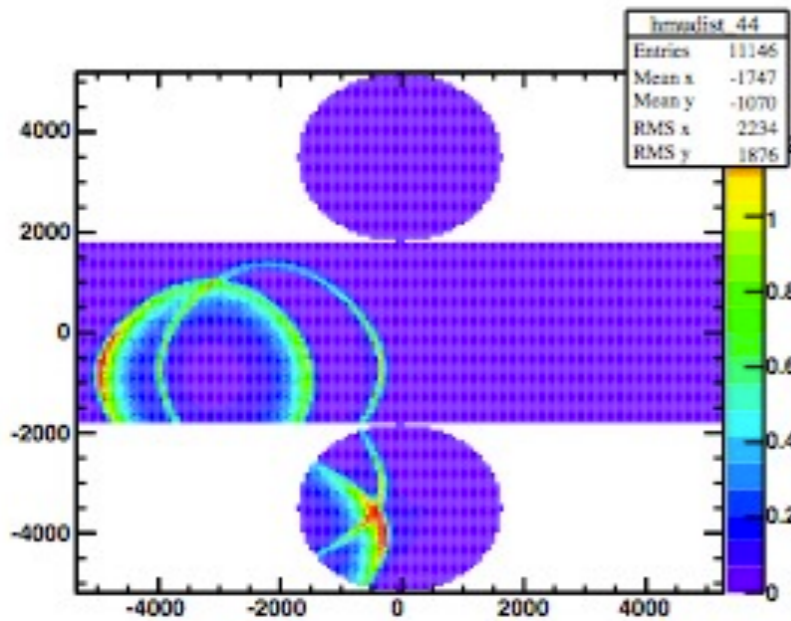
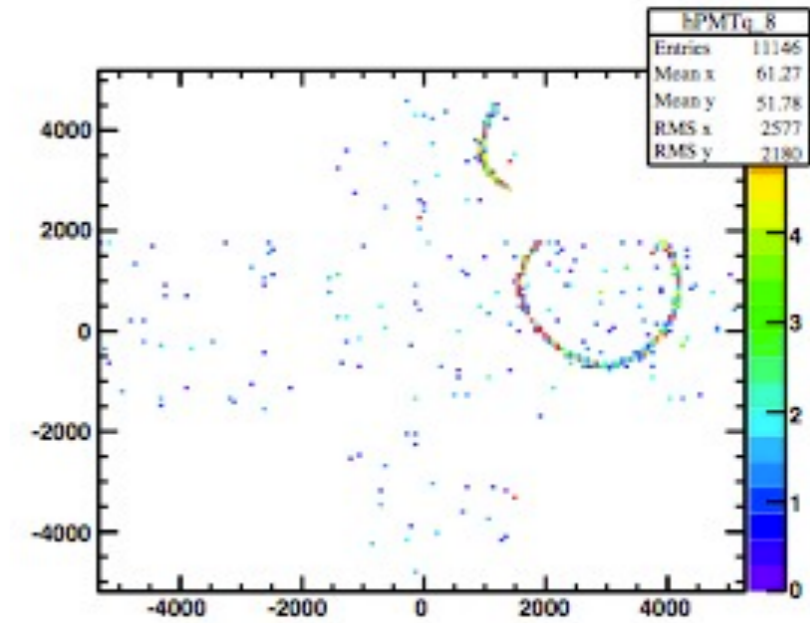
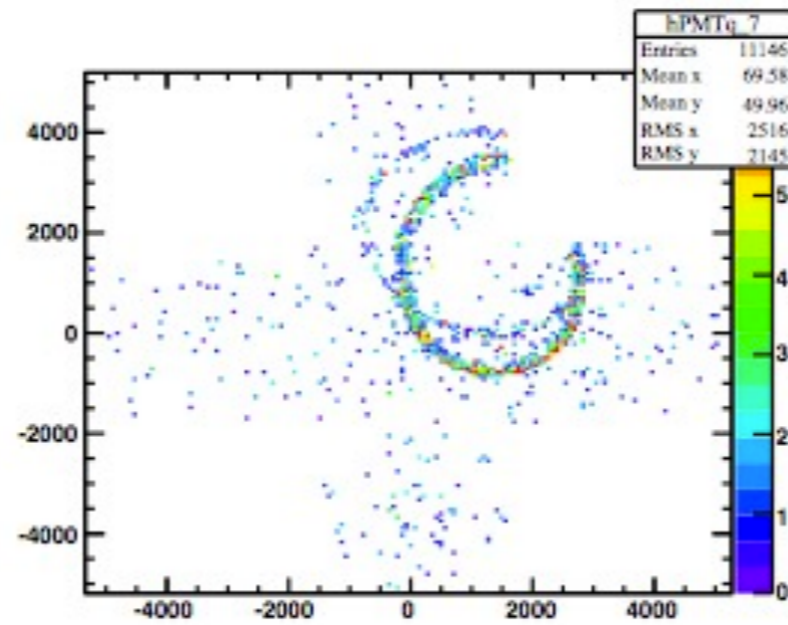
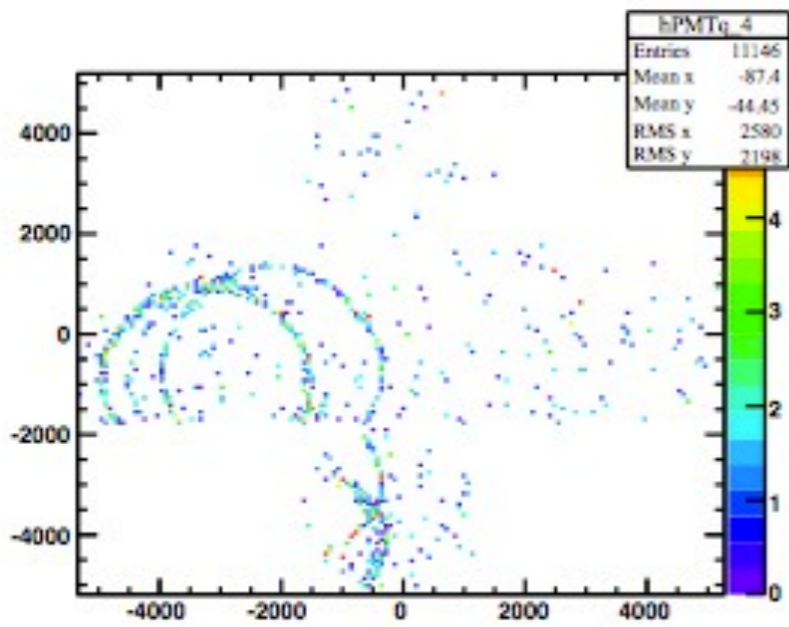
muon tracks



pion tracks



# Kinked-track $\pi^+$ Fitter



# Potential improvements

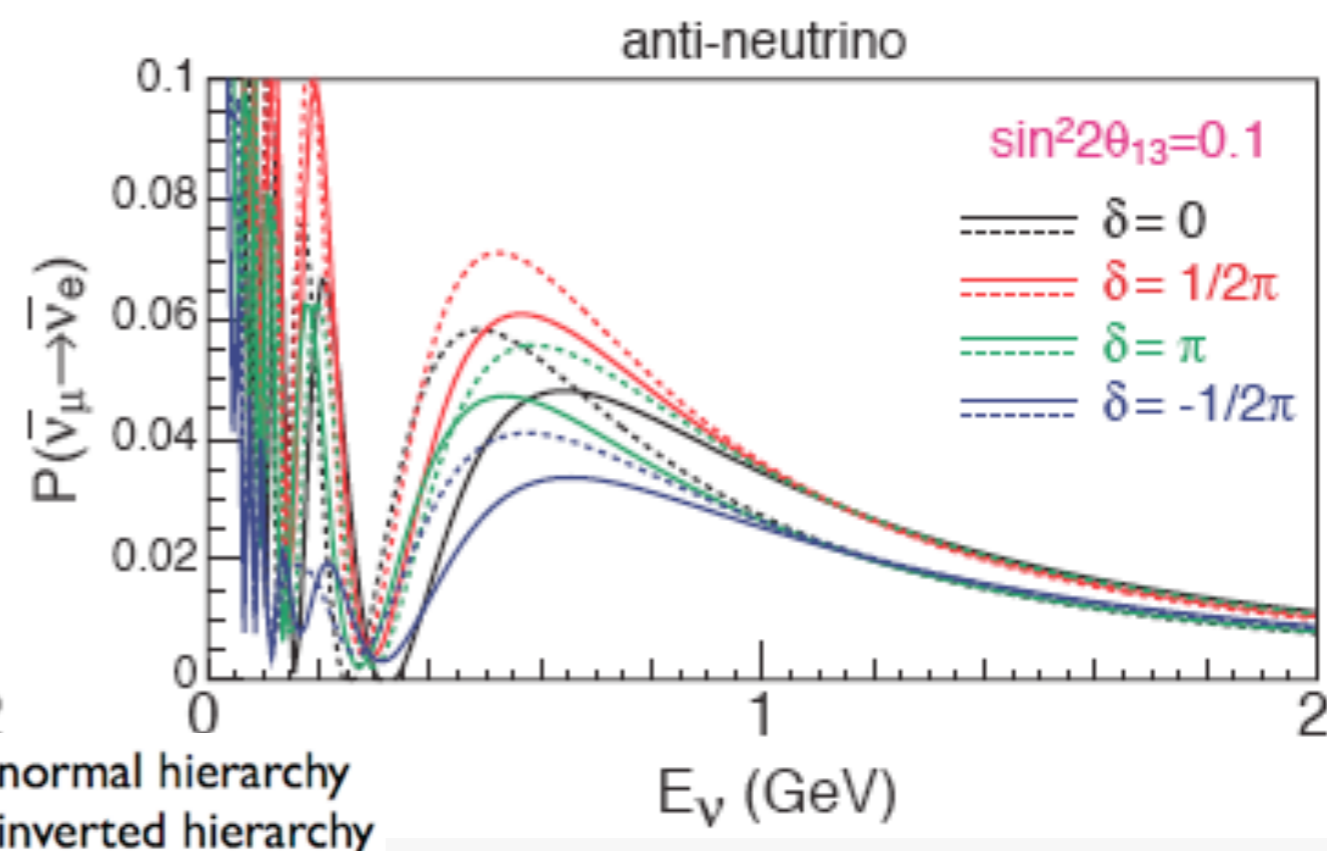
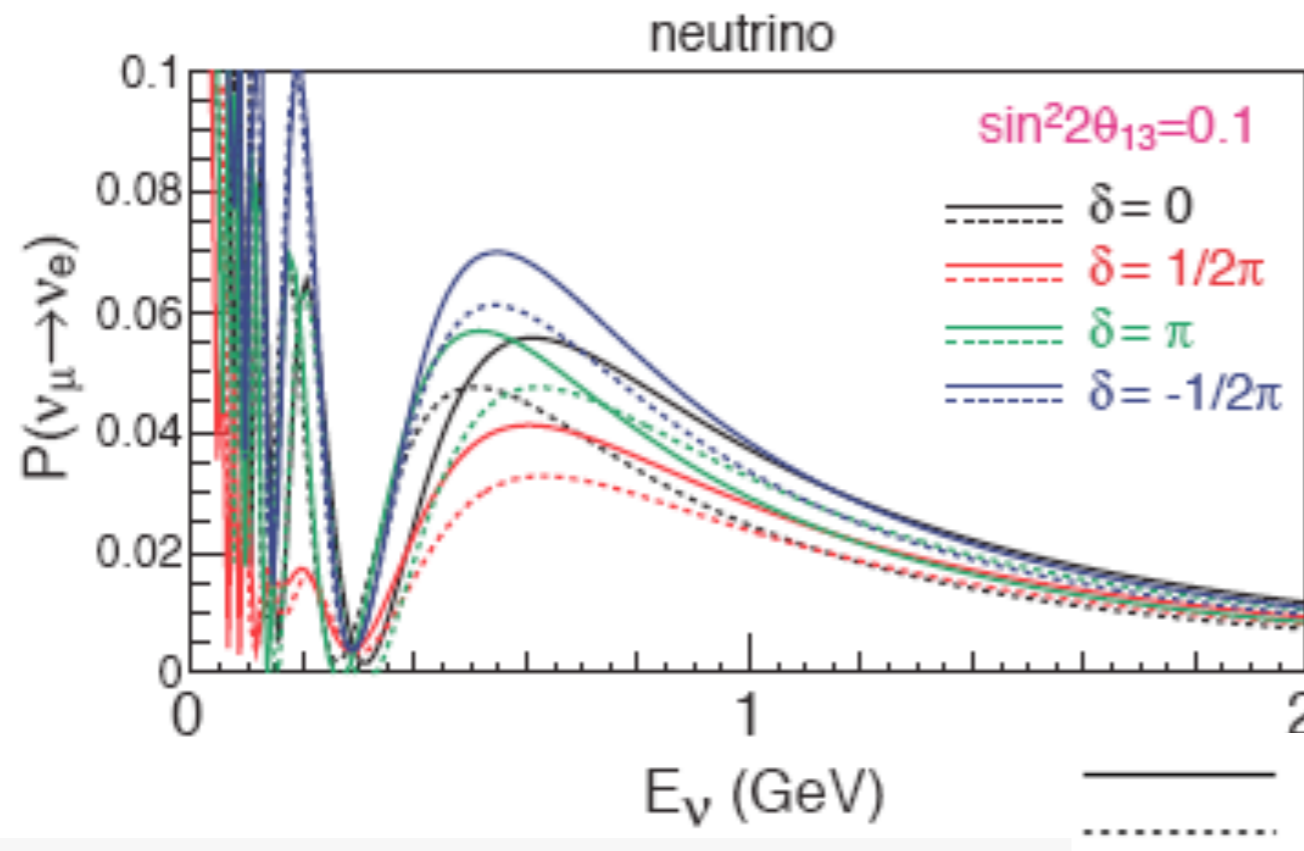
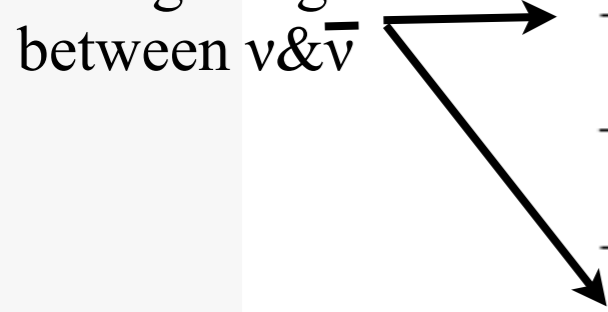
- Background suppression & measurement (fitQun)
  - $\nu_\mu$  disappearance ( $\theta_{23}$ ):  $NC\pi^+, CC\pi^+$
- Reduction of systematic uncertainties
  - Background suppression & measurement
  - Cross section modeling and measurements
- Statistics to improve:
  - T2K: runs with continuously increasing intensity
  - Hyper-K: 25 time more fiducial volume

# $\theta_{13}$ and CP

# $\nu_e$ Appearance probability

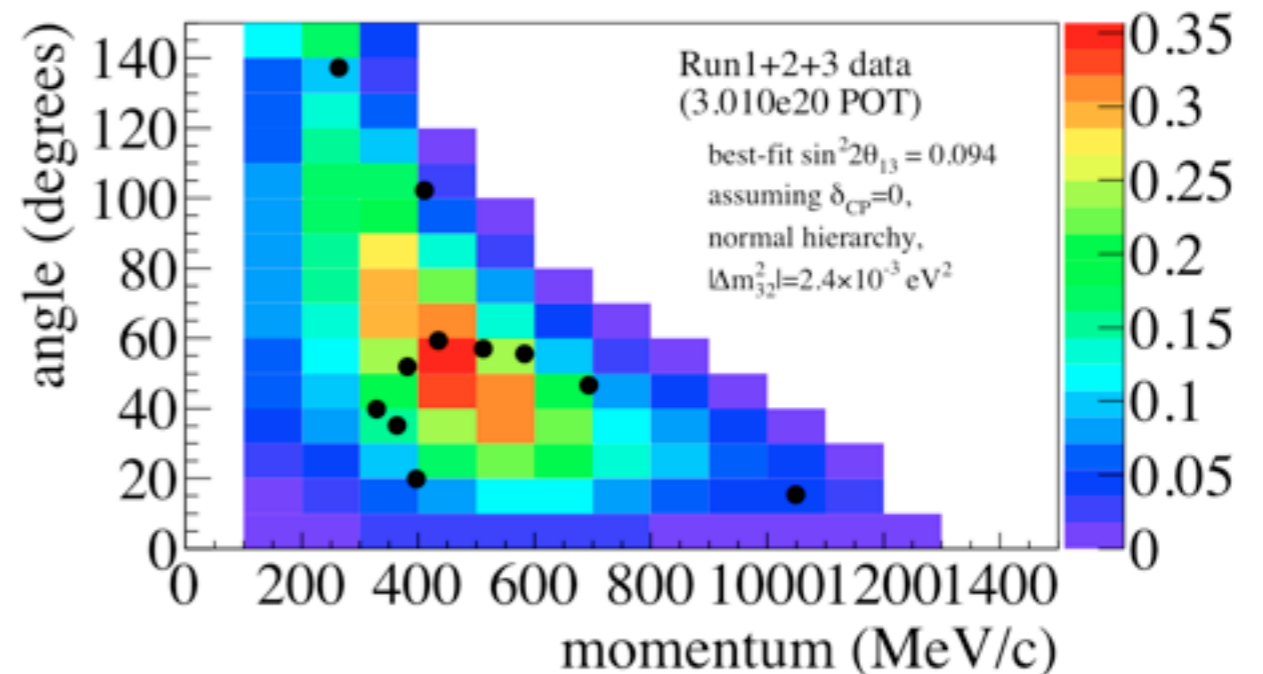
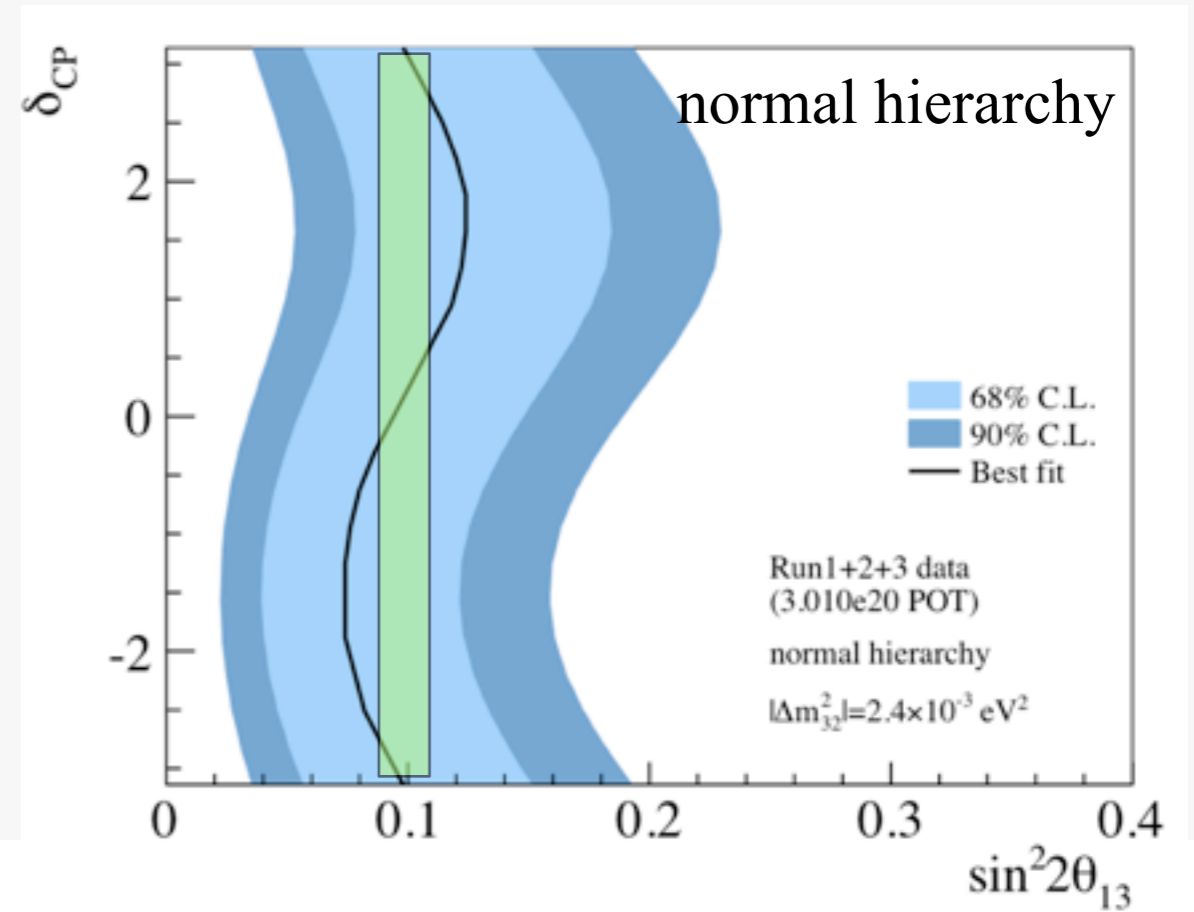
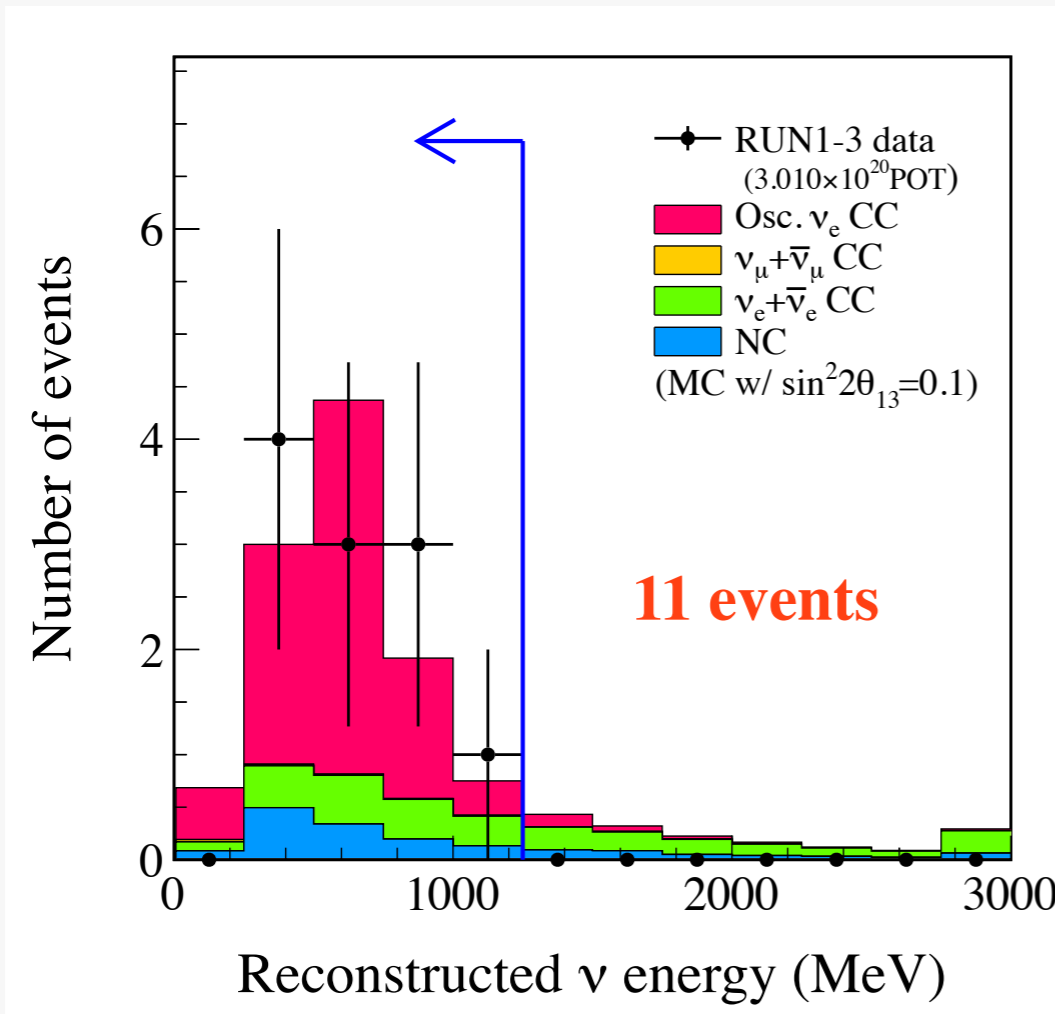
$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) = & 4C_{13}^2 S_{13}^2 S_{23}^2 \sin^2 \Delta_{31} \quad \text{Leading term} \\
 & + 8C_{13}^2 S_{12} S_{13} S_{23} (C_{12} C_{23} \cos \delta - S_{12} S_{13} S_{23}) \cos \Delta_{32} \cdot \sin \Delta_{31} \cdot \sin \Delta_{21} \quad \text{CP conserv.} \\
 & - 8C_{13}^2 C_{12} C_{23} S_{12} S_{13} S_{23} \sin \delta \cdot \sin \Delta_{32} \cdot \sin \Delta_{31} \cdot \sin \Delta_{21} \quad \text{CP violating} \\
 & + 4S_{12}^2 C_{13}^2 (C_{12}^2 C_{23}^2 + S_{12}^2 S_{23}^2 S_{13}^2 - 2C_{12} C_{23} S_{12} S_{23} S_{13} \cos \delta) \sin^2 \Delta_{21} \quad \text{Solar} \\
 & - 8C_{13}^2 S_{13}^2 S_{23}^2 \cdot \frac{aL}{4E_\nu} (1 - 2S_{13}^2) \cos \Delta_{32} \cdot \sin \Delta_{31} \quad \text{Matter effect} \\
 & + 8C_{13}^2 S_{13}^2 S_{23}^2 \frac{a}{\Delta m_{31}^2} (1 - 2S_{13}^2) \sin^2 \Delta_{31}
 \end{aligned}$$

