

# Testing Radiative Mechanism for Neutrino Mass Generation

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**Workshop on Neutrinos & New Physics**

**TRIUMF, Vancouver**

**November 12-14, 2012**

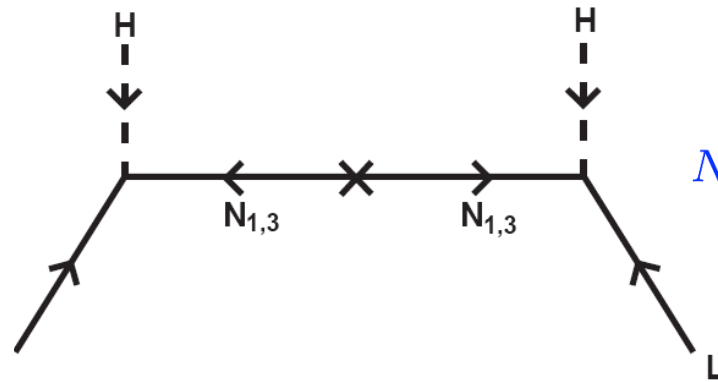


## Outline

- Some general remarks
- Specific models
- Experimental tests

# Majorana Neutrinos and Seesaw Mechanism

Type (I,III) seesaw

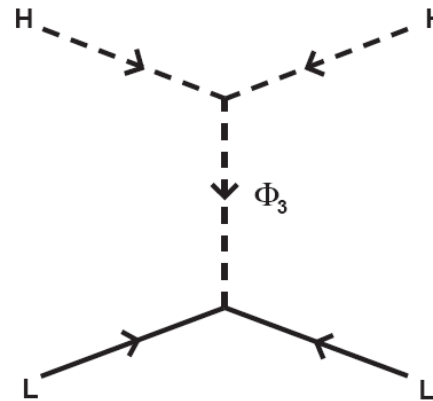


$N_1 : (1, 1, 0), N_3 : (1, 3, 0)$

Minkowski (1977)  
 Yanagida (1979)  
 Gell-Mann, Ramond, & Slansky (1980)  
 Mohapatra & Senjanovic (1980)

Foot, Lew, He, & Joshi (1989)

Type II seesaw



$\Phi_3 : (1, 3, +1)$

Mohapatra & Senjanovic (1980)  
 Schechter & Valle (1980)  
 Lazarides, Shafi, & Wetterich (1981)

$$\mathcal{L}_{\text{eff}} = \frac{LLHH}{M} \Rightarrow m_\nu \sim \frac{v^2}{M}$$

# Effective Delta(L) =2 operators for neutrino masses

Standard seesaw operator

$$\mathcal{O}_1 = L^i L^j H^k H^l \epsilon_{ik} \epsilon_{jl}$$

Operators with four fermions:

C.N. Leung, KSB (2003)

$$\mathcal{O}_2 = L^i L^j L^k e^c H^l \epsilon_{ij} \epsilon_{kl}$$

$$\mathcal{O}_3 = \{L^i L^j Q^k d^c H^l \epsilon_{ij} \epsilon_{kl}, L^i L^j Q^k d^c H^l \epsilon_{ik} \epsilon_{jl}\}$$

$$\mathcal{O}_4 = \{L^i L^j \bar{Q}_i \bar{u}^c H^k \epsilon_{jk}, L^i L^j \bar{Q}_k \bar{u}^c H^k \epsilon_{ij}\}$$

$$\mathcal{O}_5 = L^i L^j Q^k d^c H^l H^m \bar{H}_i \epsilon_{jl} \epsilon_{km}$$

$$\mathcal{O}_6 = L^i L^j \bar{Q}_k \bar{u}^c H^l H^k \bar{H}_i \epsilon_{jl}$$

$$\mathcal{O}_7 = L^i Q^j \bar{e}^c \bar{Q}_k H^k H^l H^m \epsilon_{il} \epsilon_{jm}$$

$$\mathcal{O}_8 = L^i \bar{e}^c \bar{u}^c d^c H^j \epsilon_{ij}$$

Choi, Jeong, Song (2002)  
De Gouvea, Jenkins (2008)  
Angel, Volkas (2012)

## Operators with six fermions:

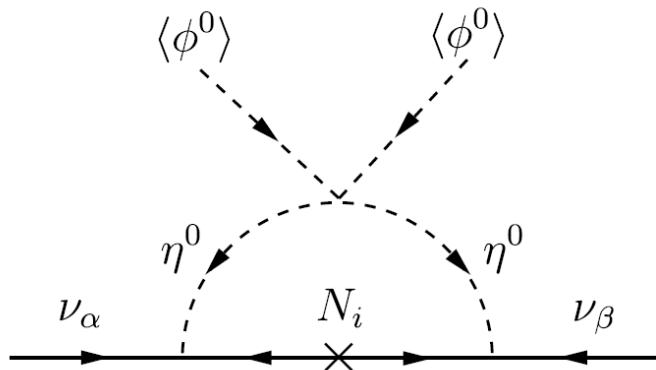
$$\begin{aligned}
 \mathcal{O}_9 &= L^i L^j L^k e^c L^l e^c \epsilon_{ij} \epsilon_{kl} \\
 \mathcal{O}_{10} &= L^i L^j L^k e^c Q^l d^c \epsilon_{ij} \epsilon_{kl} \\
 \mathcal{O}_{11} &= \{L^i L^j Q^k d^c Q^l d^c \epsilon_{ij} \epsilon_{kl}, \quad L^i L^j Q^k d^c Q^l d^c \epsilon_{ik} \epsilon_{jl}\} \\
 \mathcal{O}_{12} &= \{L^i L^j \bar{Q}_i \bar{u}^c \bar{Q}_j \bar{u}^c, \quad L^i L^j \bar{Q}_k \bar{u}^c \bar{Q}_l \bar{u}^c \epsilon_{ij} \epsilon^{kl}\} \\
 \mathcal{O}_{13} &= L^i L^j \bar{Q}_i \bar{u}^c L^l e^c \epsilon_{jl} \\
 \mathcal{O}_{14} &= \{L^i L^j \bar{Q}_k \bar{u}^c Q^k d^c \epsilon_{ij}, \quad L^i L^j \bar{Q}_i \bar{u}^c Q^l d^c \epsilon_{jl}\} \\
 \mathcal{O}_{15} &= L^i L^j L^k d^c \bar{L}_i \bar{u}^c \epsilon_{jk} \\
 \mathcal{O}_{16} &= L^i L^j e^c d^c \bar{e}^c \bar{u}^c \epsilon_{ij} \\
 \mathcal{O}_{17} &= L^i L^j d^c d^c \bar{d}^c \bar{u}^c \epsilon_{ij} \\
 \mathcal{O}_{18} &= L^i L^j d^c u^c \bar{u}^c \bar{u}^c \epsilon_{ij} \\
 \mathcal{O}_{19} &= L^i Q^j d^c d^c \bar{e}^c \bar{u}^c \epsilon_{ij} \\
 \mathcal{O}_{20} &= L^i d^c \bar{Q}_i \bar{u}^c \bar{e}^c \bar{u}^c
 \end{aligned}$$

## Dimension 5 operator through loops

$\mathcal{O}_1 = L^i L^j H^k H^l \epsilon_{ik} \epsilon_{jl}$  may arise at loops  
without light fermion mass factor

Example: Inert Doublet Model

E. Ma (2006)



$$m_\nu \approx \frac{Y_\nu^2 \lambda_5 v^2}{16\pi^2} \frac{M_R}{M_\eta^2 + M_R^2}$$

From  $m_\nu$  alone,  $M_R, M_\eta \sim 10^{12}$  GeV will work

Dark matter would require  $M_R \sim M_\eta \sim$  TeV

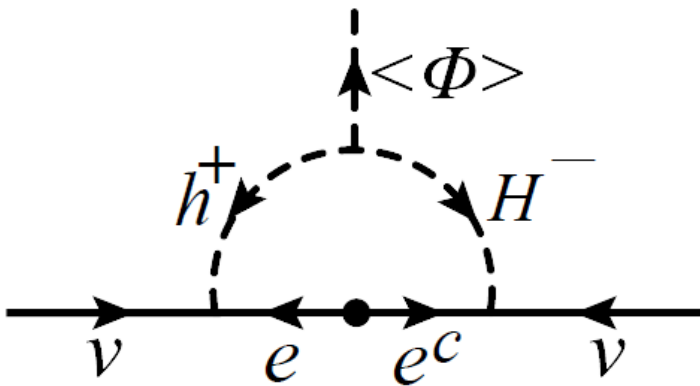
## Dimension 5 operator with one light fermion mass

$\mathcal{O}_1 = L^i L^j H^k H^l \epsilon_{ik} \epsilon_{jl}$  can arise at loops with one light fermion mass factor