Changes sign  
between  $\nu$  &  $\bar{\nu}$



# T2K $\nu_e$ appearance

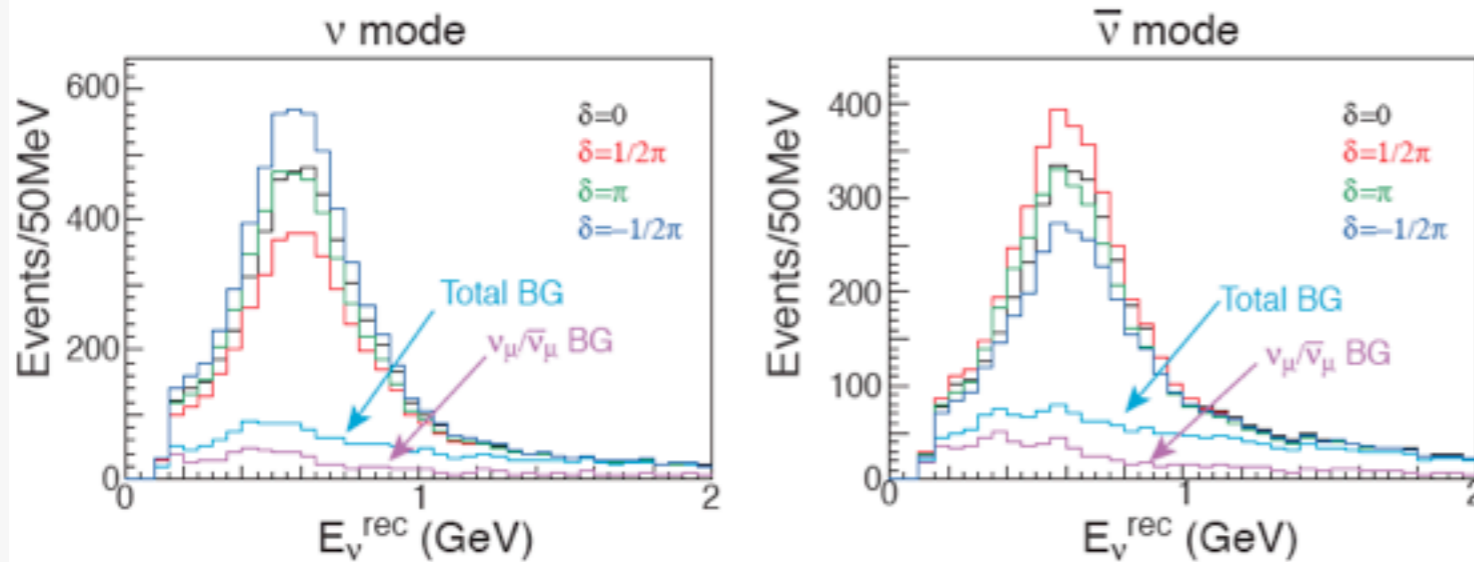
- Run1-3 (-2012) data  
-3.2 $\sigma$  evidence in  
 $\nu_e$  appearance



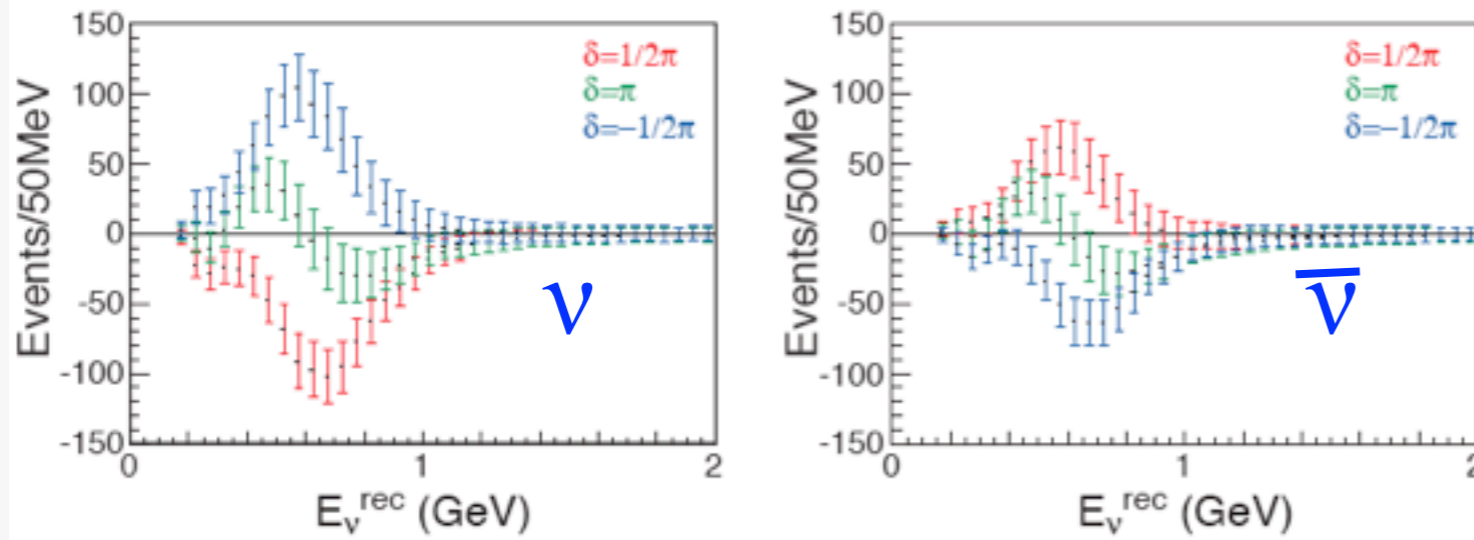


# CP violation @ HyperK

Event rate at HyperK

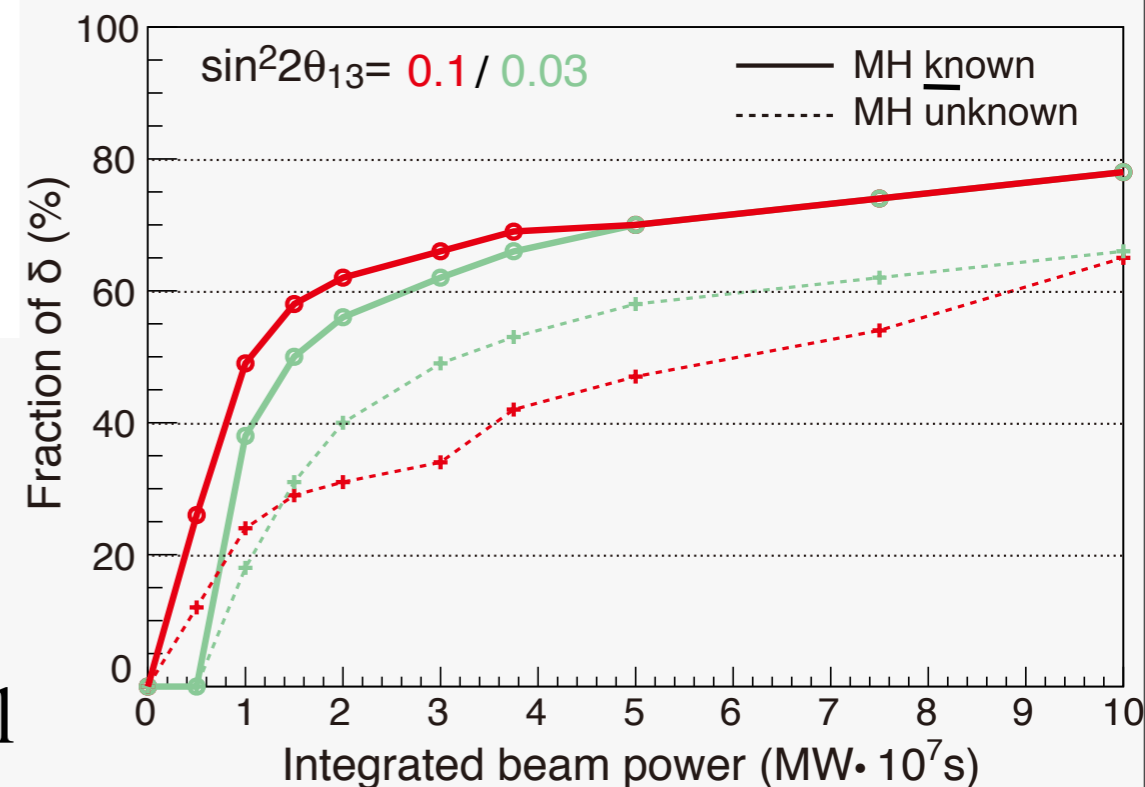


Event rate relative to  $\delta=0$



$\nu$  and  $\bar{\nu}$  rate changes with CP phase

$\delta=0$  and  $\pi$  are different in spectrum ( $\cos\delta$  term)

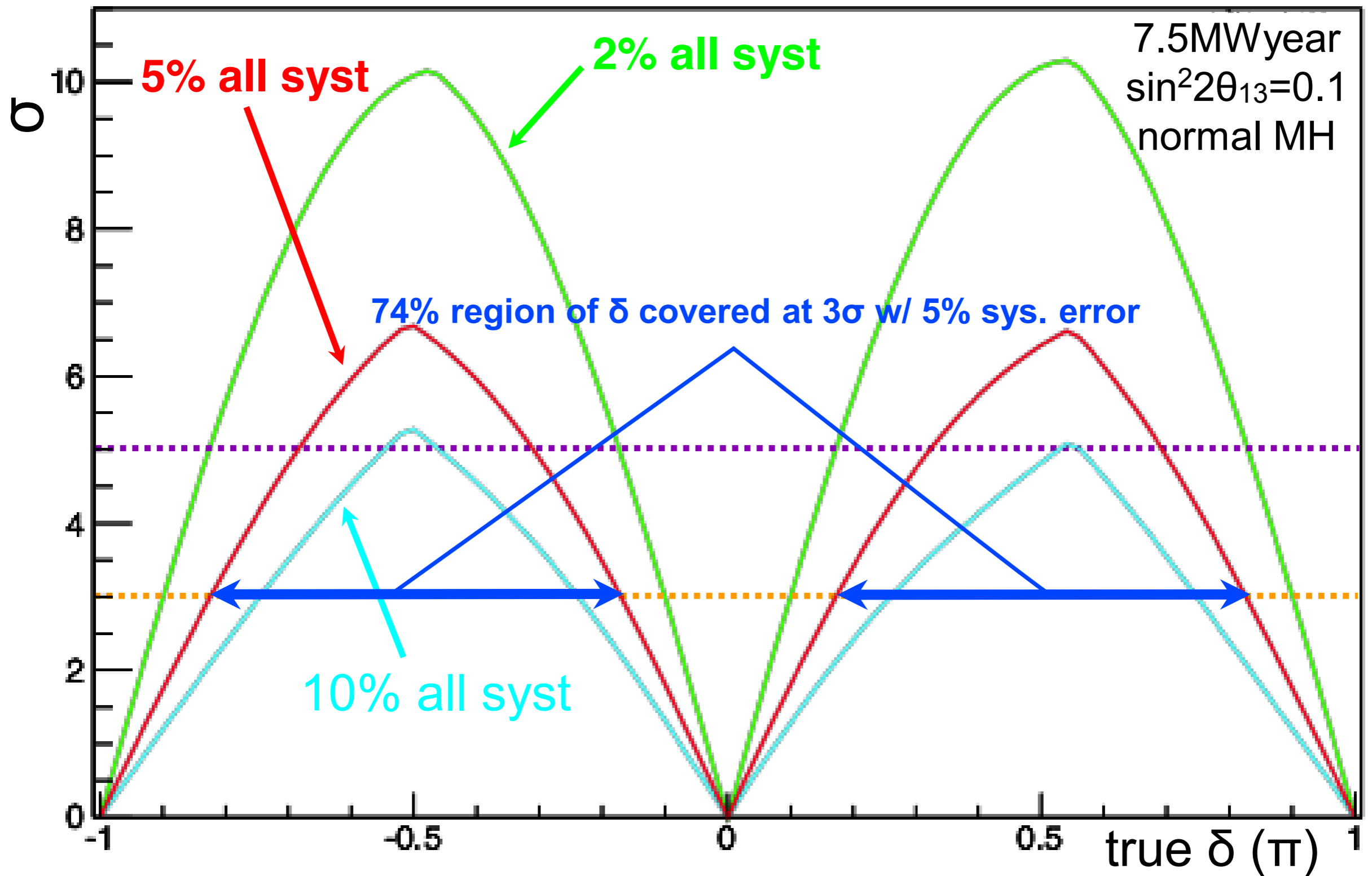


**Assumes overall 5% systematics realistic?**

Fraction of  $\delta$  phase space that can be observed at  $3\sigma$  level

# CPV Discovery Sensitivity

(w/ Mass Hierarchy known)

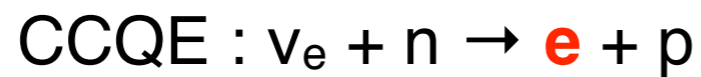
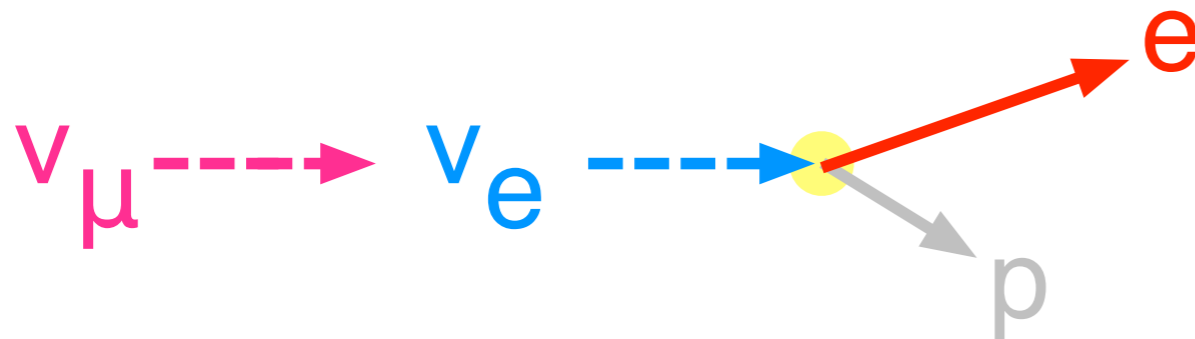