Example: General Zee model

A. Zee (1980)

Effective operator:  $\mathcal{O}_2 = L^i L^j L^k e^c H^l \epsilon_{ij} \epsilon_{kl}$



$$m_\nu \approx \frac{f Y'_\ell}{16\pi^2} (m_\ell v) \frac{\mu}{M_h^2 + M_H^2}$$

For  $Y'_\ell, f \sim 1$ ,  $M_h, M_H \sim 10^{10}$  GeV will work

If  $Y'_\ell = Y_\ell \Rightarrow m_\nu \sim m_\ell^2$

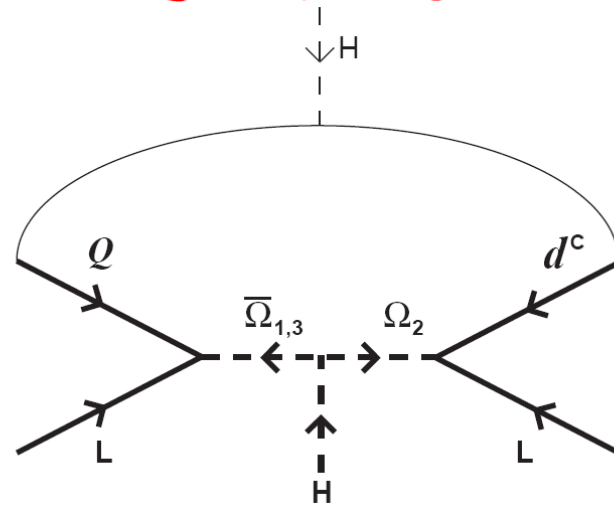
L. Wolfenstein (1981)

This version excluded by solar + KamLand data

## $\mathcal{O}_3$ model of neutrino mass

$$\mathcal{O}_3 = \{L^i L^j Q^k d^c H^l \epsilon_{ij} \epsilon_{kl}, \quad L^i L^j Q^k d^c H^l \epsilon_{ik} \epsilon_{jl}\}$$

$R$ -parity violating supersymmetric models



$$\bar{\Omega}_1 : (3^*, 1, \frac{1}{3}), \quad \Omega_2 : (3, 2, \frac{1}{6})$$

$$m_\nu \approx \frac{(\lambda')^2 m_b^2 A}{16\pi^2 \tilde{M}^2}$$

Widely studied in literature

- L. Hall, M. Suzuki (1984)
- J. Ellis, G. Gelmini, G. Ross, J. Valle (1985)
- S. Dawson (1985)
- V. Barger, G. Giudice, T. Han (1989)
- T. Banks et. al. (1995)
- F. Borzumati et. al. (1996)
- G. Bhattacharyya (1996)
- H.K. Dreiner (1997)
- R. Barbier et. al., Phys. Rept. (2005)

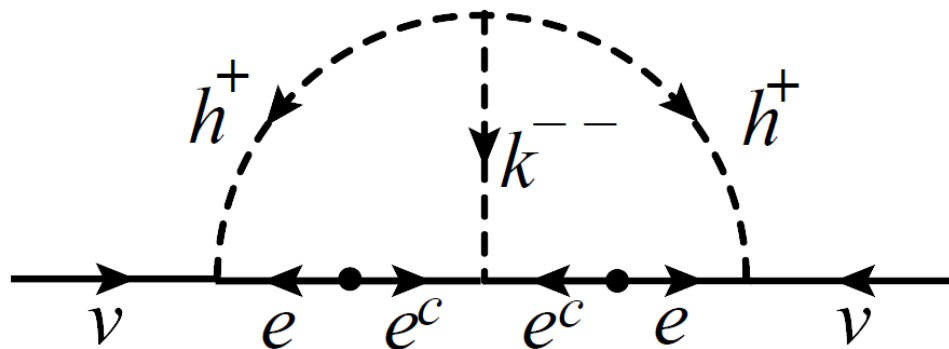


## Two-loop neutrino mass generation via $\mathcal{O}_9$

$$\mathcal{L} = f_{ij} L_i^a L_j^b h^+ \epsilon_{ab} + g_{ij} e_i^c e_j^c k^{--} + \mu h^+ h^+ k^{--} + \text{h.c.}$$



$$\mathcal{O}_9 = L^i L^j L^k e^c L^l e^c \epsilon_{ij} \epsilon_{kl}$$



A. Zee, (1985)  
Babu (1988)

Consistent with all neutrino oscillation data

Predicts doubly charged Higgs boson with TeV mass

One neutrino is nearly massless

## Two-loop neutrino mass model

$$(\mathcal{M}_\nu)_{ab} = 16\mu f_{ac} m_c g_{cd}^* I_{cd} m_d f_{bd}$$

$$I_{cd} = \int \frac{d^4 k}{(2\pi)^4} \int \frac{d^4 q}{(2\pi)^4} \frac{1}{(k^2 - m_c^2)} \frac{1}{(k^2 - m_h^2)} \frac{1}{(q^2 - m_d^2)} \frac{1}{(q^2 - m_h^2)} \frac{1}{(k - q)^2 - m_k^2}$$

$$I_{cd} \simeq I = \frac{1}{(16\pi^2)^2} \frac{1}{m_h^2} \tilde{I} \left( \frac{m_k^2}{m_h^2} \right)$$

$$\tilde{I}(r) = - \int_0^1 dx \int_0^{1-x} dy \frac{1-y}{x + (r-1)y + y^2} \log \frac{y(1-y)}{x + ry} \quad \tilde{I}(r) = \begin{cases} 1 + \frac{3}{\pi^2}(\log^2 r - 1) & \text{for } r \gg 1 \\ 1 & \text{for } r \rightarrow 0 \end{cases}$$

$$M_\nu = \xi f \omega f^T$$

$$f_{ab} = -f_{ba}, \quad \omega_{ab} = g_{ab} m_a m_b$$

## Two-loop neutrino mass model

$f$  has an eigenvector with zero eigenvalue:

$$v_0^T = (1, -\epsilon, \epsilon'); \quad f v_0 = 0 .$$

$$\epsilon = f_{e\tau} / f_{\mu\tau}, \quad \epsilon' = f = e_{\mu} / f_{\mu\tau}$$

$v_0$  is an eigenvector of  $M_\nu$  with zero eigenvalue:  $M_\nu v_0 = 0$

$$\epsilon = \frac{m_{12}m_{33} - m_{13}m_{23}}{m_{22}m_{33} - m_{23}^2} \quad \epsilon' = \frac{m_{12}m_{23} - m_{13}m_{22}}{m_{22}m_{33} - m_{23}^2}$$

$$\epsilon = \tan\theta_{12} \frac{\cos\theta_{23}}{\cos\theta_{13}} + \tan\theta_{13} \sin\theta_{23} e^{-i\delta} \quad \text{Normal hierarchy}$$

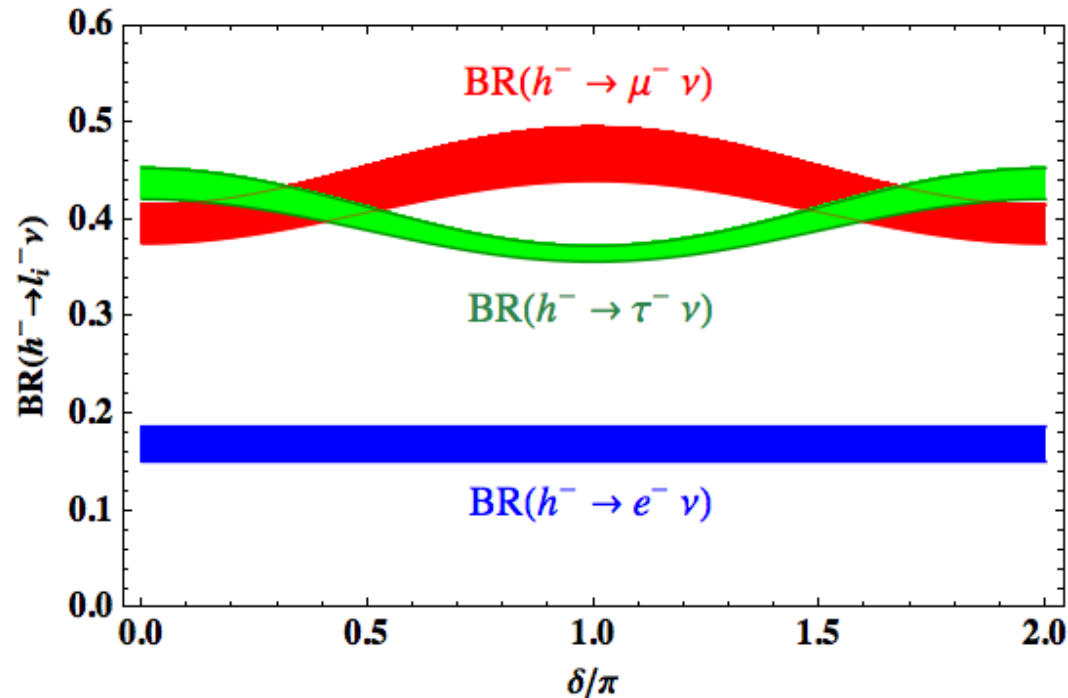
$$\epsilon' = \tan\theta_{12} \frac{\sin\theta_{23}}{\cos\theta_{13}} - \tan\theta_{13} \cos\theta_{23} e^{-i\delta}$$

## Two-loop neutrino mass model (cont.)

Inverted hierarchy:

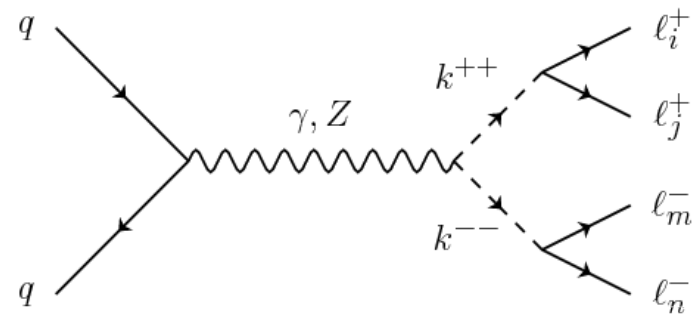
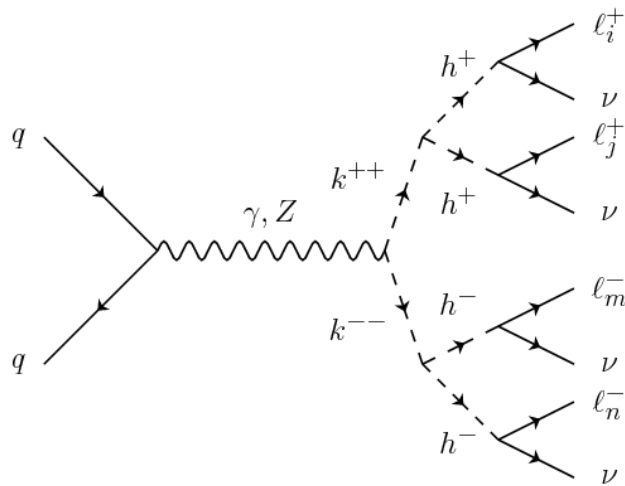
$$\epsilon = -\sin\theta_{23} \cot\theta_{13} e^{-i\delta} \quad , \quad \epsilon' = \cos\theta_{23} \cot\theta_{13} e^{-i\delta}$$

$h^- \rightarrow \ell^- \nu$  branching ratios fixed:



## Production of $h^+h^+h^-h^-$

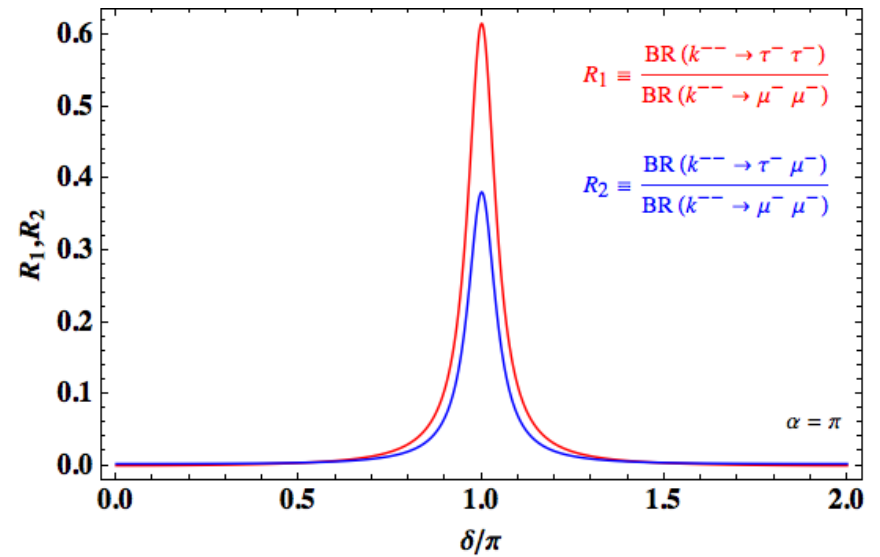
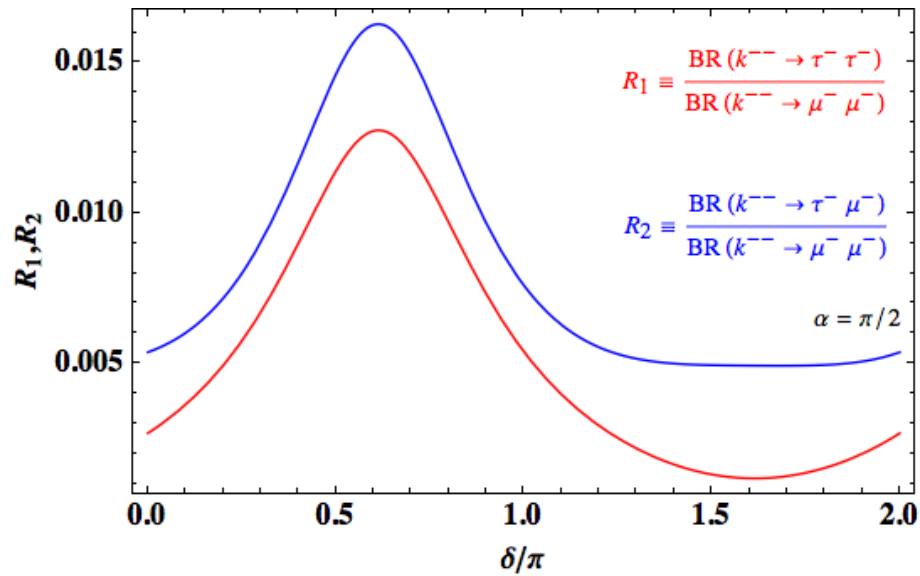
$pp \rightarrow k^{++}k^{--}$  with  $k^{++} \rightarrow h^+h^+$



CMS limit of  $m_{k^{++}} > 355$  GeV not applicable for  $k^{++} \rightarrow h^+h^+$  decay

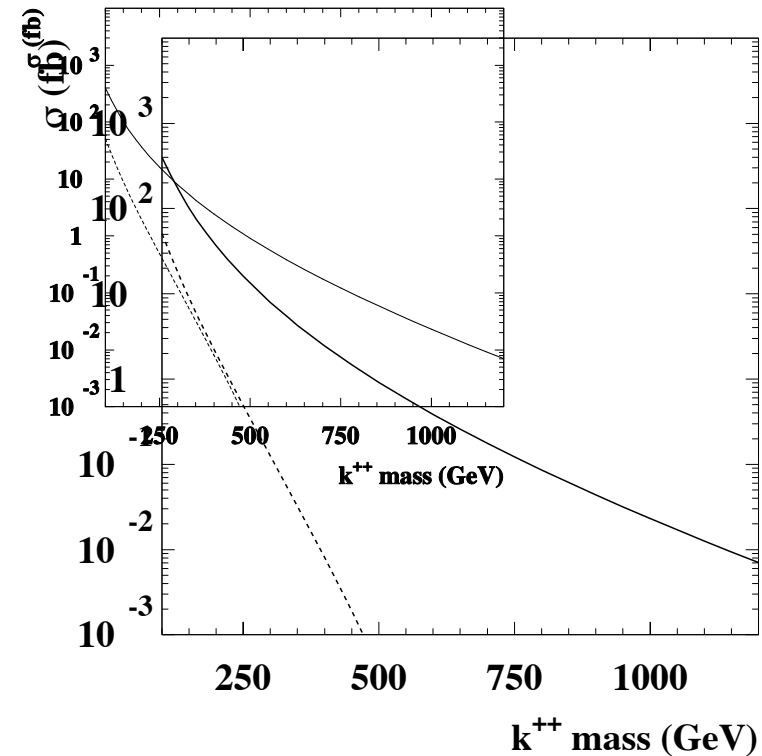
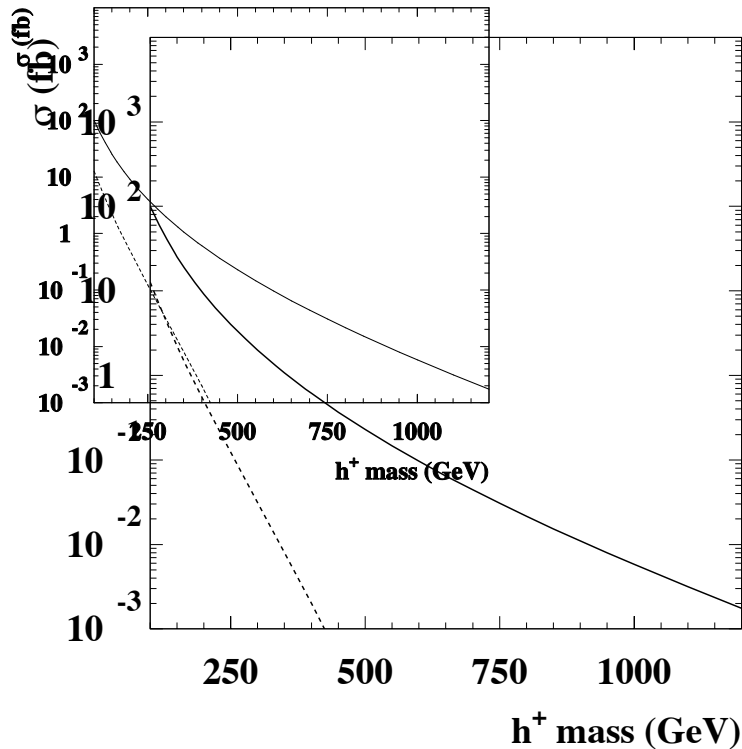
Branching ratios of  $k^{++} \rightarrow \ell^+\ell^+$  predicted from neutrino data