High Sensitivity to CPV w/  $\sim 5\%$  syst. error

# T2K Signal & Background for $\nu_e$ appearance

- Signal = **single electron event**

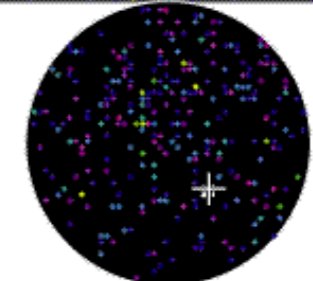
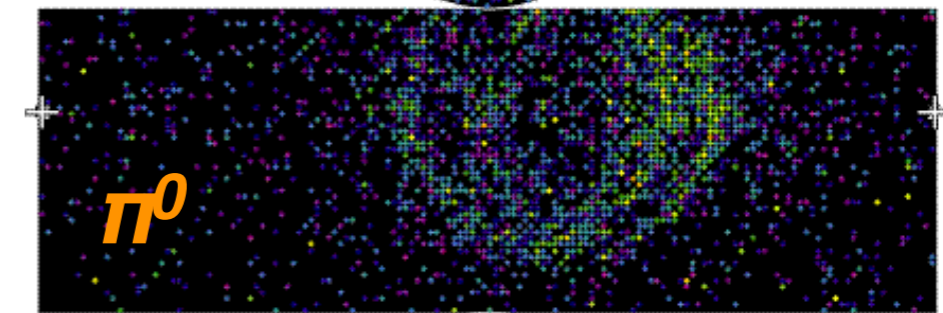
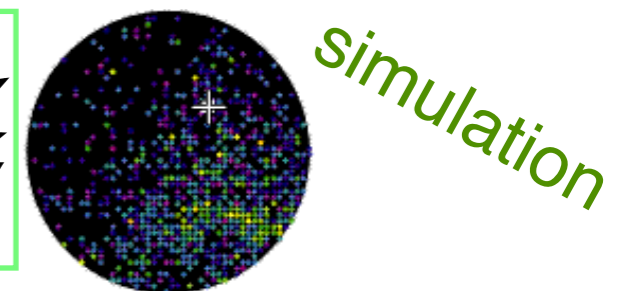
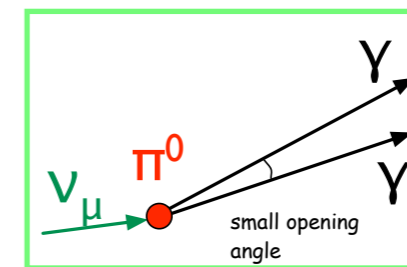
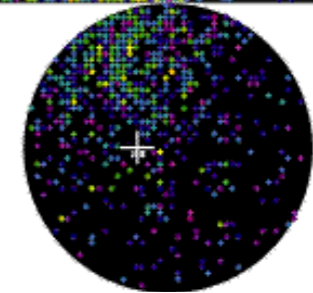
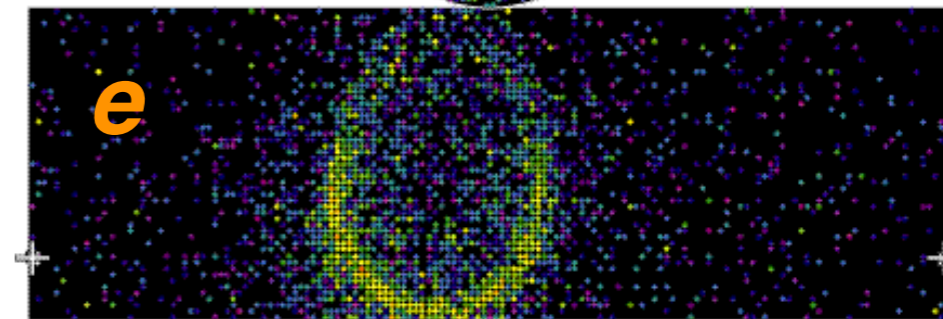
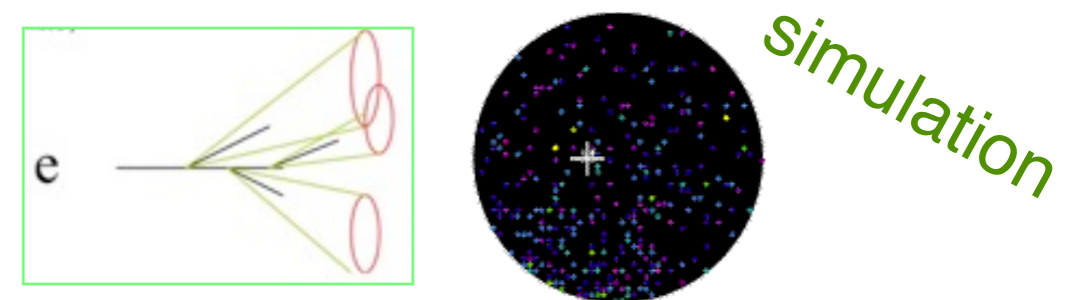
- oscillated  $\nu_e$  interaction :



(dominant process at T2K beam energy)

- Background

- intrinsic  $\nu_e$  in the beam (from  $\mu$ ,  $K$  decays)
- $\pi^0$  from NC interaction



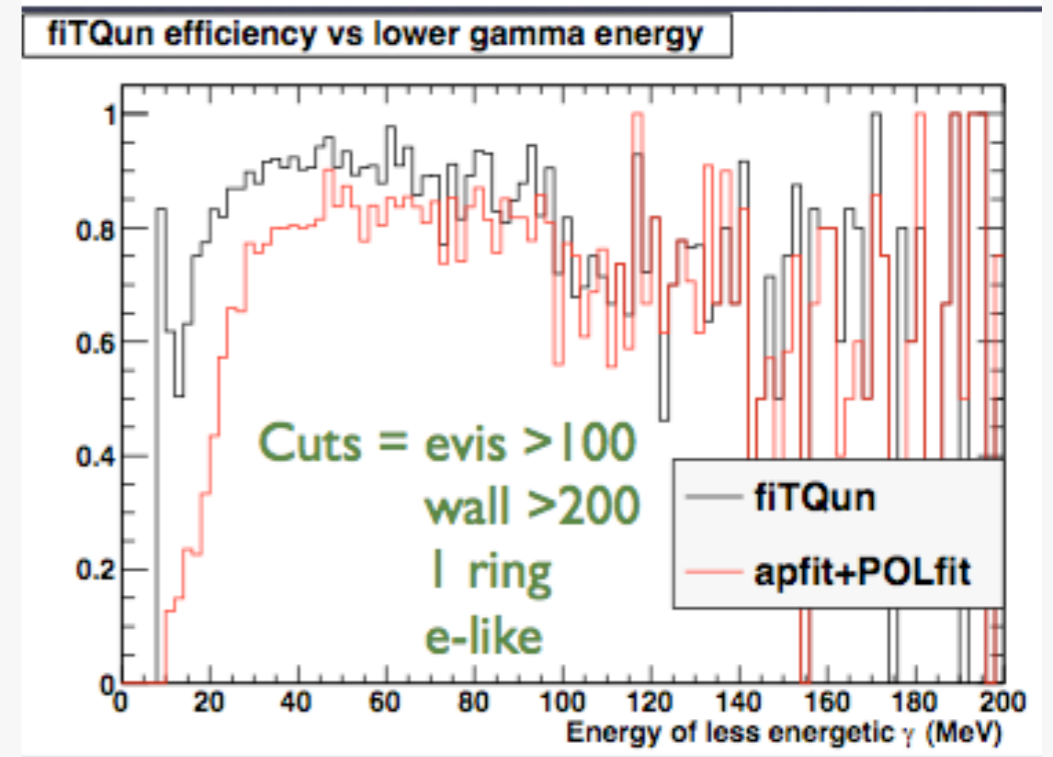
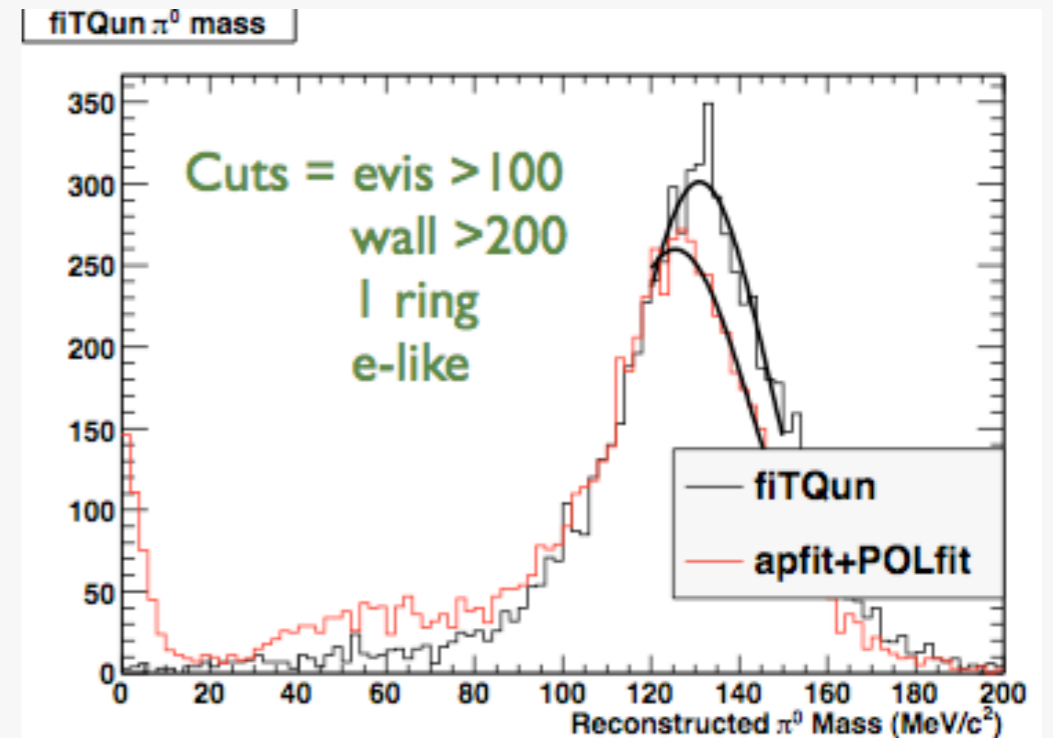
# Background suppression

- New reconstruction code suppress NC $\pi^0$  background in  $\nu_e$  appearance to  $\sim 1/3$

M. Wilking @HyperK meeting

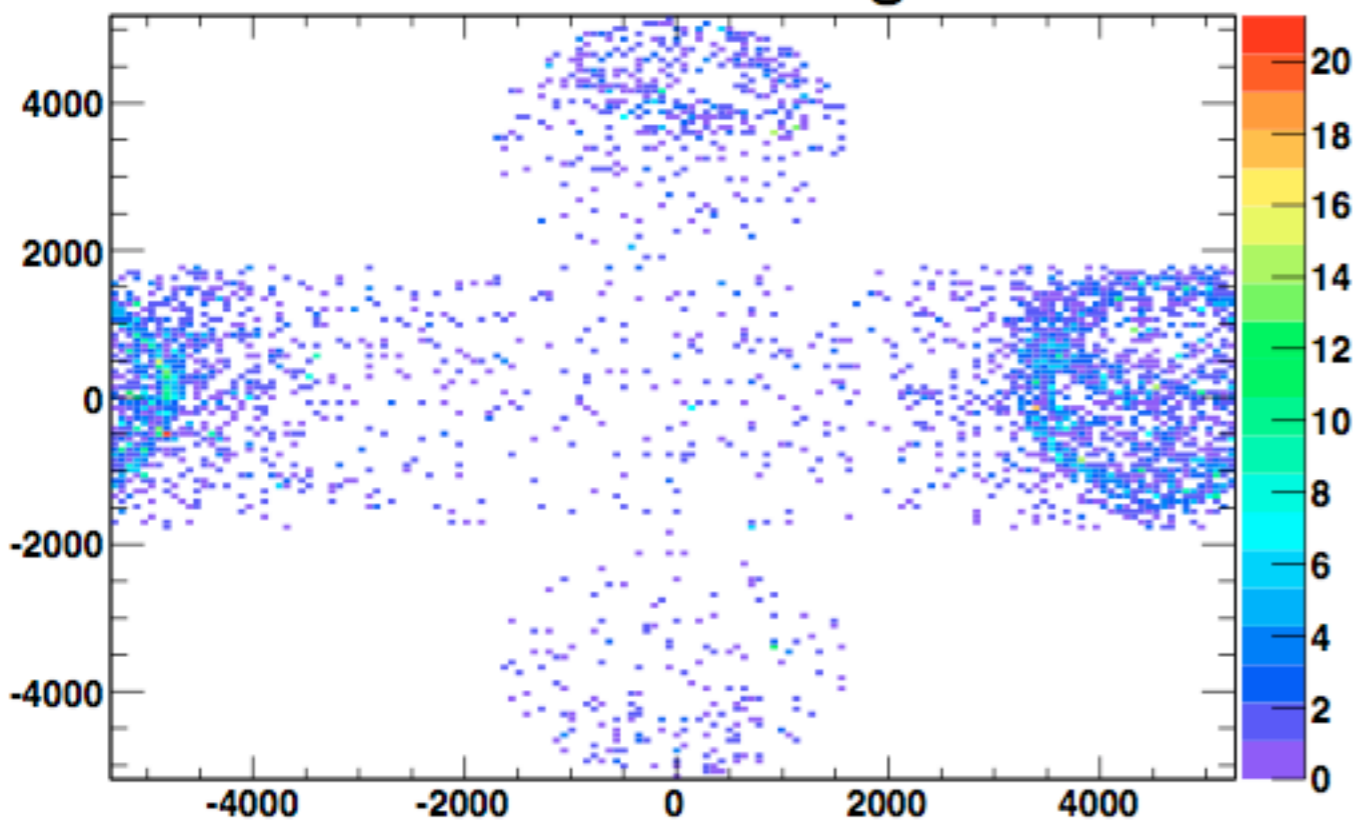
- Also effective to suppress CC $\pi^+$  in  $\nu_\mu$  disappearance

Event category	$\sin^2 2\theta_{13} = 0.0$	$\sin^2 2\theta_{13} = 0.1$
Total	$3.22 \pm 0.43$	$10.71 \pm 1.10$
$\nu_e$ signal	0.18	7.79
$\nu_e$ background	1.67	1.56
$\nu_\mu$ background (mainly NC $\pi^0$ )	1.21	1.21
$\bar{\nu}_\mu + \bar{\nu}_e$ background	0.16	0.16

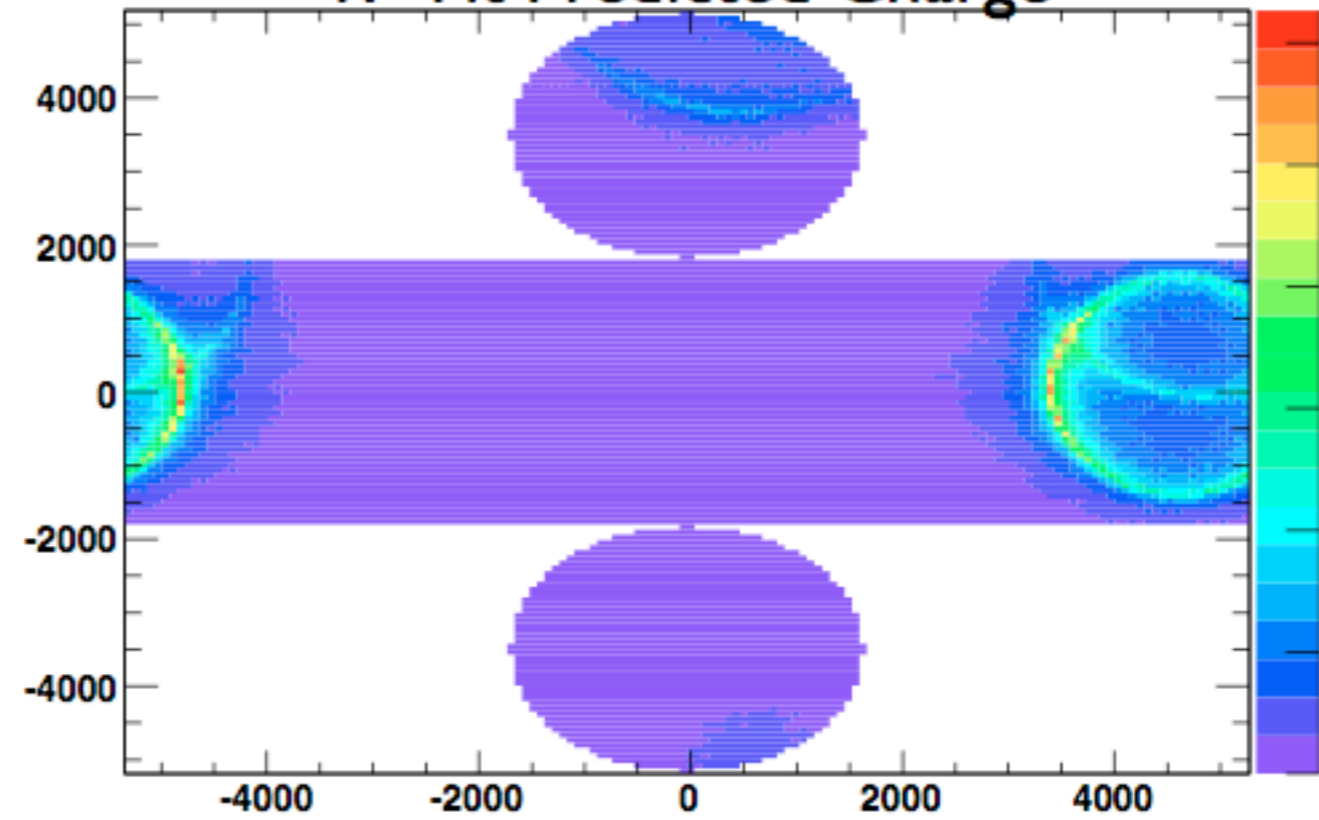


# fitQun $\pi^0$ fitter

Measured Charge



$\pi^0$  Fit Predicted Charge



# Potential improvements

- Background suppression & measurement (fitQun)
  - ve appearance ( $\theta_{13}$ ,  $\delta_{cp}$ ):  $NC\pi^0$  [down to <30%]
- Reduction of systematic uncertainties
  - Background suppression & measurement
  - Cross section modeling and measurements
- Statistics to improve:
  - T2K: runs with continuously increasing intensity
  - Hyper-K: 25 time more fiducial volume

# Atmospheric neutrino

mass hierarchy

$\theta_{23} < 45^\circ$  or  $\theta_{23} > 45^\circ$

# Atmospheric neutrino

$$\frac{\Phi(\nu_e)}{\Phi_0(\nu_e)} - 1 \approx P_2 \cdot (r \cdot \cos^2 \theta_{23} - 1) \quad \theta_{12}$$

$$-r \cdot \sin \tilde{\theta}_{13} \cdot \cos^2 \tilde{\theta}_{13} \cdot \sin 2\theta_{23} \cdot (\cos \delta \cdot R_2 - \sin \delta \cdot I_2) \quad \text{interference } (\delta_{cp})$$

$$+2 \sin^2 \tilde{\theta}_{13} \cdot (r \cdot \sin^2 \theta_{23} - 1) \quad \theta_{13} \text{ resonance}$$

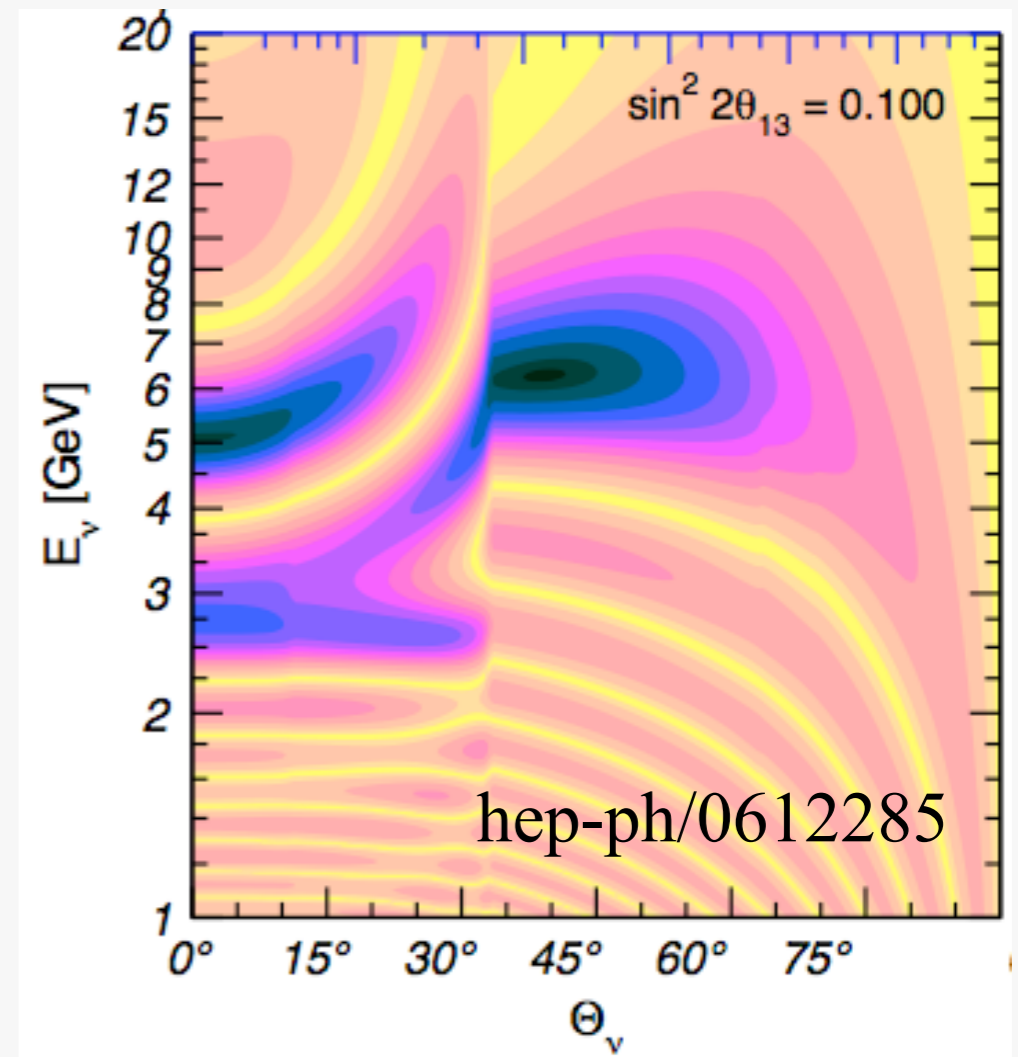
- Resonant oscillation (oscillogram)

- Zenith angle dependence

- 0-33°: core+mantle
- 33°-75°: mantle

- $r = \Phi_\mu / \Phi_e$

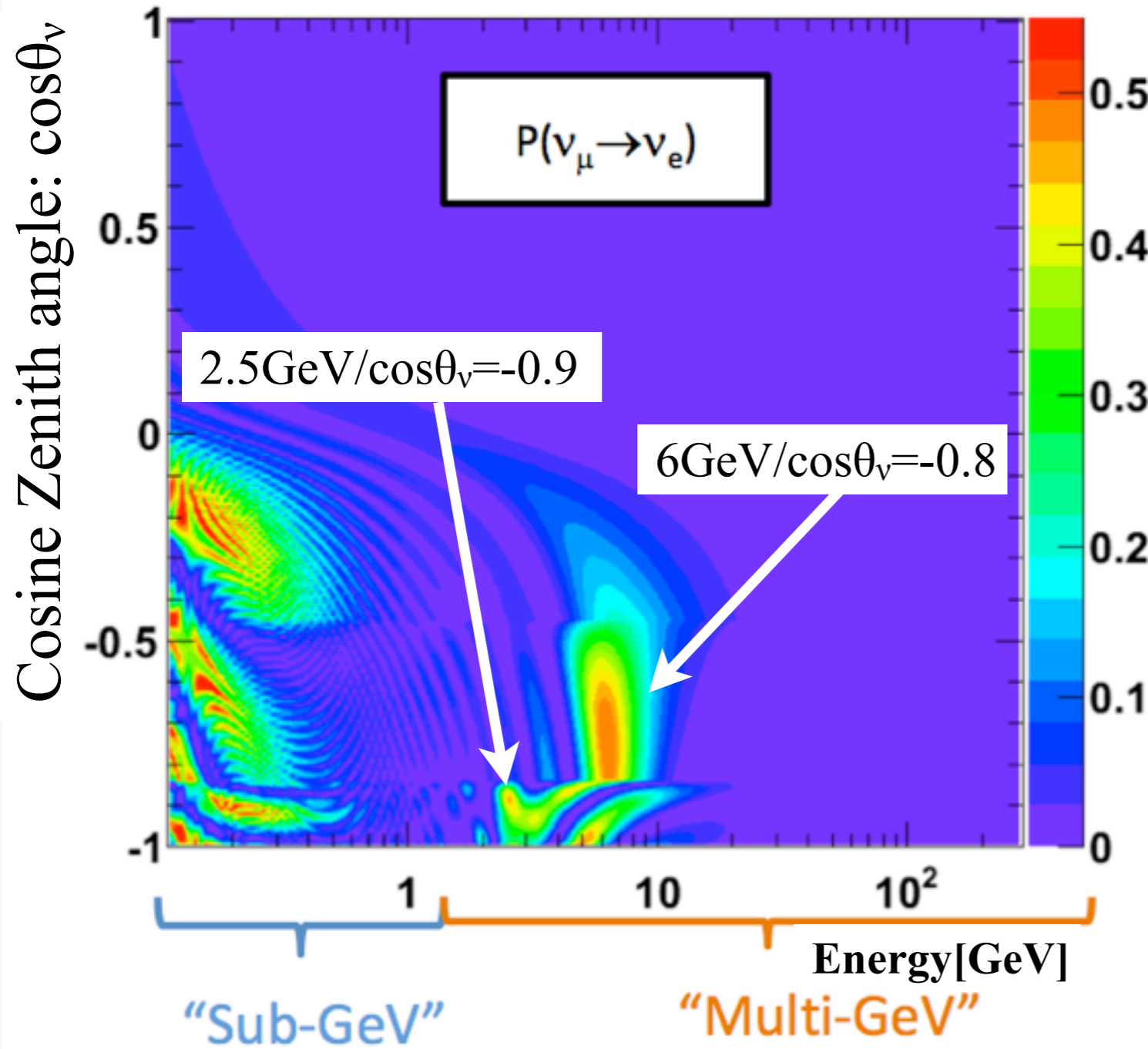
- $r = 2.1 @ 1\text{GeV}, 2.4 @ 3\text{GeV}, 2.6 @ 6\text{GeV}$
- cancellation for  $\sin^2 \theta_{23} < 0.5$





# Resonant $\nu_\mu \rightarrow \nu_e$ oscillation

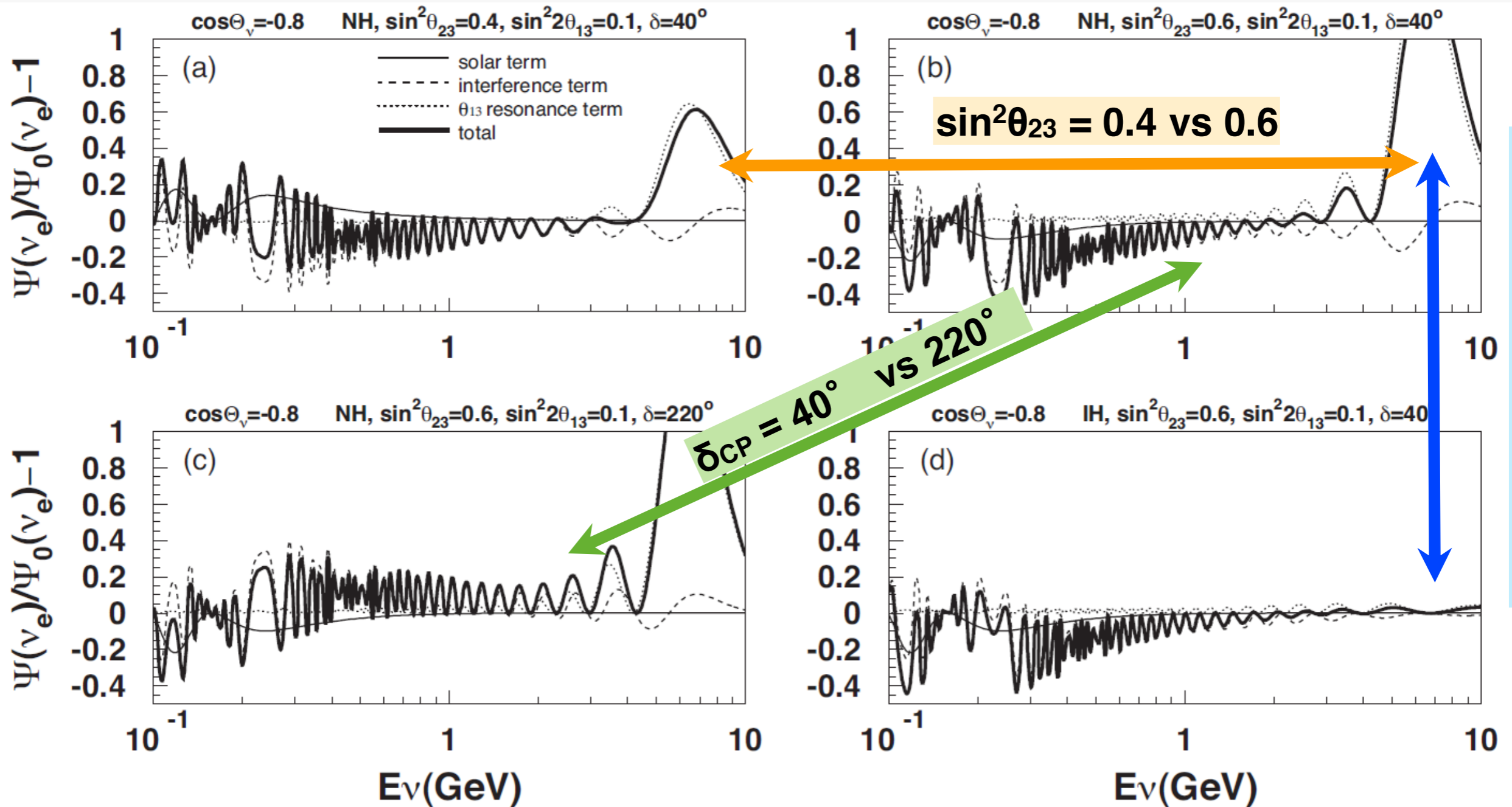
R. Wendel@HyperK meeting



- $\nu_\mu \rightarrow \nu_e$  resonance
  - $2.5 \text{ GeV} / \cos\theta_\nu = -0.9$
  - $6 \text{ GeV} / \cos\theta_\nu = -0.8$
- Resonance depends on
  - $\cos\theta_\nu$  and  $E_\nu$ 
    - multi-GeV, upward
  - $\nu / \bar{\nu}$ , mass hierarchy
    - $\nu$  for normal hierarchy
    - $\bar{\nu}$  for inverted hierarchy

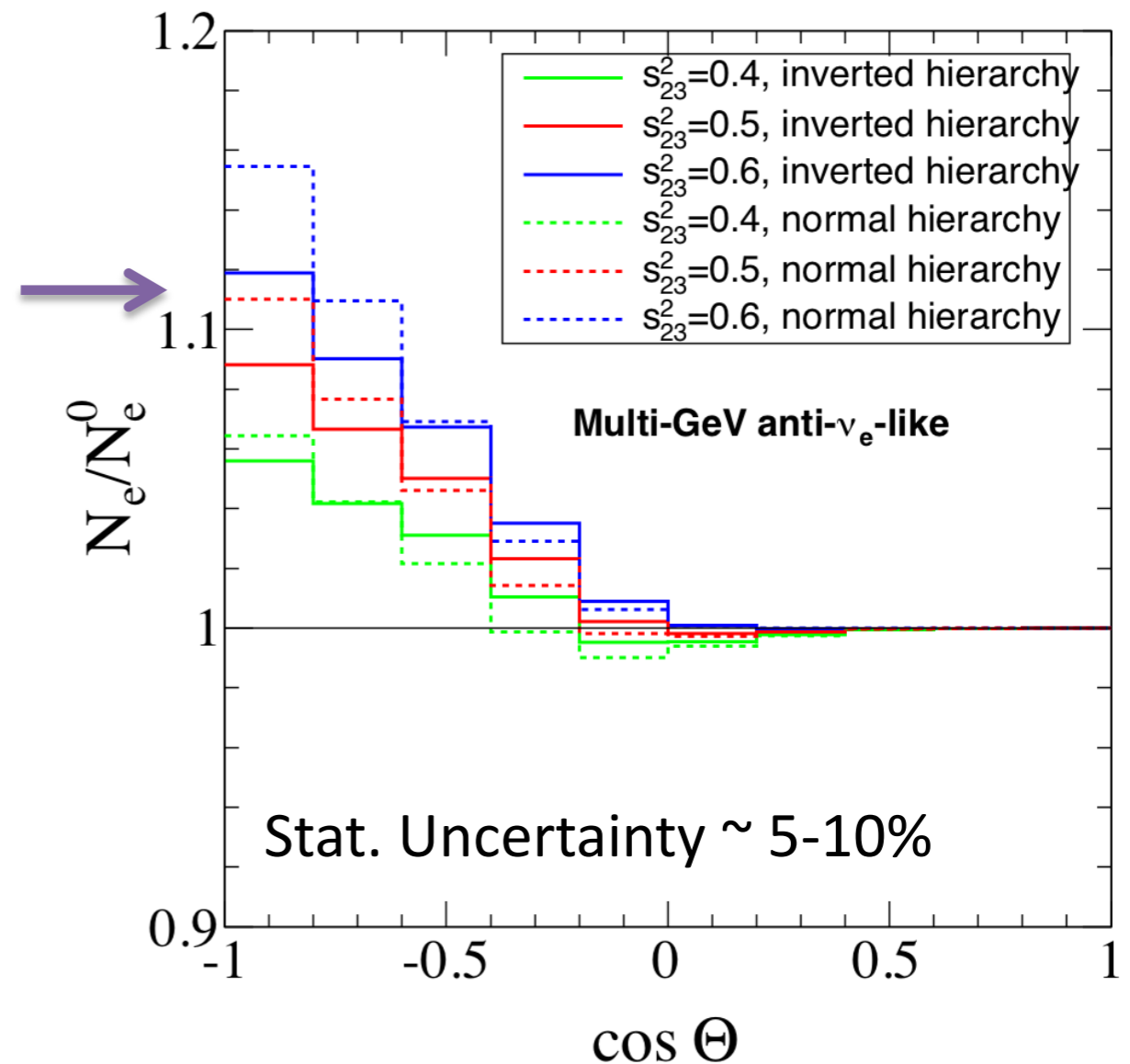
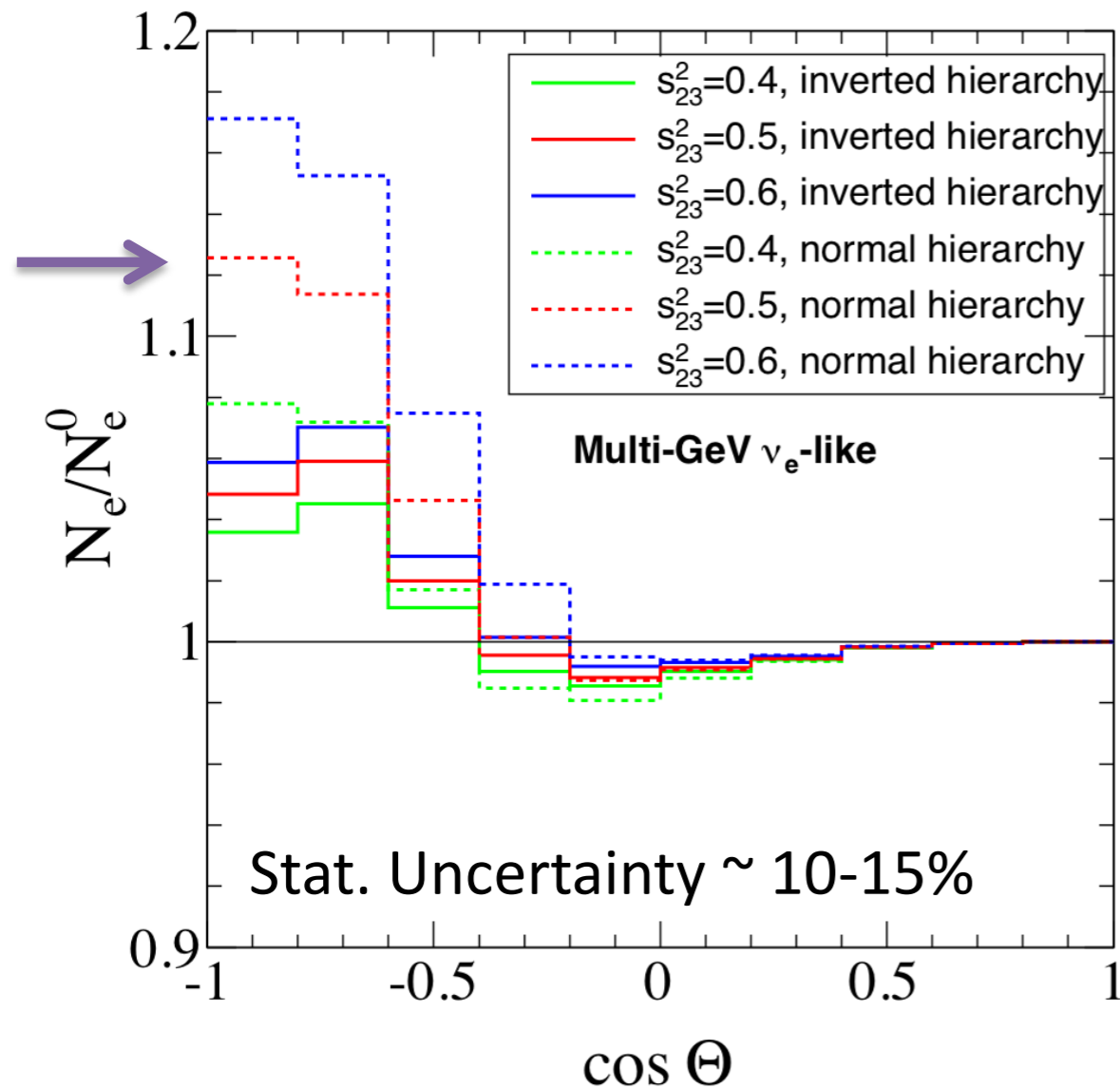
# atmospheric $\nu_e$ appearance