# Branching ratio for $k^{++} \rightarrow l^+ l^+$



Relative branching ratio gives insight into Majorana phases

# Cross section for $h^+$ and $k^{++}$ at LHC and Tevatron



LHC: Solid line, Tevatron: dashed

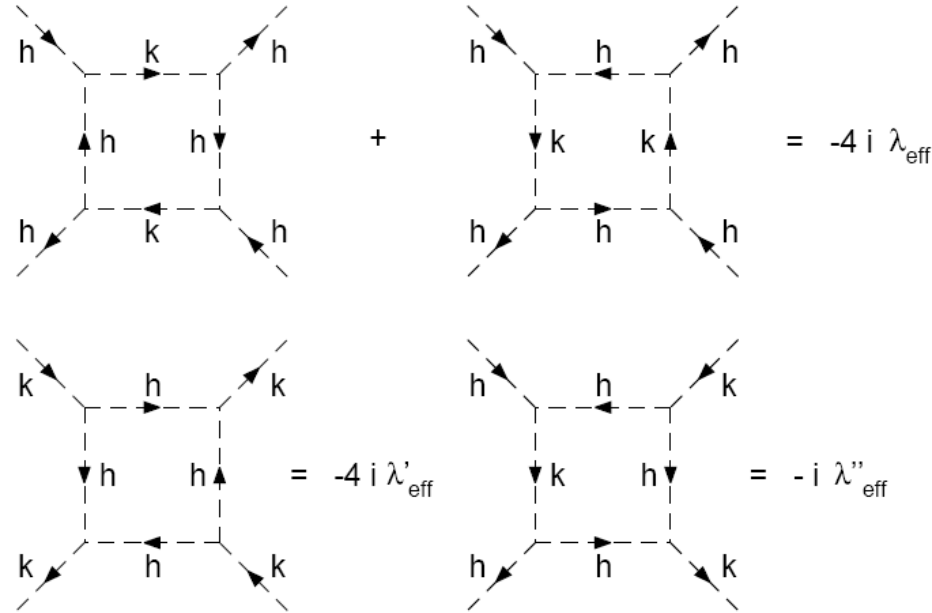
Detailed study:

K.S. Babu, C. Macesanu (2005)

D. Sierra, M. Hirsch (2006)

M. Nebot, J. Oliver, D. Paolo, A. Santamaria (2008)

# Perturbativity constraints on $\mu$



$$-\mathcal{L}_{\text{eff}} = \lambda_{\text{eff}}(h^+)^2(h^-)^2 + \lambda'_{\text{eff}}(k^{++})^2(k^{--})^2 + \lambda''_{\text{eff}}(h^+h^-)(k^{--}k^{++})$$

$$\mu < \begin{cases} m_h \times (6\pi^2)^{1/4} & \text{if } m_k \approx m_h \\ m_h \times (2\pi^2)^{1/4} & \text{if } m_k \ll m_h \\ m_h \times (24\pi^2)^{1/4} & \text{if } m_k \gg m_h . \end{cases}$$



## Lepton flavor violation constraints

Process	Experiment (90% CL)	Bound (90% CL)
$\mu^- \rightarrow e^+ e^- e^-$	$\text{BR} < 1.0 \times 10^{-12}$	$ g_{e\mu} g_{ee}^*  < 2.3 \times 10^{-5} (m_k/\text{TeV})^2$
$\tau^- \rightarrow e^+ e^- e^-$	$\text{BR} < 3.6 \times 10^{-8}$	$ g_{e\tau} g_{ee}^*  < 0.010 (m_k/\text{TeV})^2$
$\tau^- \rightarrow e^+ e^- \mu^-$	$\text{BR} < 2.7 \times 10^{-8}$	$ g_{e\tau} g_{e\mu}^*  < 0.006 (m_k/\text{TeV})^2$
$\tau^- \rightarrow e^+ \mu^- \mu^-$	$\text{BR} < 2.3 \times 10^{-8}$	$ g_{e\tau} g_{\mu\mu}^*  < 0.008 (m_k/\text{TeV})^2$
$\tau^- \rightarrow \mu^+ e^- e^-$	$\text{BR} < 2.0 \times 10^{-8}$	$ g_{\mu\tau} g_{ee}^*  < 0.008 (m_k/\text{TeV})^2$
$\tau^- \rightarrow \mu^+ e^- \mu^-$	$\text{BR} < 3.7 \times 10^{-8}$	$ g_{\mu\tau} g_{e\mu}^*  < 0.008 (m_k/\text{TeV})^2$
$\tau^- \rightarrow \mu^+ \mu^- \mu^-$	$\text{BR} < 3.2 \times 10^{-8}$	$ g_{\mu\tau} g_{\mu\mu}^*  < 0.010 (m_k/\text{TeV})^2$
$\mu^+ e^- \rightarrow \mu^- e^+$	$G_{M\bar{M}} < 0.003 G_F$	$ g_{ee} g_{\mu\mu}^*  < 0.2 (m_k/\text{TeV})^2$

# Universality and LFV constraints

SM Test	Experiment	Bound (90%CL)
lept./hadr. univ.	$\sum_{q=d,s,b}  V_{uq}^{exp} ^2 = 0.9992 \pm 0.0011$	$ f_{e\mu} ^2 < 0.015 (m_h/\text{TeV})^2$
$\mu/e$ universality	$g_\mu^{exp}/g_e^{exp} = 1.0001 \pm 0.0020$	$  f_{\mu\tau} ^2 -  f_{e\tau} ^2  < 0.05 (m_h/\text{TeV})^2$
$\tau/\mu$ universality	$g_\tau^{exp}/g_\mu^{exp} = 1.0004 \pm 0.0022$	$  f_{e\tau} ^2 -  f_{e\mu} ^2  < 0.06 (m_h/\text{TeV})^2$
$\tau/e$ universality	$g_\tau^{exp}/g_e^{exp} = 1.0004 \pm 0.0023$	$  f_{\mu\tau} ^2 -  f_{e\mu} ^2  < 0.06 (m_h/\text{TeV})^2$

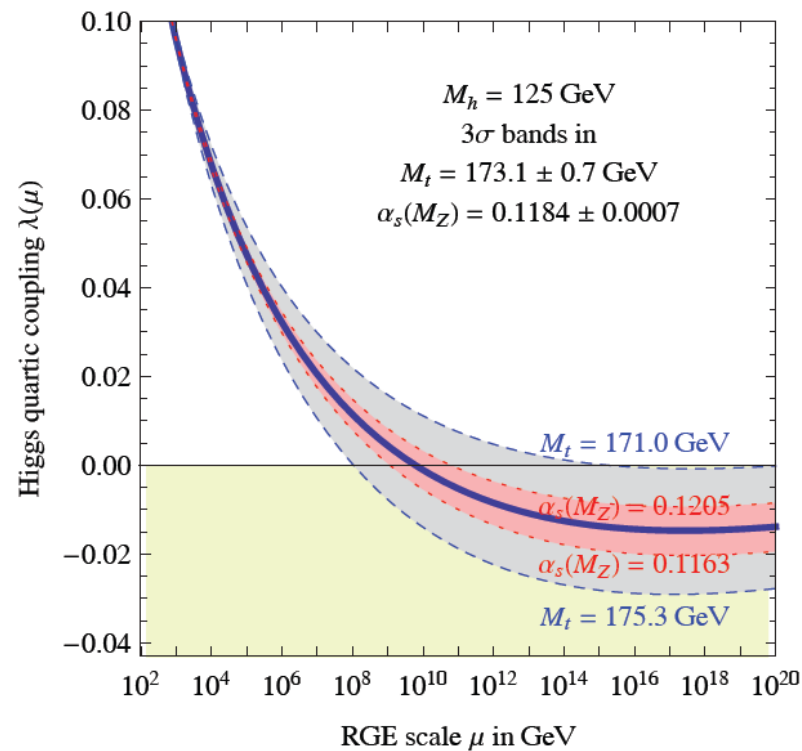
Table II: Constraints from universality of charged currents.