$\cos\theta_v = -0.8$



Sensitive to mass hierarchy,  $\theta_{23}$  octant,  $\delta_{CP}$

# Expected Effects : Single-Ring e-like and anti-e-like samples



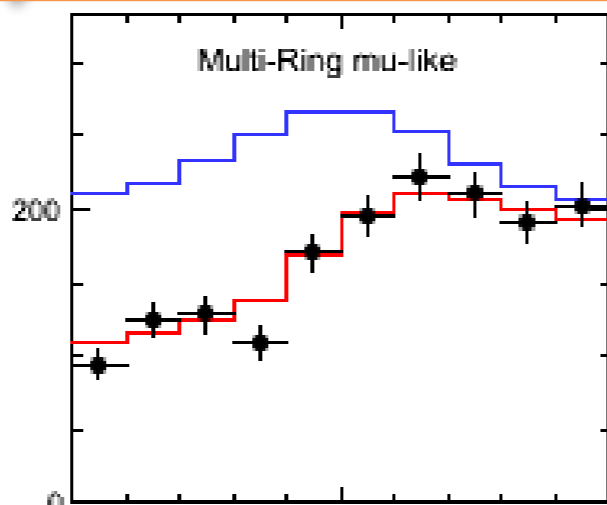
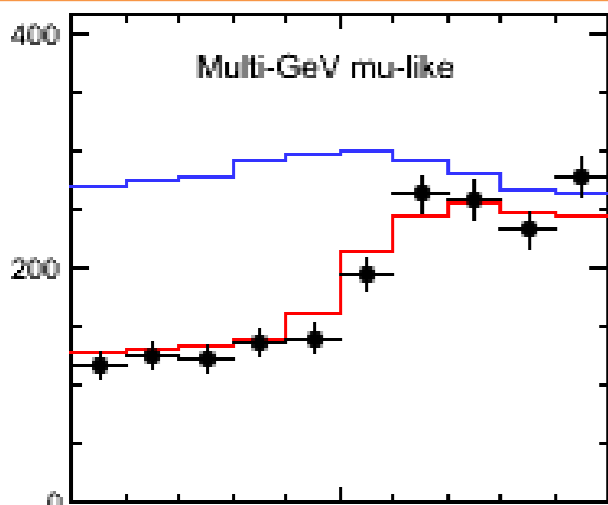
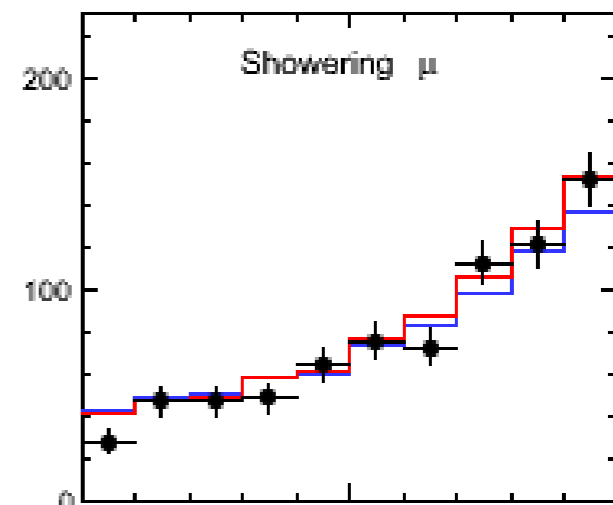
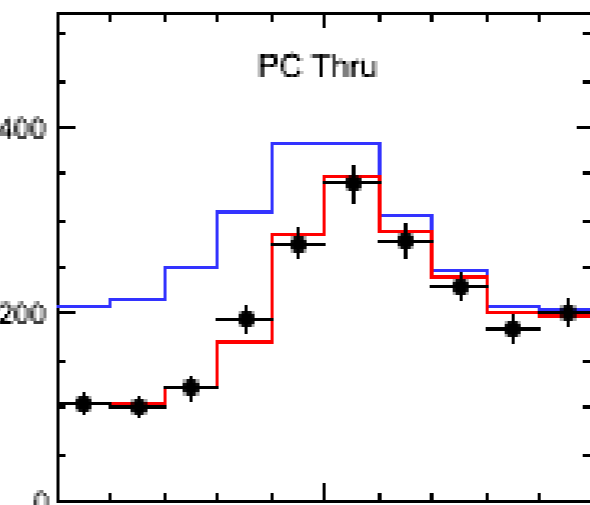
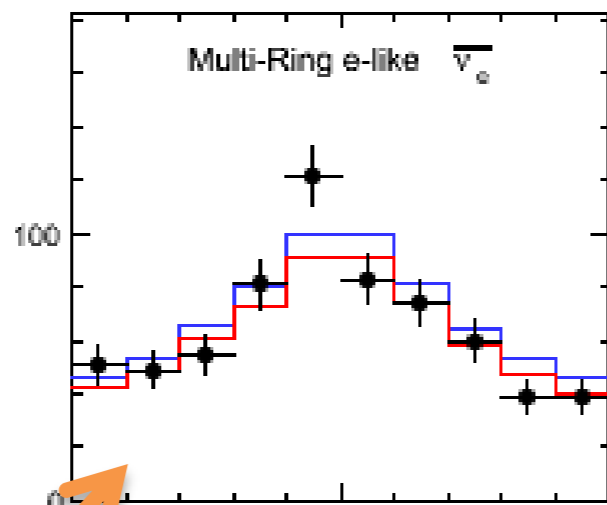
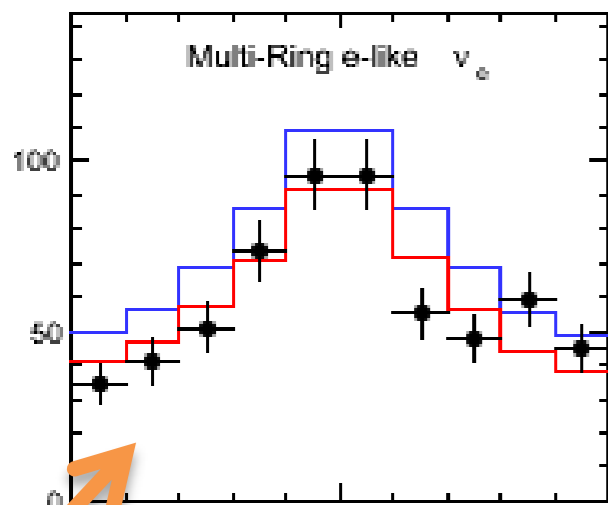
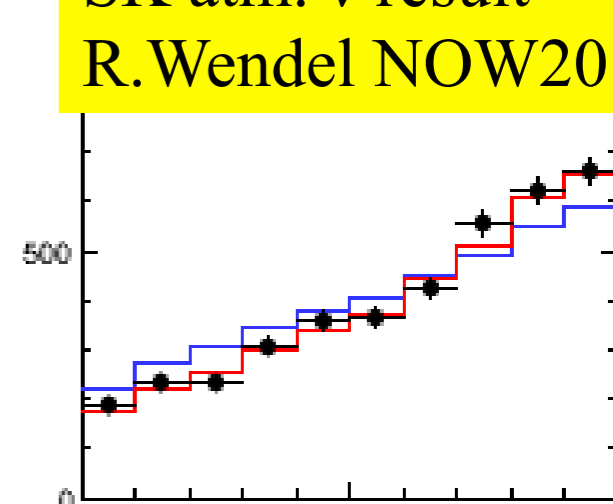
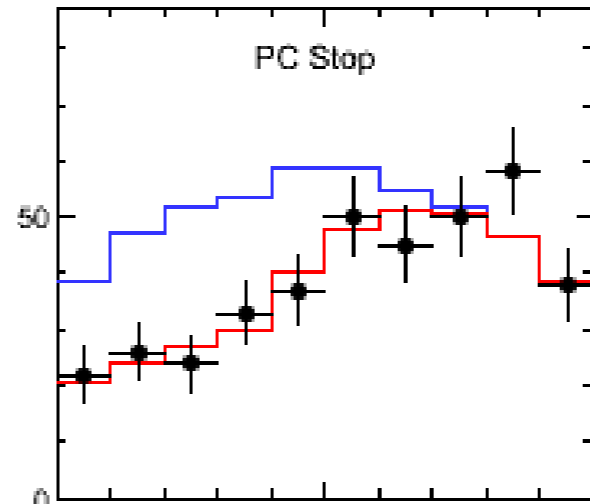
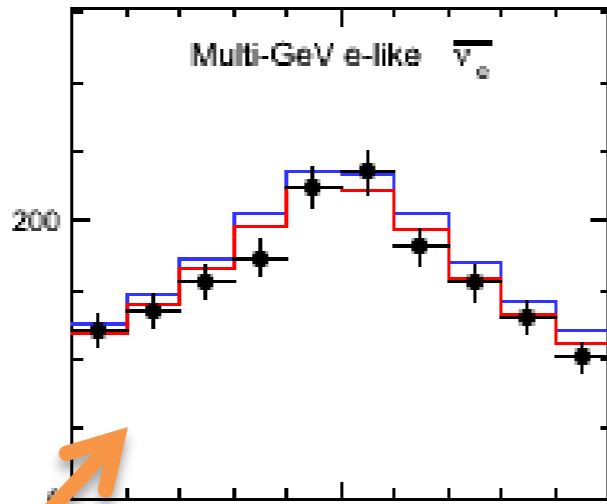
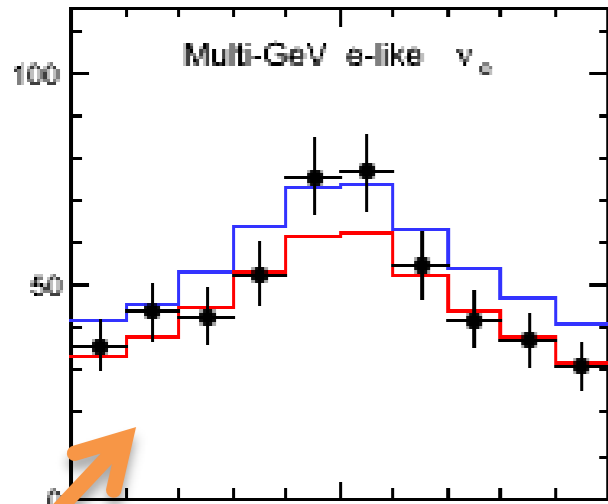
→ Equivalent MC

$\nu_e$ -like:  $CC1\pi^+ \Leftrightarrow$  with decay electron

$\bar{\nu}_e$ -like  $CC1\pi^- \Leftrightarrow$  without decay electron

R. Wendel NOW2012

SK atm.  $\nu$  result  
R. Wendel NOW2012

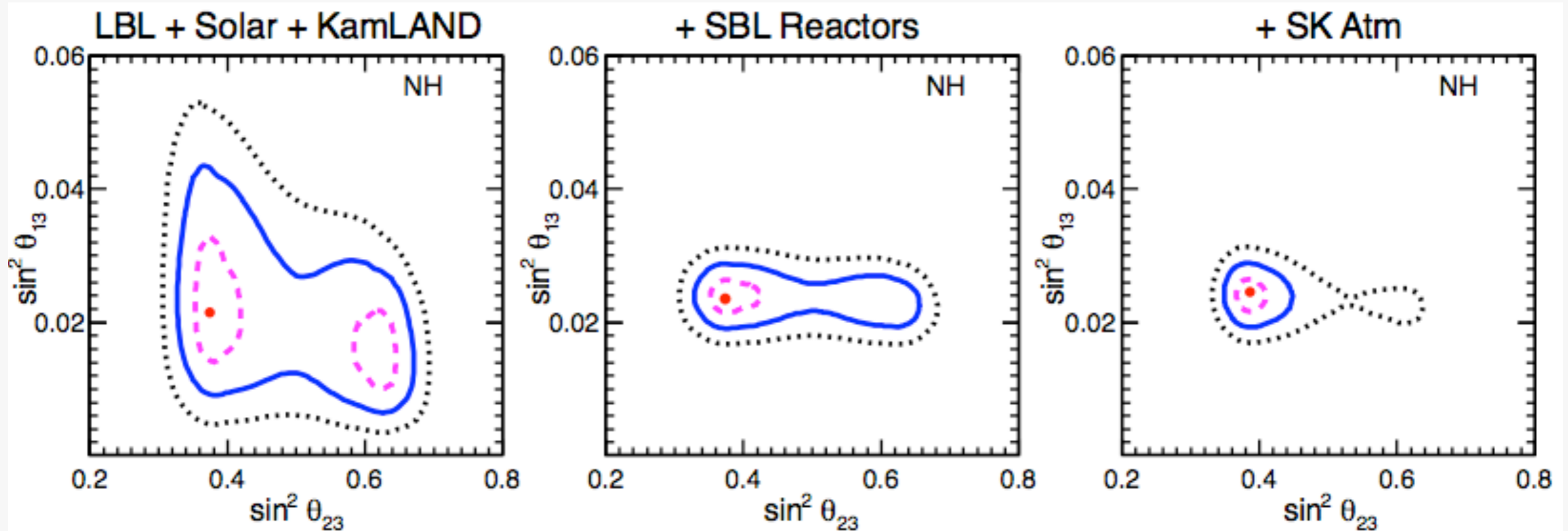


- Data
- Unoscillated Expectation
- Oscillated Expectation
- ➔ Electron appearance region



SK I-IV: 3903 days, 34,000 events

# Systematics in Atm. $\nu$



In practice, it is difficult to infer —from atmospheric data— clean  $3\nu$  information beyond the dominant parameters ( $\Delta m^2$ ,  $\theta_{23}$ ). Subdominant oscillation effects are often smeared out over wide energy-angle spectra of events, and can be partly mimicked by systematic effects. For this reason, “hints” coming from current atmospheric data should be taken with a grain of salt, and should be possibly supported by independent datasets. **Fogli et. al. ArXiv:1205.5254v3**

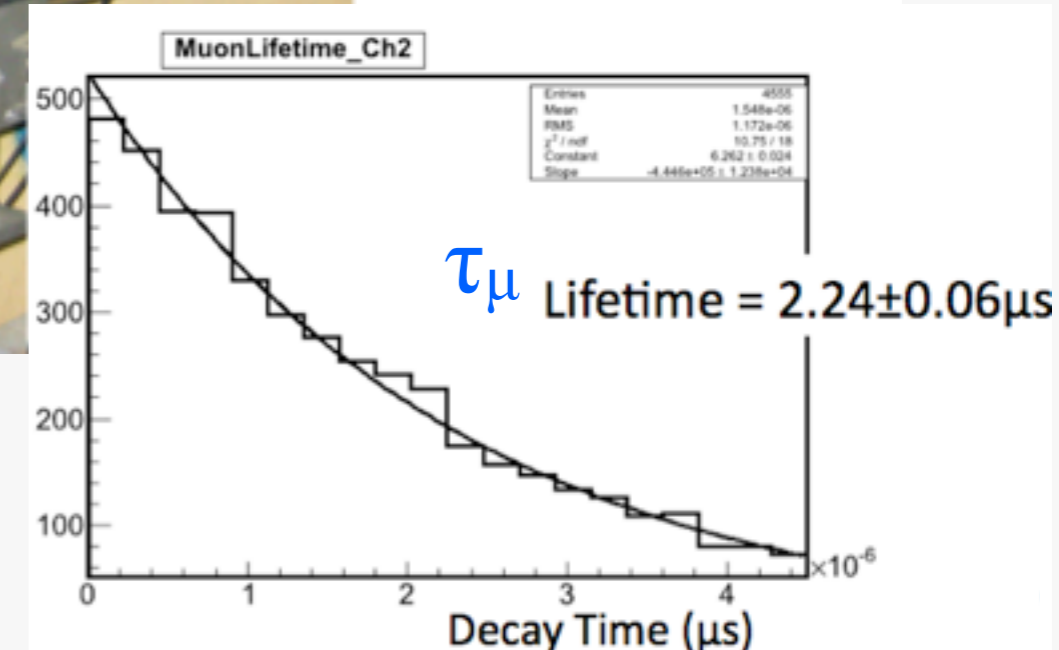
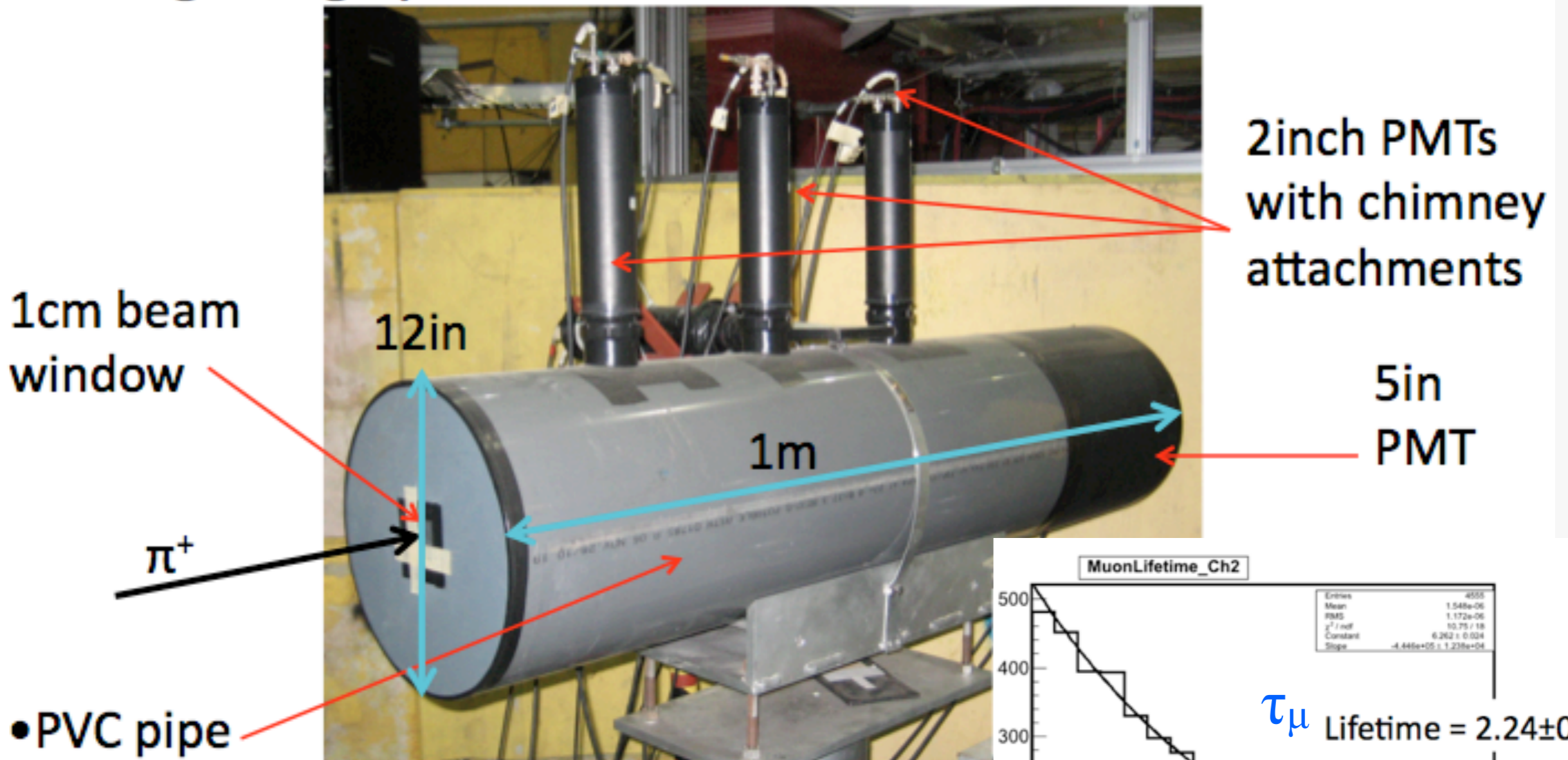
# Potential Improvement: atm. $\nu$ CC1 $\pi^\pm$

- Exclusive CC1 $\pi^\pm$  reconst. successfully done @ MiniBooNE
  - FitQun reconstruction is based on the same method.
- Energy and zenith angle reconstruction
  - Pion Cerenkov ring before interaction  $\rightarrow P_\pi$
  - Full reconstruction except for the recoil nucleon:
    - $\Delta E_\nu \sim 50 \text{ MeV} : \Delta E_\nu / E_\nu = 50 / 2500 \sim 2\%$
    - $P_T(\text{nucleon}) \sim 300 \text{ MeV}/c : \Delta \theta_\nu / \theta_\nu \sim 300 / 2500 \sim 12\%$
- Background suppression: clean measurement
  - Particle identification
    - $\nu_e$ : CC $\pi^+$  with decay  $e$ , anti- $\nu_e$ : CC $\pi^-$  without decay  $e$
  - NC( $\pi^0$ ) rejection using fitQun reconstruction
- Spectrum before oscillation given by down-ward going  $\nu$ 's

# TRIUMF $\pi$ beam test (M11)

S.Berkman et.al.

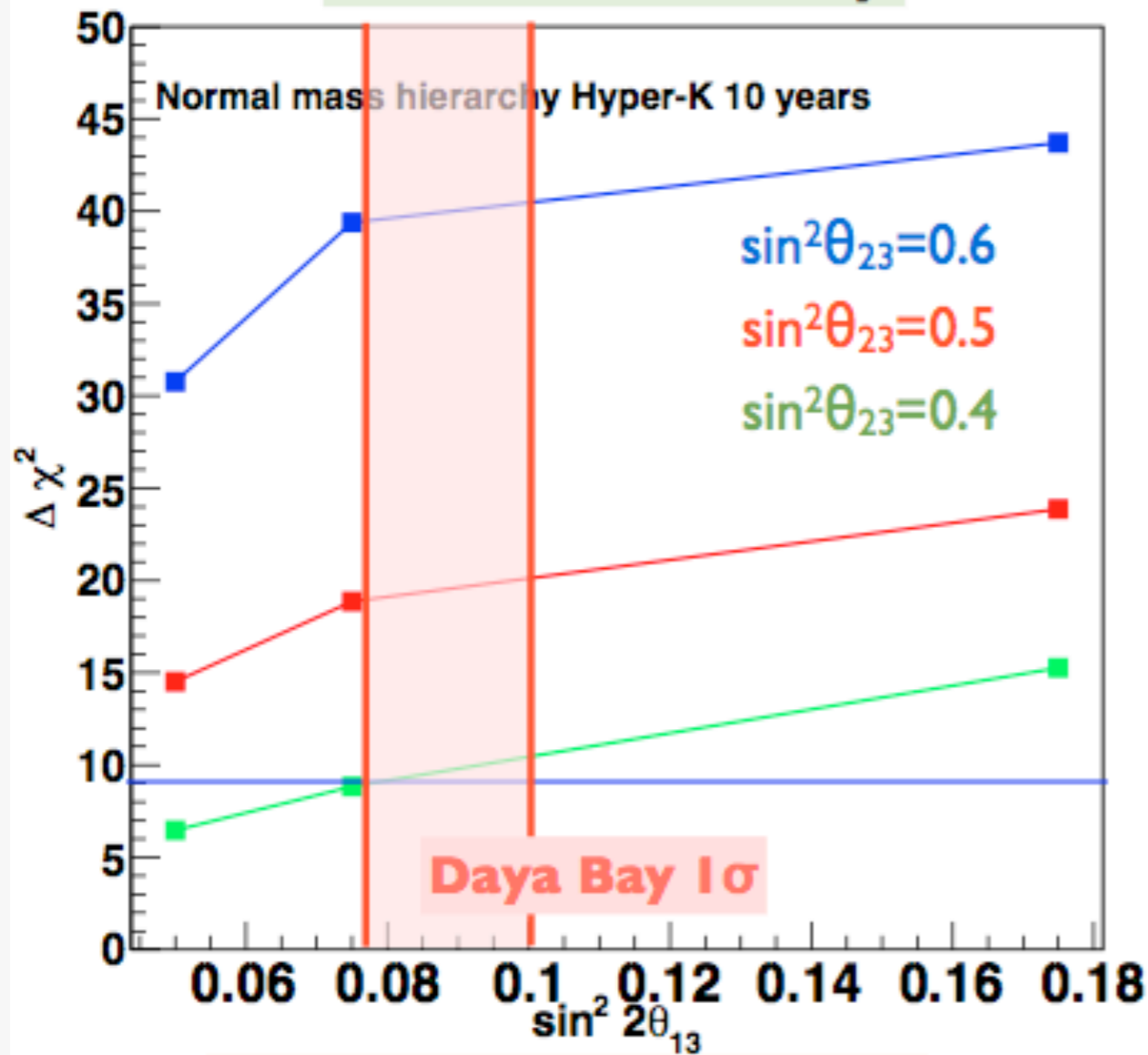
- Integrating cylinder filled with water



Measured probability for  $\pi$  to stop and decay in water for  $P_\pi=0.1-0.3\text{GeV}/c$

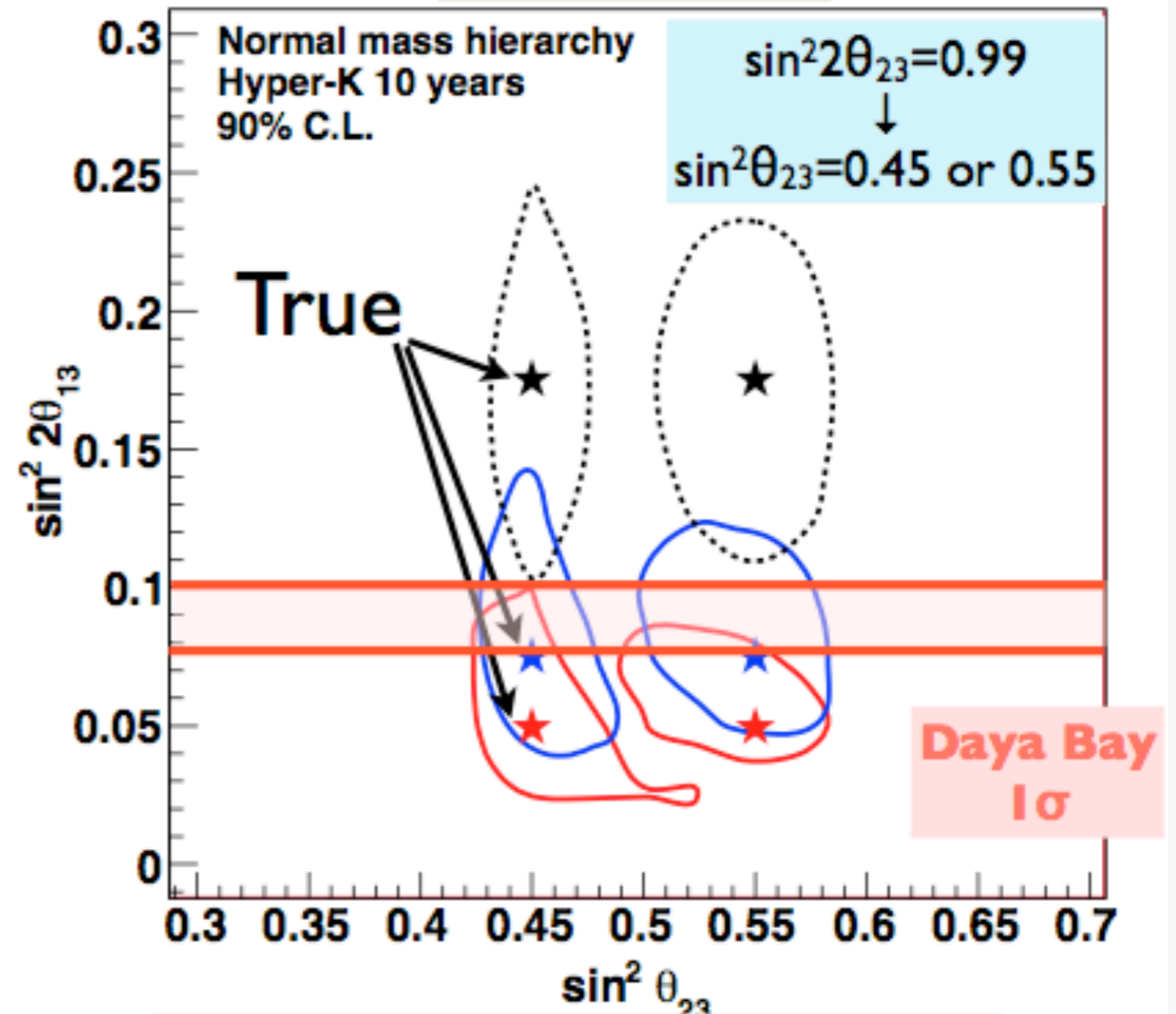
# Atmospheric $\nu$ at Hyper-K

## Mass hierarchy



>3 $\sigma$  with 5-10 years

## $\theta_{23}$ octant



Resolved if  $\sin^2 2\theta_{23} < 0.99$

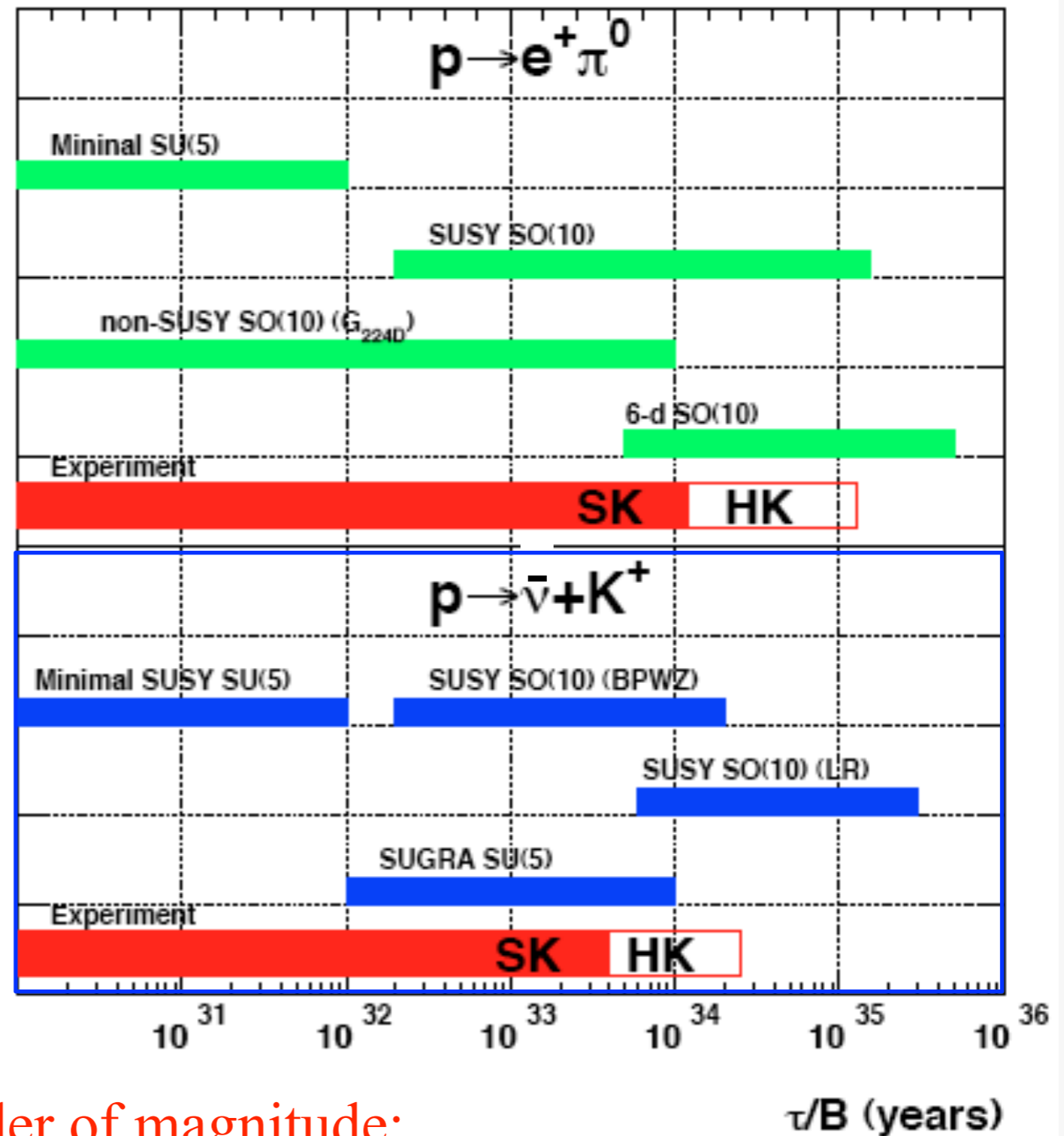
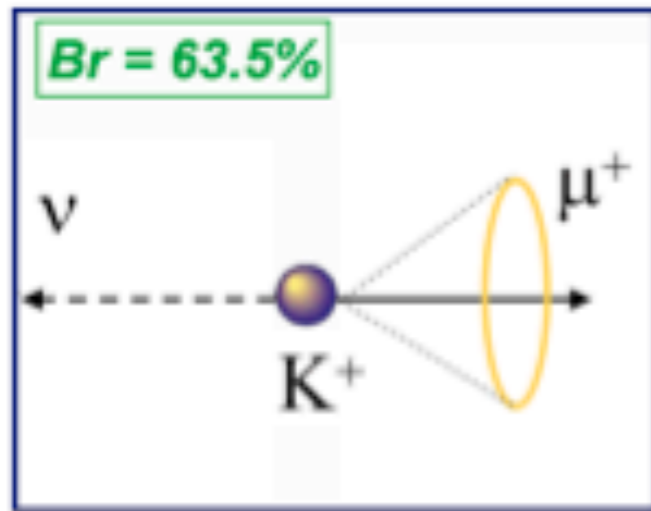
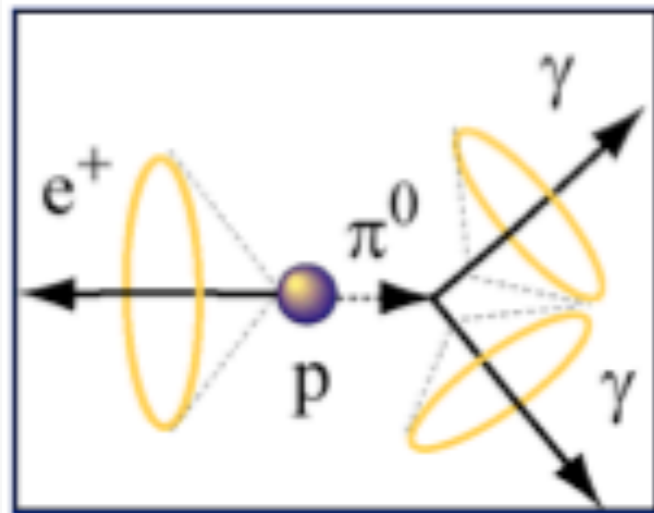
Without assuming fitQun but 5% syst.

M.Yokoyama@Neutrino 2012



# Proton decays

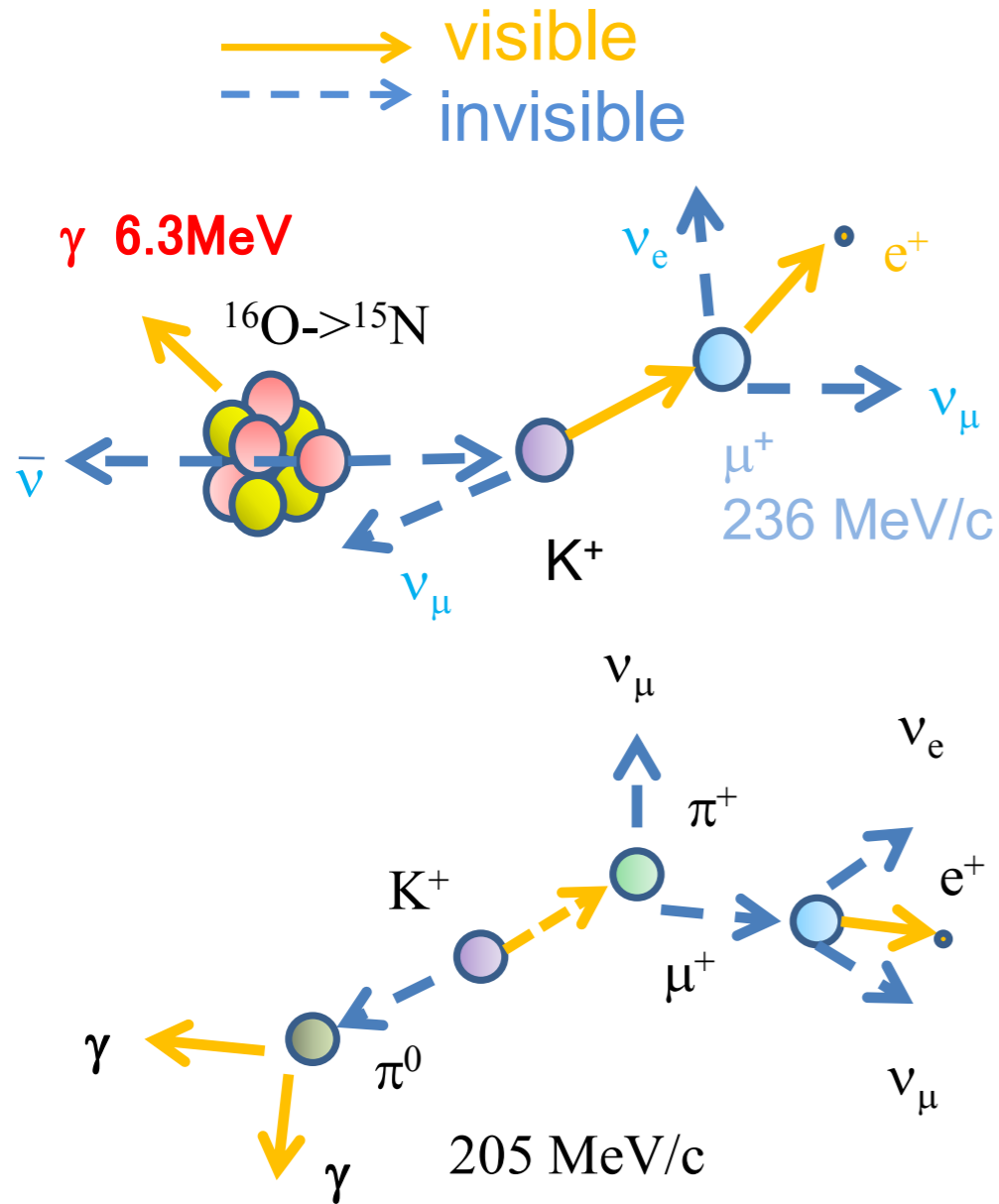
# Proton decays



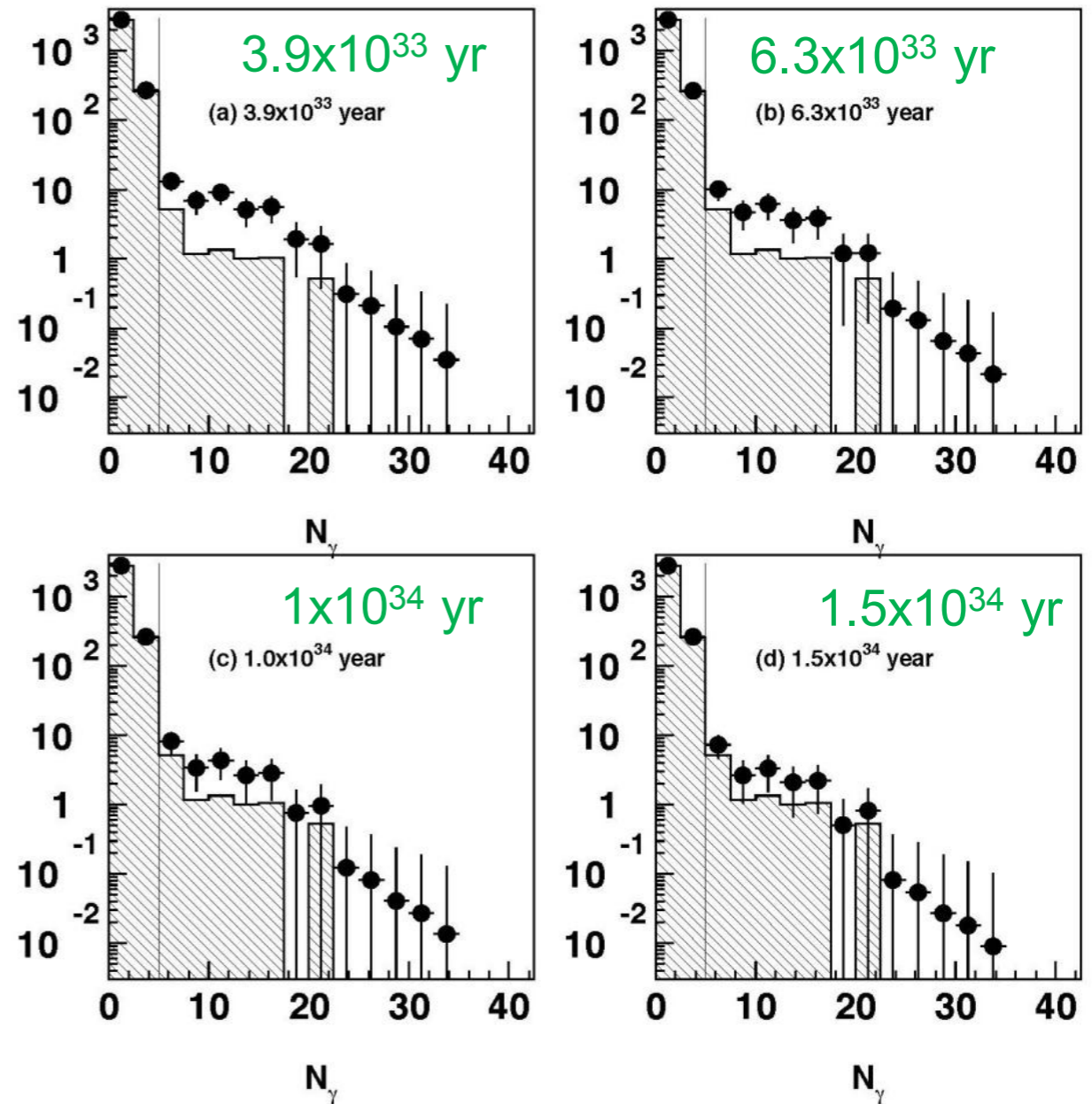
Expand search by another order of magnitude:  
Getting deep into the remaining GUT window

# $p \rightarrow \nu K^+$ mode at Hyper-K

Nakahata@Lyon



Number of hits of the prompt gamma signal (10 years)

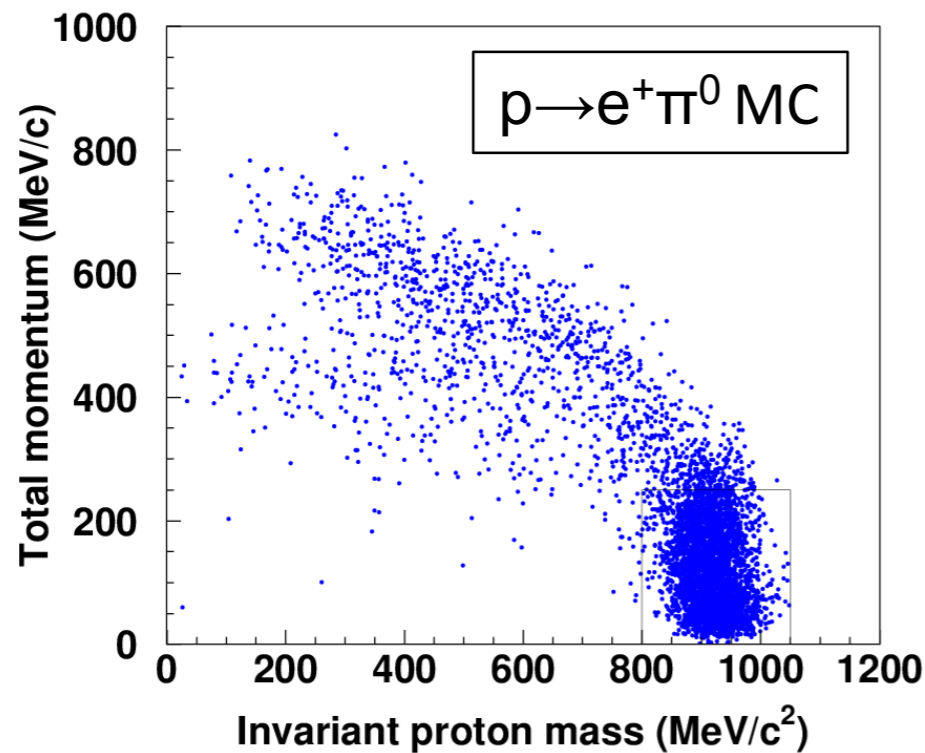
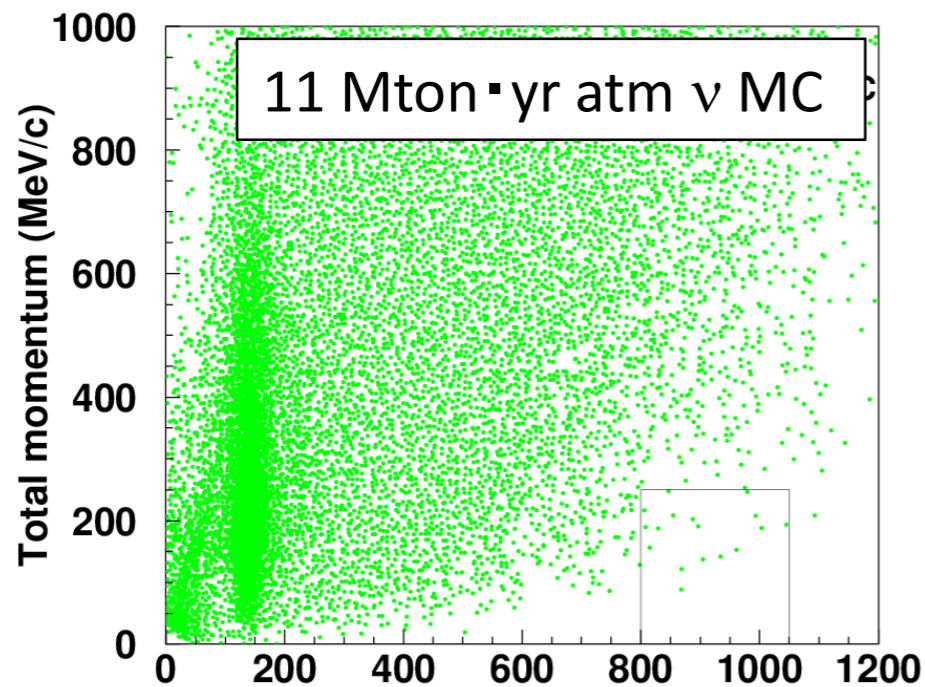