Experiment	Bound (90%CL)
$\delta a_e = (12 \pm 10) \times 10^{-12}$	$r ( f_{e\mu} ^2 +  f_{e\tau} ^2) + 4 ( g_{ee} ^2 +  g_{e\mu} ^2 +  g_{e\tau} ^2) < 5.5 \times 10^3 (m_k/\text{TeV})^2$
$\delta a_\mu = (21 \pm 10) \times 10^{-10}$	$r ( f_{e\mu} ^2 +  f_{\mu\tau} ^2) + 4 ( g_{e\mu} ^2 +  g_{\mu\mu} ^2 +  g_{\mu\tau} ^2) < 7.9 (m_k/\text{TeV})^2$
$BR(\mu \rightarrow e\gamma) < 1.2 \times 10^{-11}$	$r^2  f_{e\tau}^* f_{\mu\tau} ^2 + 16  g_{ee}^* g_{e\mu} + g_{e\mu}^* g_{\mu\mu} + g_{e\tau}^* g_{\mu\tau} ^2 < 3.4 \times 10^{-5} (m_k/\text{TeV})^4$
$BR(\tau \rightarrow e\gamma) < 1.1 \times 10^{-7}$	$r^2  f_{e\mu}^* f_{\mu\tau} ^2 + 16  g_{ee}^* g_{e\tau} + g_{e\mu}^* g_{\mu\tau} + g_{e\tau}^* g_{\tau\tau} ^2 < 1.7 (m_k/\text{TeV})^4$
$BR(\tau \rightarrow \mu\gamma) < 4.5 \times 10^{-8}$	$r^2  f_{e\mu}^* f_{e\tau} ^2 + 16  g_{e\mu}^* g_{e\tau} + g_{\mu\mu}^* g_{\mu\tau} + g_{\mu\tau}^* g_{\tau\tau} ^2 < 0.7 (m_k/\text{TeV})^4$

Table III: Constraints from lepton number violating photon interactions.

# Boundedness of Higgs potential

In the Standard Model with  $m_H = 125$  GeV Higgs potential becomes unbounded at  $\mu \sim 10^{10}$  GeV



deGrassi et. al. (2012)

## Boundedness in Two-loop model

$$\begin{aligned} V = & -\mu_1^2 \Phi^\dagger \Phi + \mu_2^2 h^+ h^- + \mu_3^2 k^{++} k^{--} + \{\mu h^+ h^+ k^{--} + h.c.\} \\ & + \frac{\lambda_1}{2} (\Phi^\dagger \Phi)^2 + \frac{\lambda_2}{2} (h^+ h^-)^2 + \frac{\lambda_3}{2} (k^{++} k^{--})^2 \\ & + \lambda_4 (\Phi^\dagger \Phi) (h^+ h^-) + \lambda_5 (\Phi^\dagger \Phi) (k^{++} k^{--}) + \lambda_6 (h^+ h^-) (k^{++} k^{--}) \end{aligned}$$

Boundedness  $\Rightarrow$

$$\lambda_1 > 0, \quad \lambda_2 > 0, \quad \lambda_3 > 0$$

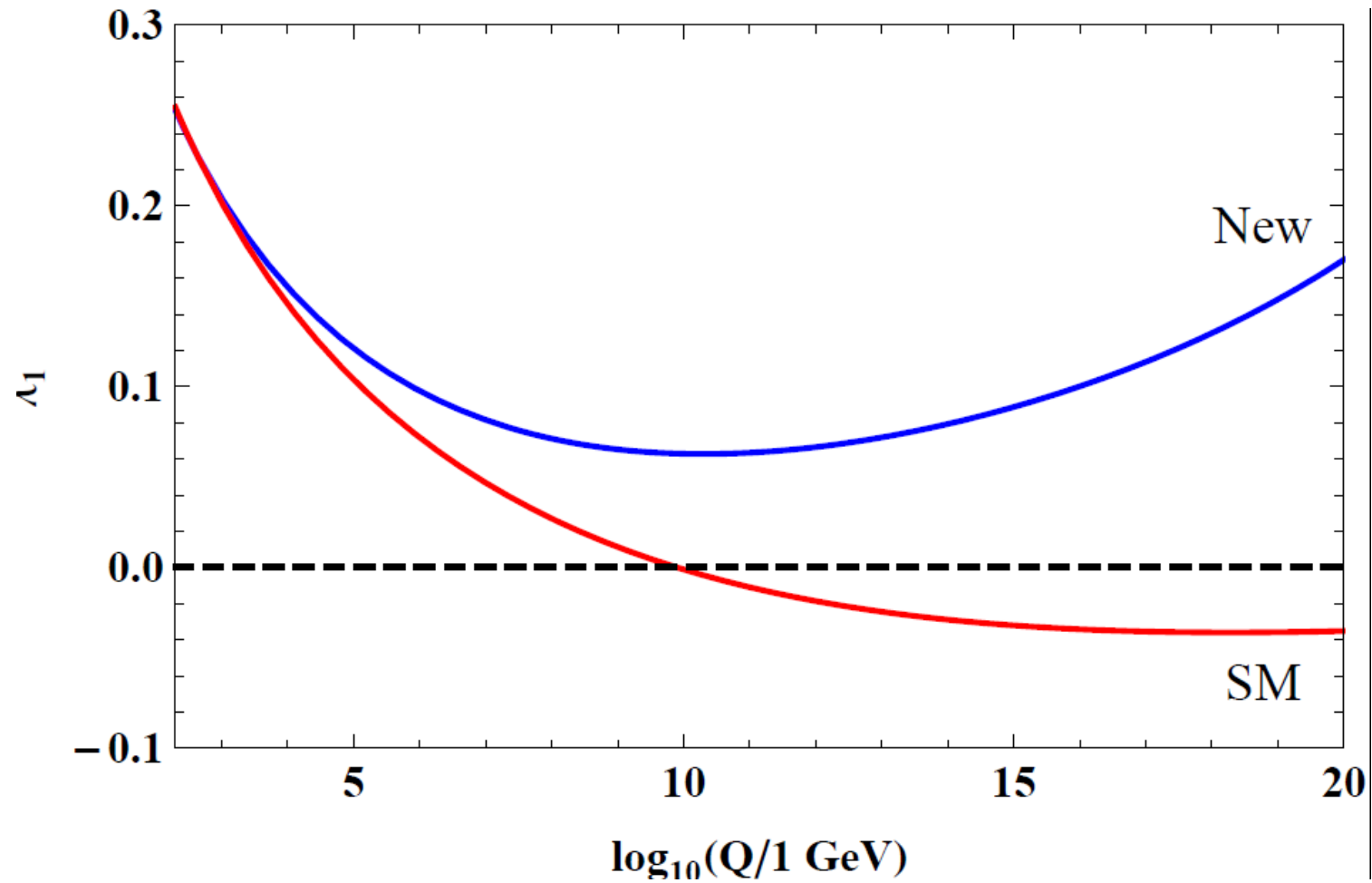
$$\lambda_4 + \sqrt{\lambda_1 \lambda_2} \geq 0, \quad \lambda_3 + \sqrt{\lambda_1 \lambda_3} \geq 0, \quad \lambda_6 + \sqrt{\lambda_2 \lambda_3} \geq 0$$

$$\lambda_4 \sqrt{\lambda_3} + \lambda_6 \sqrt{\lambda_1} + \lambda_5 \sqrt{\lambda_2} + \sqrt{\lambda_1 \lambda_2 \lambda_3} \geq 0 \quad \text{or} \quad \text{Det}(\hat{\lambda}) \geq 0$$