Discovery potential ( $3\sigma$ ):  $1.0 \times 10^{34}$  years

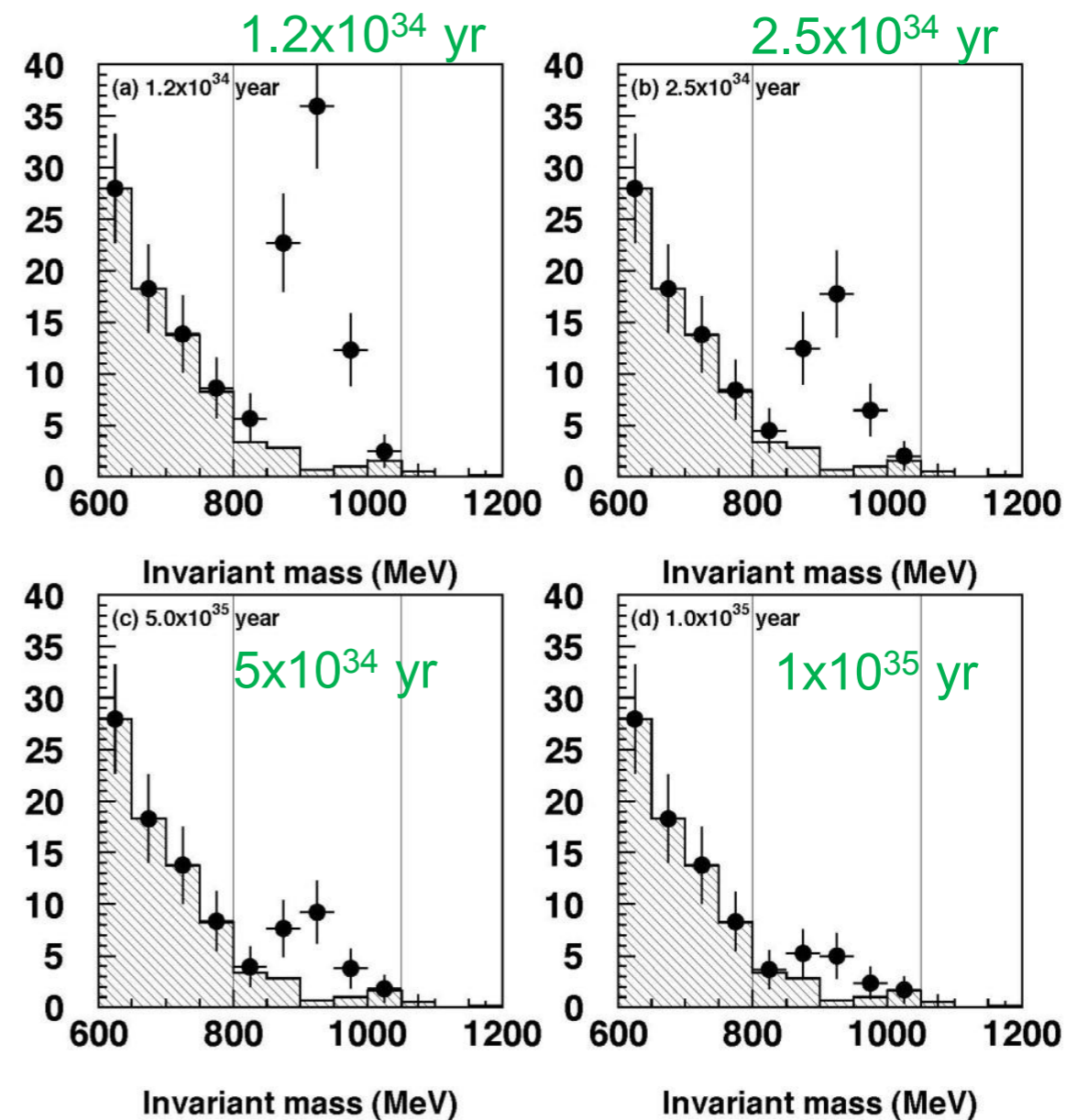
Sensitivity (90% C.L.):  $2.5 \times 10^{34}$  years

with 10 years run (5.6 Mton  $\cdot$  yrs)

# $p \rightarrow e^+ + \pi^0$ mode at Hyper-K



Reconstructed proton mass(10 years)  
(Total  $p < 250$  MeV/c)



Discovery potential ( $3\sigma$ ):  $5.7 \times 10^{34}$  years

Sensitivity (90%C.L.):  $1.3 \times 10^{35}$  years

with 10 years run (5.6 Mton·yrs)

# Potential improvements

- $p \rightarrow K\nu$ 
  - Improved 6MeV  $\gamma$  tagging efficiency with fitQun
    - detect low energy  $\gamma$  before K decay
    - good K decay vertex reconstruction to reject fake  $\gamma$
    - Full reconstruction of  $K^+ \rightarrow \pi^+ \pi^0$
- $p \rightarrow e\pi^0$ 
  - Background rejection
    - better energy/momentum resolution with fitQun
- Gd to veto associated neutrons from atm. $\nu$  BG?

# Summary

- Precision  $\nu$  oscillation measurements provides clues on origin of the flavour (family) structure.
- T2K continue accumulating data with increasing  $\nu$  flux and future proposal of Hyper-K prepared.
  - $\theta_{23}$ : long baseline  $\nu_{\mu}$  disappearance
  - $\delta_{cp}$ : long baseline  $\nu_e$  appearance
  - mass hierarchy/ $\theta_{23}$  octant: atmospheric  $\nu$
  - proton decays: explore GUT

# Summary (cont.)

- Controlling systematics is crucial:
  - Progress in cross section modeling/measurements
  - New reconstruction code
- New reconstruction code being prepared: fitQun
  - better reconstruction
    - momentum, vertex, particle identification
  - background tagging and rejection
    - $\pi^0$  for accelerator  $\nu_e$  appearance (to less than 1/3)
    - $\pi^\pm$  for accelerator  $\nu_\mu$  disappearance
    - $\pi^\pm$  for atm.  $\nu_e$  appearance: tag CC1 $\pi^+$  and CC1 $\pi^-$

# Backups



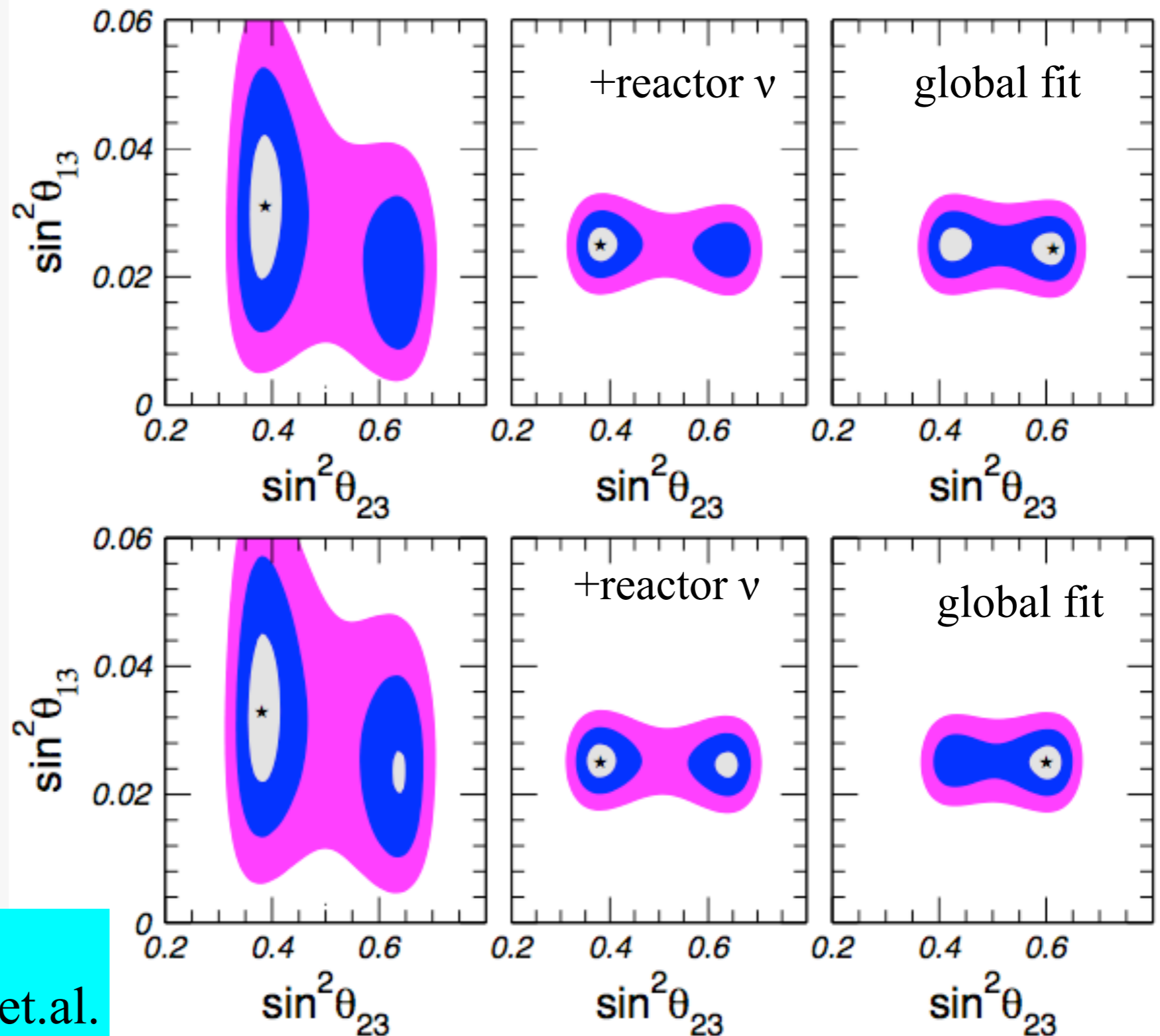
# Global fit (2012)

Normal hierarchy

arXiv:1205.4218  
(Valencia group)

Inverted hierarchy

$\theta_{23} > 45^\circ$  preferred:  
different from Fogli et.al.

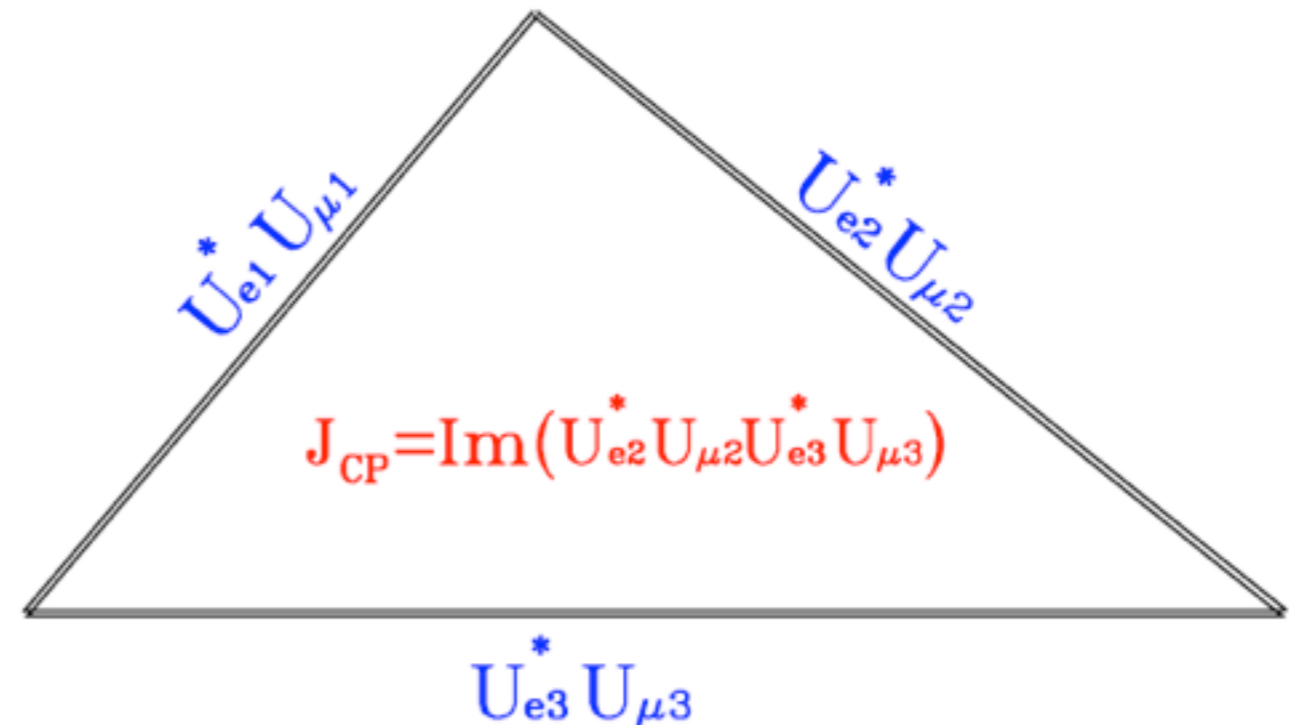


# Lepton Unitary Triangle

- Large  $\theta_{13}$  makes unitary triangle measurement realistic now: explore physics beyond SM
  - Sterile neutrino (right-handed  $\nu$ ) breaks unitarity
  - 4th generation, new interaction, etc.
  - Two speakers independently pointed out at Neutrino2012

- 4 coefficients overconstrain the unitarity triangle:

$$U_{e1}^* U_{\mu 1} + U_{e2}^* U_{\mu 2} + U_{e3}^* U_{\mu 3} = 0$$



# Mass hierarchy

- Matter resonance: atmospheric  $\nu$ 
  - SuperK/HyperK, INO, PINGU, ...
- Long baseline  $\nu$ 
  - Comparison between short and long baselines
    - T2K/HK vs. NOVA/LBNE
- Comparison between  $\Delta m^2_{32}$  and  $\Delta m^2_{13}$ 
  - Reactor  $\Delta m^2_{13}$  vs. T2K  $\Delta m^2_{32}$
  - Daya-Bay2: Reactor  $\Delta m^2_{13}$  and  $\Delta m^2_{12}$  at 60km
- Cosmology:  $\Sigma m_\nu$
- Double  $\beta$  decays

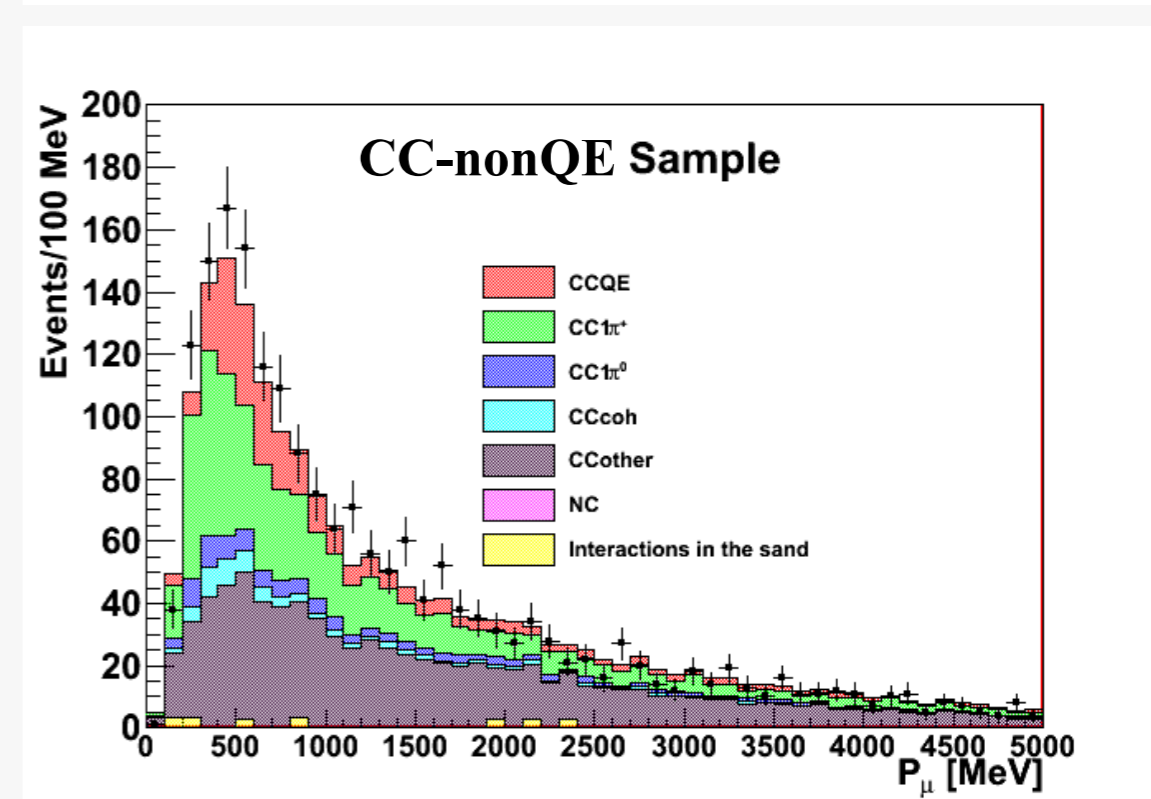
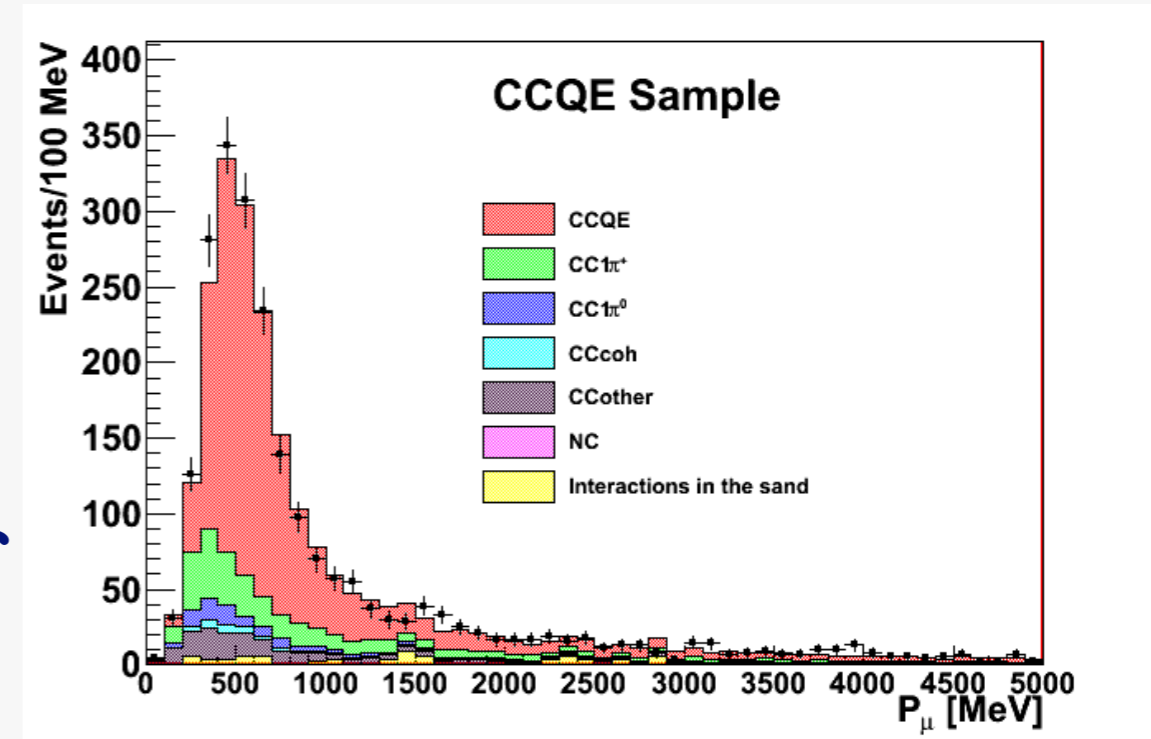
# Cosmological $\nu$ mass

Probe	Current $\sum m_\nu$ (eV)	Forecast $\sum m_\nu$ (eV)	Key Systematics	Current Surveys	Future Surveys
CMB Primordial	1.3	0.6	Recombination	WMAP, Planck	None
CMB Primordial + Distance	0.58	0.35	Distance measurements	WMAP, Planck	None
Lensing of CMB	$\infty$	0.2 – 0.05	NG of Secondary anisotropies	Planck, ACT [39], SPT [96]	EBEX [57], ACTPol, SPTPol, POLAR-BEAR [5], CMBPol [6]
Galaxy Distribution	0.6	0.1	Nonlinearities, Bias	SDSS [58, 59], BOSS [82]	DES [84], BigBOSS [81], DESpec [85], LSST [92], Subaru PFS [97], HETDEX [35]
Lensing of Galaxies	0.6	0.07	Baryons, NL, Photometric redshifts	CFHT-LS [23], COSMOS [50]	DES [84], Hyper SuprimeCam, LSST [92], Euclid [88], WFIRST [100]
Lyman $\alpha$	0.2	0.1	Bias, Metals, QSO continuum	SDSS, BOSS, Keck	BigBOSS [81], TMT [99], GMT [89]
21 cm	$\infty$	0.1 – 0.006	Foregrounds, Astrophysical modeling	GBT [11], LOFAR [91], PAPER [53], GMRT [86]	MWA [93], SKA [95], FFTT [49]
Galaxy Clusters	0.3	0.1	Mass Function, Mass Calibration	SDSS, SPT, ACT, XMM [101] Chandra [83]	DES, eRosita [87], LSST
Core-Collapse Supernovae	$\infty$	$\theta_{13} > 0.001^*$	Emergent $\nu$ spectra	SuperK [98], ICECube [90]	Noble Liquids, Gad-zooks [7]

Normal hierarchy:  $\sum m_\nu > 0.05 \text{ eV}$ , Inverted hierarchy:  $\sum m_\nu > 0.1 \text{ eV}$

# Systematic uncertainties

- Cross section uncertainty dominates:
  - Precise study of cross section in the near detector is critical!