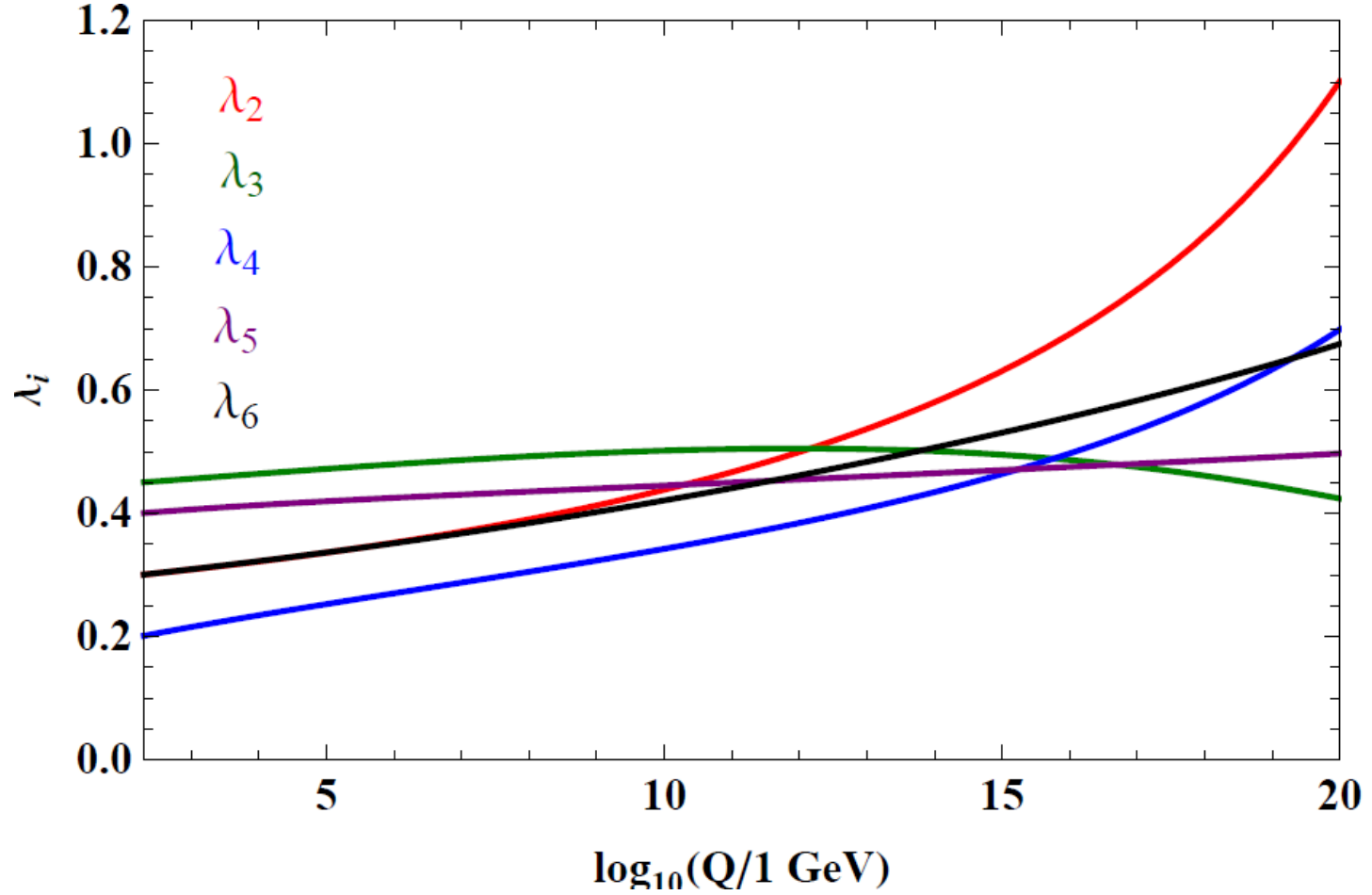
Related to Co-positivity of matrices

**Hadeler (1983)**

## Evolution of $\lambda_1$



J. Julio, KB (2012)



$$\lambda_1 = 0.258, \quad \lambda_2 = 0.3, \quad \lambda_3 = 0.45, \\ \lambda_4 = 0.2, \quad \lambda_5 = 0.4, \quad \lambda_6 = 0.3, \quad h_{\mu\tau} = 0.25.$$

# Neutrino mass model with Leptoquarks

K.S. Babu, J. Julio (2010)

$\mathcal{O}_8$  can be induced via scalar leptoquarks

$$\mathcal{O}_8 = L^i \bar{e}^c \bar{u}^c d^c H^j \epsilon_{ij}$$

$M_\nu$  arises as  $d = 7$  operators at two loop

$$m_\nu \sim \frac{m_t m_b m_\tau \mu v}{(16\pi^2)^2 M^4}$$

Leptoquark within reach of LHC

Leptoquark branching ratios probe  $m_\nu$

## Leptoquark model for neutrino mass

Add two leptoquark fields to standard model

$$\Omega(3, 2, 1/6) = \begin{bmatrix} \omega^{2/3} \\ \omega^{-1/3} \end{bmatrix}, \quad \chi^{-1/3}(3, 1, -1/3)$$

$$\mathcal{L}_{\text{Yukawa}} = Y_{ij} L_i^\alpha d_j^c \Omega^\beta \epsilon_{\alpha\beta} + F_{ij} e_i^c u_j^c \chi^{-1/3} + h.c.$$

$$V = \mu \Omega^\dagger H \chi^{-1/3} + h.c.$$

$B$  is unbroken, but  $L$  is softly broken

$$\mathcal{L}_\nu = Y_{ij} (\nu_i d_j^c \omega^{-1/3} - e_i d_j^c \omega^{2/3}) + F_{ij} e_i^c u_j^c \chi^{-1/3} + h.c.$$

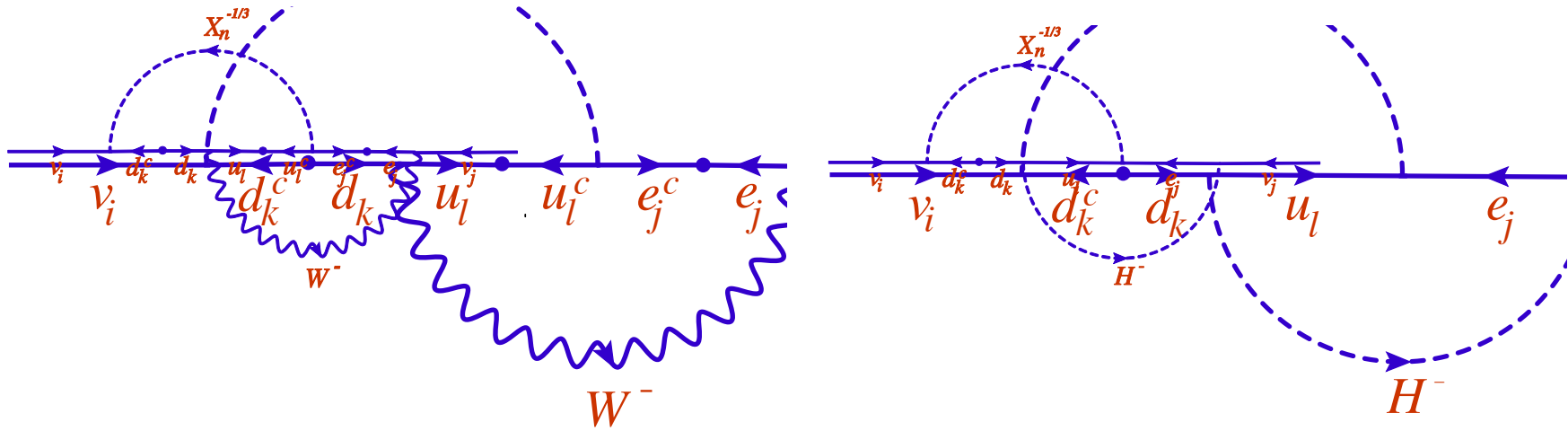
$$V = \mu (\omega^{-2/3} H^+ + \omega^{1/3} H^0) \chi^{-1/3} + h.c.$$

$$M_{\text{LQ}}^2 = \begin{pmatrix} m_\omega^2 & \mu v \\ \mu^* v & m_\chi^2 \end{pmatrix}$$

Mass eigenvalues  $M_{1,2}^2$ , mixing  $\theta$



# Neutrino mass generation



$$M_\nu = \hat{m}_0 \hat{I} \left[ Y D_d V^T D_u F^\dagger D_\ell + D_\ell F^* D_u V D_d Y^T \right]$$

$$\hat{m}_0 = \left( \frac{C g^2 \sin 2\theta}{(16\pi^2)^2} \right) \left( \frac{m_t m_b m_\tau}{M_1^2} \right)$$

$$D_u = \text{diag.} \left[ \frac{m_u}{m_t}, \frac{m_c}{m_t}, 1 \right], \quad D_d = \text{diag.} \left[ \frac{m_d}{m_b}, \frac{m_s}{m_b}, 1 \right], \quad D_\ell = \text{diag.} \left[ \frac{m_e}{m_\tau}, \frac{m_\mu}{m_\tau}, 1 \right]$$

$\hat{I}$ : Loop integral function

## Neutrino phenomenology

$$M_\nu \simeq m_0 \begin{pmatrix} 0 & \frac{1}{2} \frac{m_\mu}{m_\tau} xy & \frac{1}{2} y \\ \frac{1}{2} \frac{m_\mu}{m_\tau} xy & \frac{m_\mu}{m_\tau} xz & \frac{1}{2} z + \frac{1}{2} \frac{m_\mu}{m_\tau} x \\ \frac{1}{2} y & \frac{1}{2} z + \frac{1}{2} \frac{m_\mu}{m_\tau} x & 1 + w \end{pmatrix} .$$

$$x \equiv \frac{F_{23}^*}{F_{33}^*}, \quad y \equiv \frac{Y_{13}}{Y_{33}}, \quad z \equiv \frac{Y_{23}}{Y_{33}}, \quad m_0 = 2 \hat{m}_0 F_{33}^* Y_{33} \hat{I}$$

$$w \equiv \frac{F_{32}^* Y_{32}}{F_{33}^* Y_{33}} \left( \frac{m_c}{m_t} \right) \left( \frac{m_s}{m_b} \right) \frac{I_{jk2}}{I_{jk3}}$$

Note that (1,1) entry is zero

This matrix has zero determinant (if  $w \ll 1$ )  $\Rightarrow$

$$m_1 = 0, \quad \tan^2 \theta_{13} = \frac{m_2}{m_3} \sin^2 \theta_{12}$$

$$\beta = 2\delta + \pi, \quad \alpha = 0 \quad (\text{Majorana phases})$$

# Predictions for $w \gg 1$

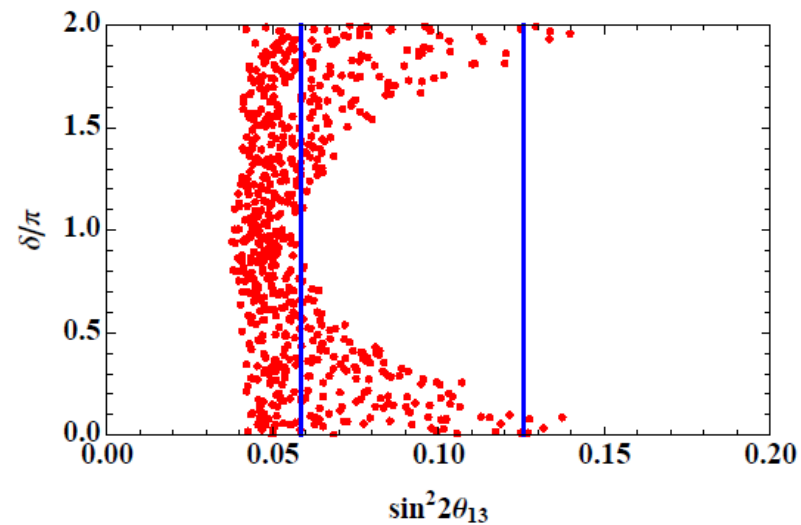
- For  $w \gg 1$ ,

$$w \equiv \frac{F_{32}^* Y_{32}}{F_{33}^* Y_{33}} \left( \frac{m_c}{m_t} \right) \left( \frac{m_s}{m_b} \right) \frac{l_{jk2}}{l_{jk3}} \gg 1 \quad \rightarrow \quad |F_{33} Y_{33}| \ll |F_{32} Y_{32}|$$

- This could generate  $(M_\nu)_{13} \simeq (M_\nu)_{11} \simeq 0$ .

Glashow, Frampton, & Marfatia (2002)

Xing (2002)



- The value of  $\theta_{13}$  is consistent with current measurements (the blue lines correspond to  $2\sigma$  allowed value from Daya Bay).

## Leptoquark branching ratios

$$\Gamma(\omega^{2/3} \rightarrow e^+b) : \Gamma(\omega^{2/3} \rightarrow \mu^+b) : \Gamma(\omega^{2/3} \rightarrow \tau^+b) = |y|^2 : |z|^2 : 1$$

Measuring any of the branching ratios will fix  $\delta$

Measuring two ratios overconstrains and gives checks

$$\Gamma(X_a^{-1/3} \rightarrow \mu^-t) : \Gamma(X_a^{-1/3} \rightarrow \tau^-t) = |x|^2 : 1$$

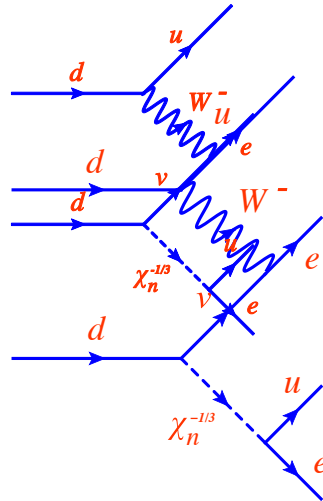
Further checks provided

Since  $|x| \gg 1$ ,  $\mu$  will dominate final states

# Neutrinoless double beta decay

$(M_\nu)_{11} \simeq 0 \Rightarrow$  No neutrino mass contribution to  $\beta\beta_{0\nu}$

**Vector-scalar exchange:**



**K.S. Babu, R.N. Mohapatra (1995)**

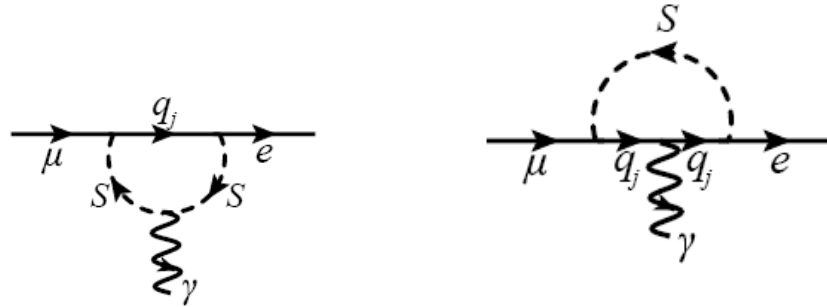
$$\mathcal{L}_{\text{eff}} = \frac{G_F^2}{2} \epsilon \bar{u} \gamma^\mu (1 - \gamma_5) d \left[ \bar{u} (1 + \gamma_5) d \bar{e} \gamma_\mu (1 - \gamma_5) \frac{1}{\not{q}} e^c + \frac{1}{4} \bar{u} \sigma^{\alpha\beta} (1 + \gamma_5) d \bar{e} \gamma_\mu (1 - \gamma_5) \frac{1}{\not{q}} \sigma_{\alpha\beta} e^c \right]$$

$$\epsilon = \frac{Y_{11}^* F_{11}}{2\sqrt{2} M_1^2 G_F} \sin 2\theta \left( 1 - \frac{M_1^2}{M_2^2} \right)$$

$$|Y_{11}^* F_{11}| < 8.4 \times 10^{-7} \left( \frac{M_1}{1 \text{ TeV}} \right)^2 \left( \frac{M_2}{1 \text{ TeV}} \right)^2 \left( \frac{1 \text{ TeV}}{\mu} \right)$$

Normal mass hierarchy and observable  $\beta\beta_{0\nu}$  possible

# $\mu \rightarrow e\gamma$ Predictions



Process	BR	Constraint
$\mu \rightarrow e\gamma$	$< 1.2 \times 10^{-11}$	$\frac{ F_3(x_b)Y_{13}^*Y_{23} ^2}{m_\Omega^4} + \frac{ \frac{1}{12}F_{11}F_{21}^* + \frac{1}{12}F_{12}F_{22}^* + F_4(x_t)F_{13}F_{23}^* ^2}{m_\chi^4} < \frac{3.1 \times 10^{-19}}{\text{GeV}^4}$
$\tau \rightarrow e\gamma$	$< 1.1 \times 10^{-7}$	$\frac{ F_3(x_b)Y_{13}^*Y_{33} ^2}{m_\Omega^4} + \frac{ \frac{1}{12}F_{11}F_{31}^* + \frac{1}{12}F_{12}F_{32}^* + F_4(x_t)F_{13}F_{33}^* ^2}{m_\chi^4} < \frac{1.6 \times 10^{-14}}{\text{GeV}^4}$
$\tau \rightarrow \mu\gamma$	$< 4.5 \times 10^{-8}$	$\frac{ F_3(x_b)Y_{23}^*Y_{33} ^2}{m_\Omega^4} + \frac{ \frac{1}{12}F_{21}F_{31}^* + \frac{1}{12}F_{22}F_{32}^* + F_4(x_t)F_{23}F_{33}^* ^2}{m_\chi^4} < \frac{6.7 \times 10^{-15}}{\text{GeV}^4}$

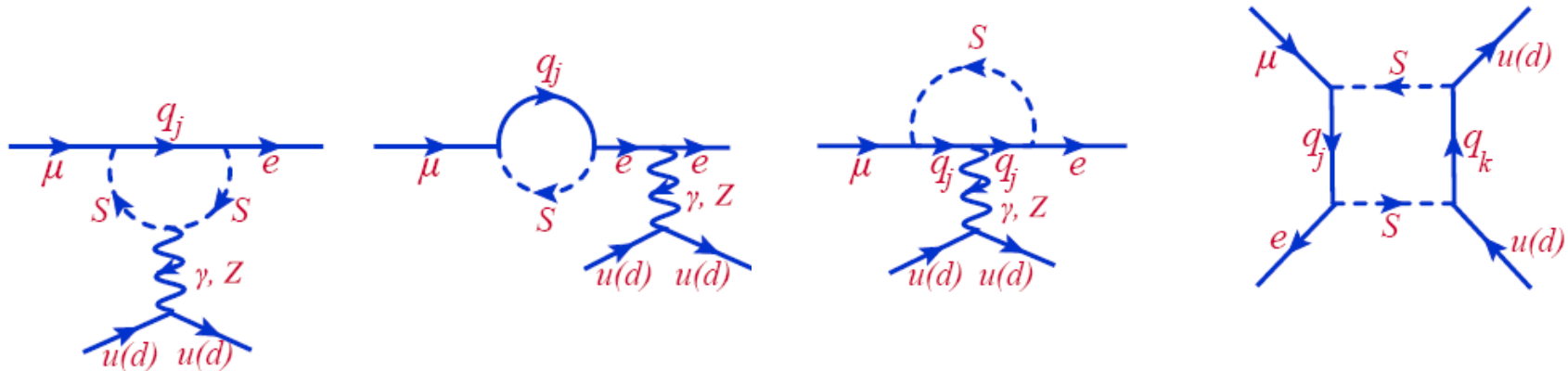
Table 1: The constraints from  $\ell_i \rightarrow \ell_j\gamma$ .