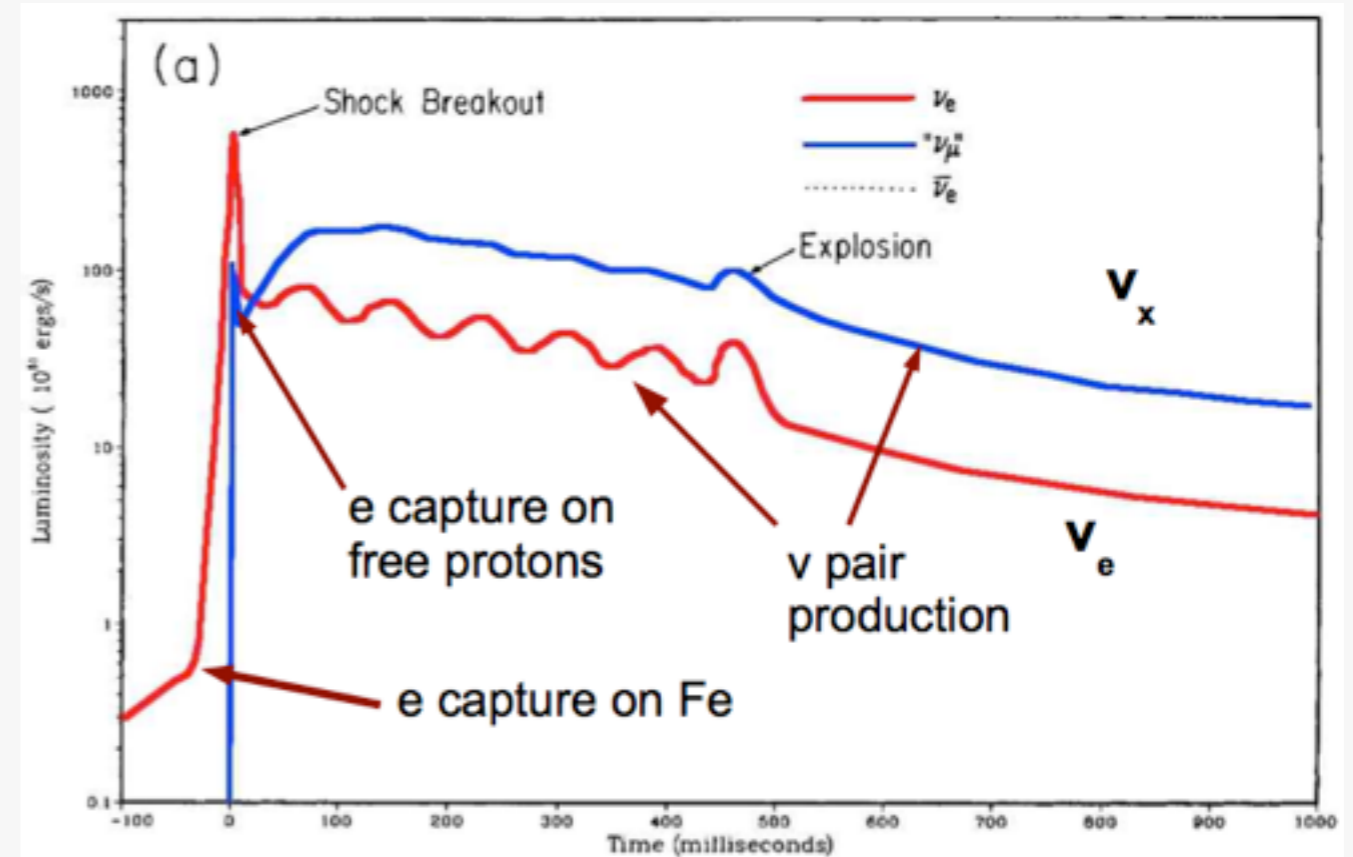
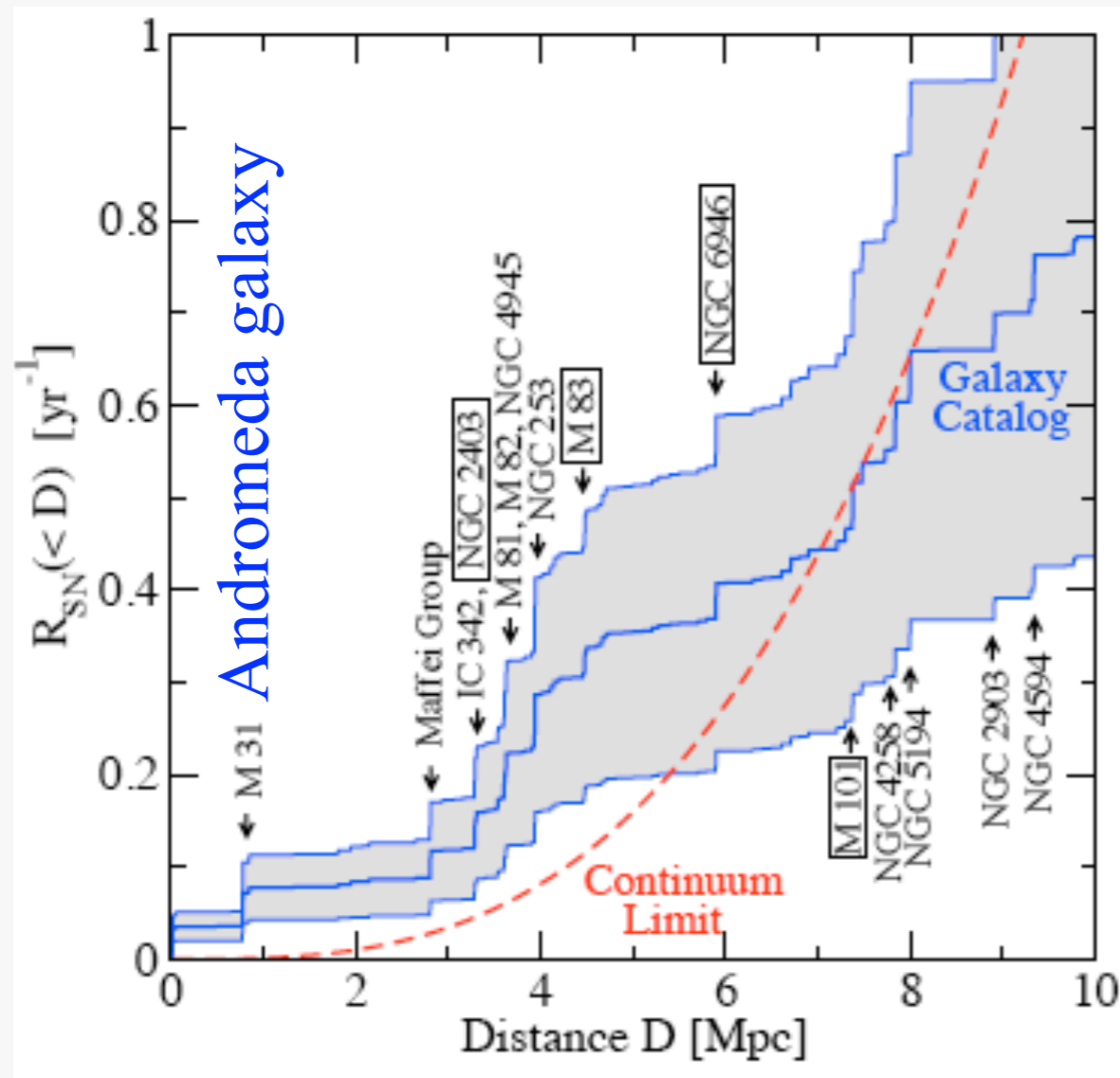


	$\sin^2 2\theta_{13}=0.1$	$\sin^2 2\theta_{13}=0.0$
Flux+Xsec in T2K fit	5.7%	8.7%
Xsec (from other exp.)	7.5%	5.9%
SK + FSI	3.9%	7.7%
<b>Total</b>	<b>10.3%</b>	<b>13.4%</b>

# Other Physics by Hyper-K

- Supernova neutrinos
  - Sensitivity reaches beyond Andromeda Galaxy
  - Supernova Relic Neutrino detection
- Solar neutrinos
  - Day–Night asymmetry measurement
- Indirect WIMP search (solar, galactic)
  - Reaching interesting region in SD interaction
- Others
  - GRB- $\nu$ , solar flare  $\nu$ ,  $n$ - $\bar{n}$  oscil., monopole, Q-ball,...

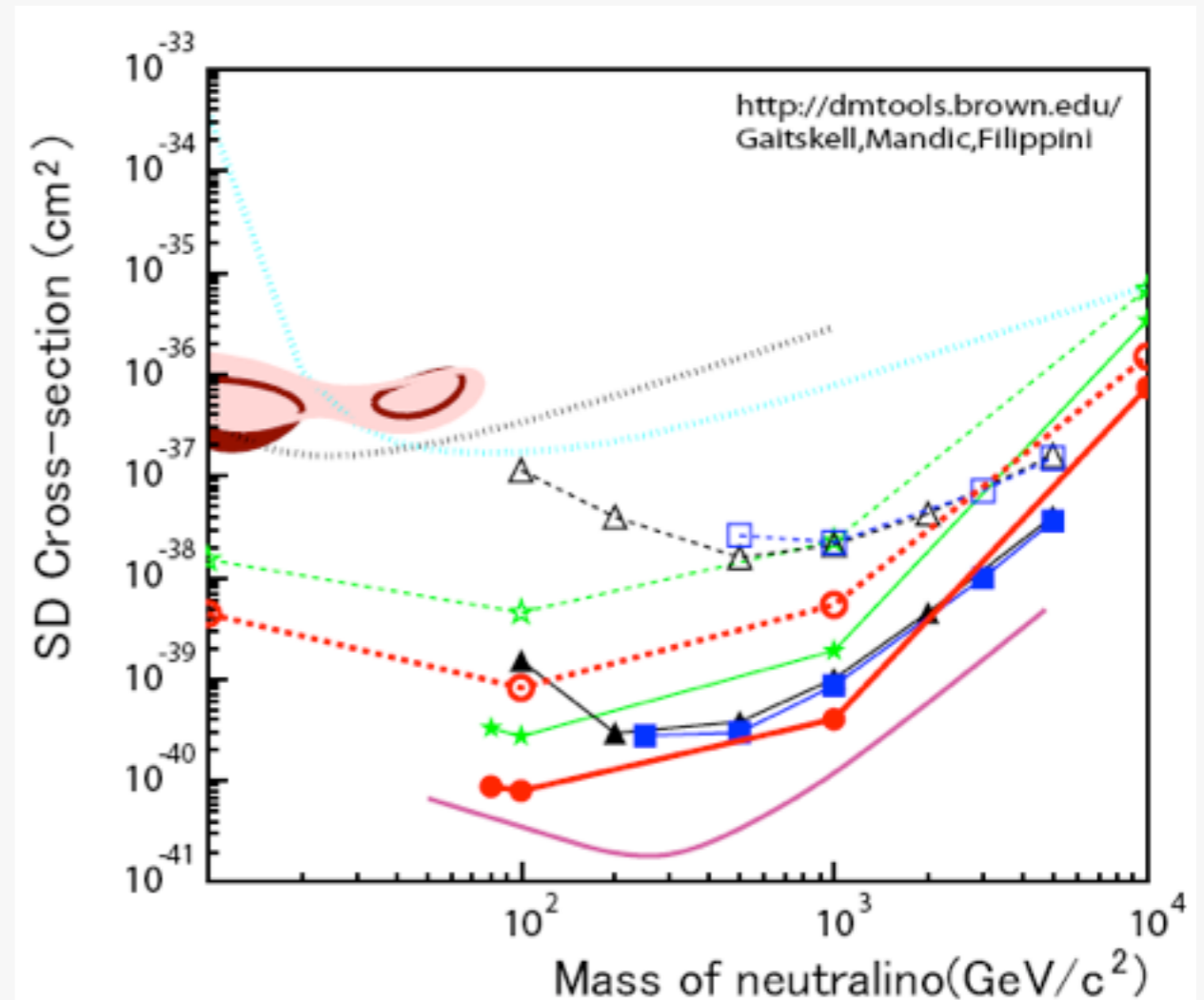
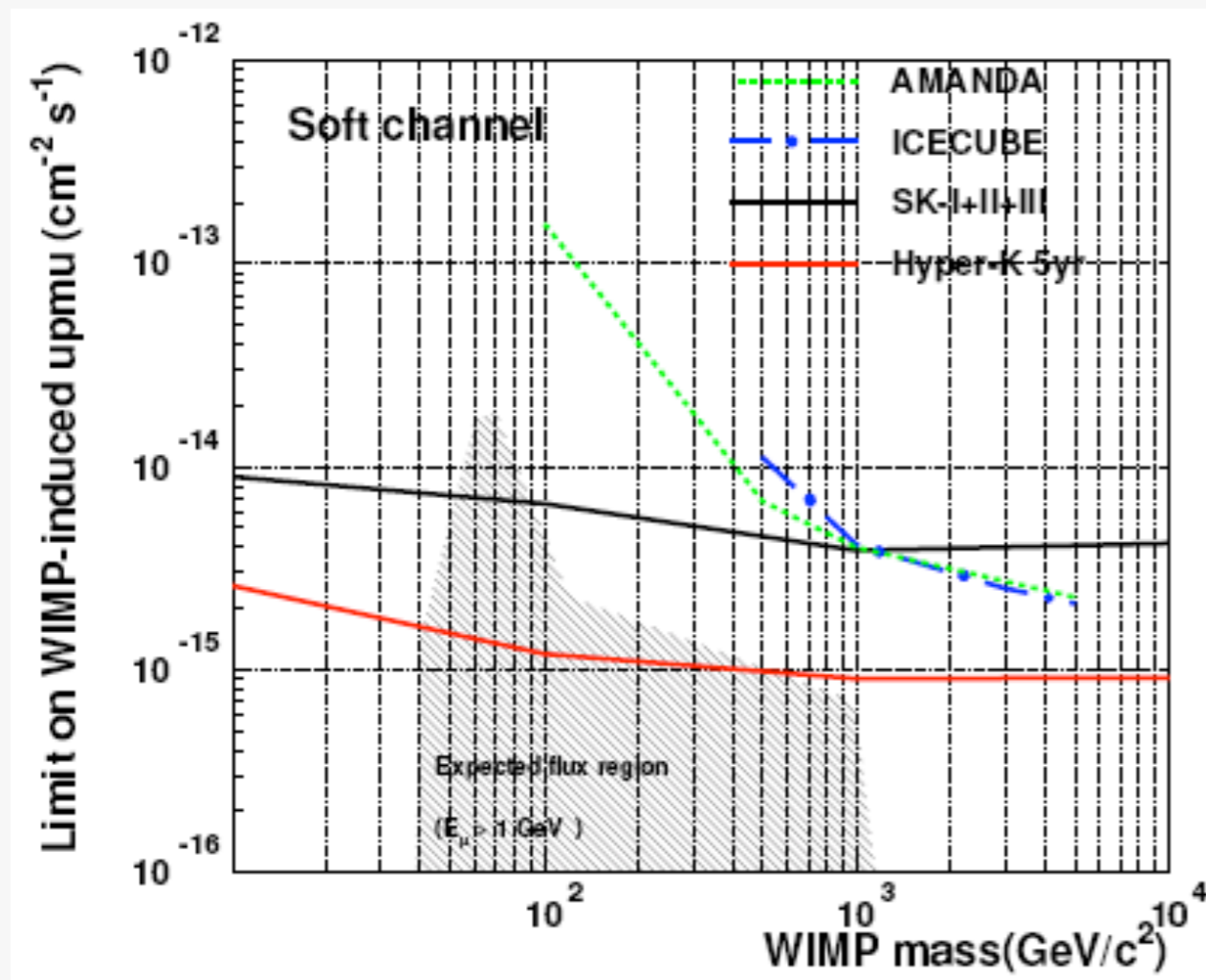
# Supernova neutrinos



- Expect 30–50  $\nu$  events even from Andromeda galaxy
  - Every 10–20 years instead of 30–50 years from our galaxy
- Sensitivity reaches to detect supernova relic neutrinos

# Indirect WIMP search

- Indirect WIMP search reaching predicted region for spin dependent interaction





# Open Meeting for the Hyper-Kamiokande Project

21-23 August 2012 *Kavli Institute for the Physics and Mathematics of the Universe (Kavli IPMU), The University of Tokyo*

Asia/Tokyo timezone

## Overview

Important Dates

Call for Abstracts

⋮ View my abstracts

⋮ Submit a new abstract

Timetable

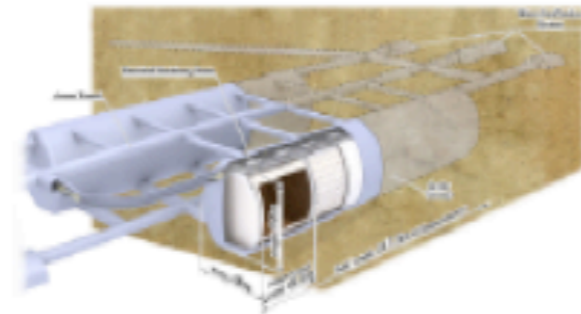
Contribution List

Registration

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Access

Accommodation



## Overview

We will hold an International Open Working Group Meeting for the Hyper-Kamiokande project. Hyper-K, which we are currently developing, is designed to be the next decade's flagship experiment for the study of neutrino oscillations, nucleon decays, and astrophysical neutrinos.

The goal of this meeting is to discuss the physics potentials of Hyper-K, the design of the detector, and necessary R&D items including:

- cavern excavation
- tank liner material and its design
- photo-sensors and their support structure
- DAQ electronics and computers
- calibration systems
- water purification systems
- software development, and so on.

Aug.21-23,2012 @ Kavli-IPMU:

<http://indico.ipmu.jp/indico/conferenceDisplay.py?confId=7>