**GIM suppression!**  $A \propto \left( \frac{m_b^2}{M_{LQ}^2} \right)$

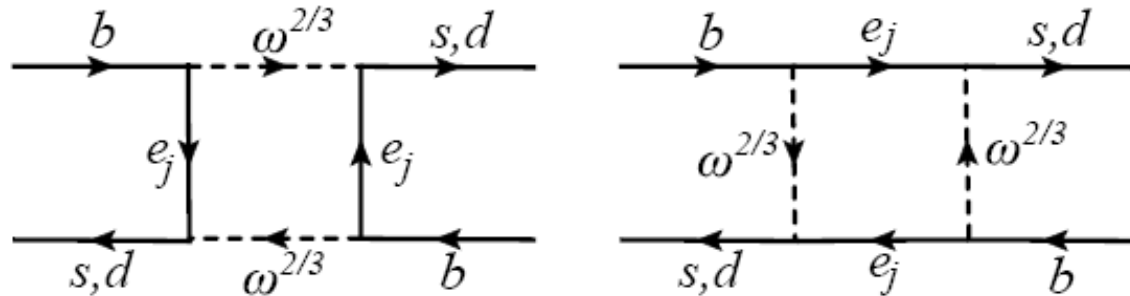
# μ – e Conversion in Nuclei

$$\text{BR}(\mu N \rightarrow e N) = \frac{|\vec{p}_e| E_e m_\mu^3 G_F^2 \alpha^3 Z_{\text{eff}}^4 F_p^2}{8\pi^2 Z \Gamma_N} \left| g_L^u(A+Z) + g_L^d(2A-Z) + 2Z\Delta g_L \right|^2 + L \leftrightarrow R$$

Element	BR	$Z_{\text{eff}}$	$F_p$	Constraint
$^{48}\text{Ti}$	$< 4.3 \times 10^{-12}$	17.61	0.53	$\left  \frac{a_j^L Y_{ij}^* Y_{2j}}{m_\Omega^2} + \frac{a_j^R F_{1j} F_{2j}^*}{m_\chi^2} \right ^2 + \left  \frac{b_j^L Y_{ij}^* Y_{2j}}{m_\Omega^2} + \frac{b_j^R F_{1j} F_{2j}^*}{m_\chi^2} \right ^2 < \frac{5.2 \times 10^{-16}}{\text{GeV}^4}$
$^{208}\text{Pb}$	$< 4.6 \times 10^{-11}$	33.81	0.15	$\left  \frac{a_j^L Y_{ij}^* Y_{2j}}{m_\Omega^2} + \frac{a_j^R F_{1j} F_{2j}^*}{m_\chi^2} \right ^2 + \left  \frac{b_j^L Y_{ij}^* Y_{2j}}{m_\Omega^2} + \frac{b_j^R F_{1j} F_{2j}^*}{m_\chi^2} \right ^2 < \frac{9.7 \times 10^{-14}}{\text{GeV}^4}$



# $B_s - \bar{B}_s$ mixing



$$\mathcal{L}_{\text{eff}}^{\text{new}} = -\frac{(Y_{i2}Y_{i3}^*)^2}{128\pi^2 m_\omega^2} (\bar{s}_R \gamma^\mu b_R) (\bar{s}_R \gamma_\mu b_R)$$

$$\langle B_s | -\mathcal{L}_{\text{eff}}^{\text{new}} | \bar{B}_s \rangle = \frac{(Y_{i2}Y_{i3}^*)^2}{192\pi^2 m_\omega^2} m_{B_s} f_{B_s}^2 B_1^{B_s}(\mu) \eta_1^{B_s}(\mu)$$

New CP violation can be as large as 40%

LQ mass  $< 500$  GeV needed



## **DØ Dimuon data:**

$$A_{sl}^b = \frac{N^{++} - N^{--}}{N^{++} + N^{--}} = -(9.57 \pm 2.51 \pm 1.46) \times 10^{-3}$$

$$A_{sl}^b(SM) = -2.3_{-0.6}^{+0.5} \times 10^{-4} \quad 3.2 \sigma \text{ discrepancy}$$

## **New contributions:**

$$\Delta M_s = \Delta M_s^{\text{SM}} \left| 1 + h_s e^{2i\sigma} \right|$$

Best fit parameters: (Ligeti, Papucci, Perez, Zupan, 2010)

$$\{h_s = 0.5, \sigma = 120^\circ\} \text{ or } \{h_s = 1.8, \sigma = 100^\circ\}$$

$h_s = 0.42, \sigma = 120^\circ$  realized with leptoquarks

Predicts  $B_s \rightarrow \tau^+ \tau^-$  decay at the percent level:

$$\text{BR}(B_s \rightarrow \tau^+ \tau^-) = 0.28\% \left( \frac{|Y_{32} Y_{33}|}{0.07} \right)^2 \left( \frac{300 \text{ GeV}}{m_\omega} \right)^4 \left( \frac{f_{B_s}}{0.24 \text{ GeV}} \right)^2$$

# Radiative Dirac Neutrino Masses

K.S. Babu, X.G. He (1988)

Left-right symmetric model with minimal Higgs content:

$$SU(3)_C \times SU(2)_L \times SU(2)_R \times U(1)_{B-L}$$

$$\text{Higgs: } (1, 2, 1, 1/2) + (1, 1, 2, 1/2)$$

Fermion mass generation via singlet fermions:

$$P(3, 1, 1, 2/3) + N(3, 1, 1, -1/3) + E(1, 1, 1, 1)$$

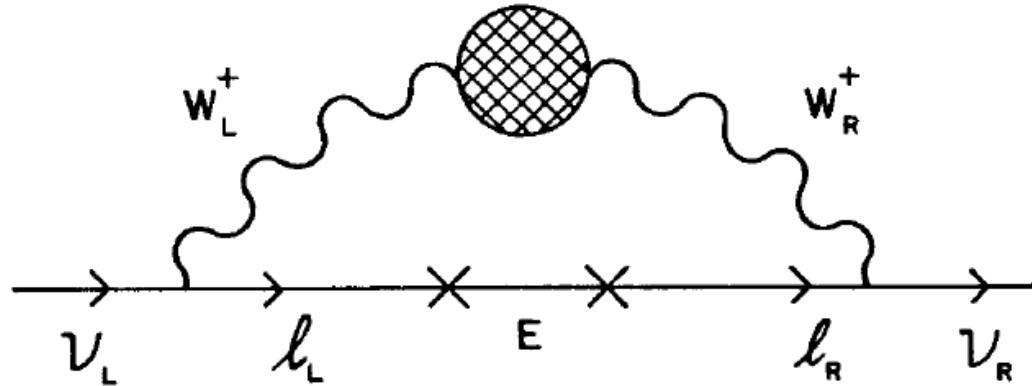
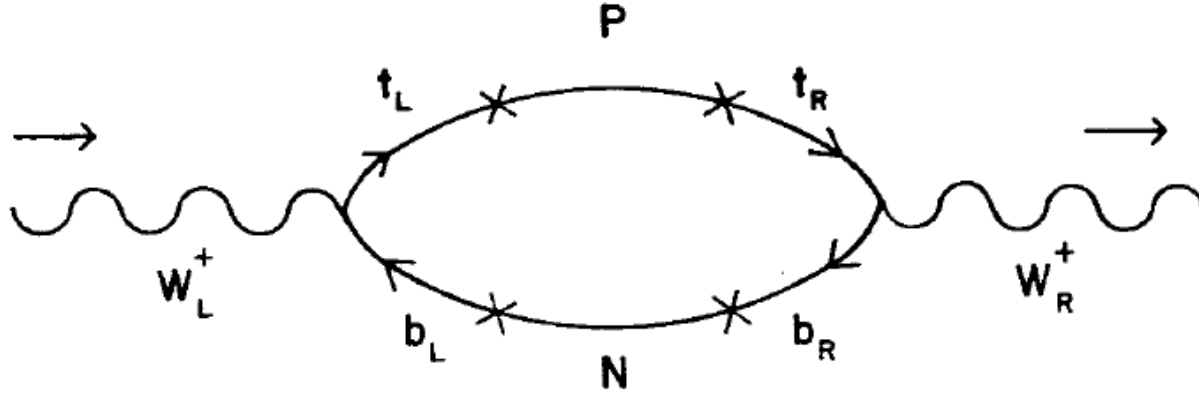
All charged fermions get mass via seesaw

Neutrinos are Dirac particles here, no seesaw for neutrinos!

Chang, Mohapatra (1987)

Davidson, Wali (1988)

Babu Mohapatra (1988)



$$m_\nu \simeq \frac{g^4}{512\pi^4} \left( \frac{m_t m_b}{M_{W_L}^2} \right) \frac{M^2}{M_{W_R}^2} m_\ell$$

## Summary and Conclusions

- Radiative neutrino mass generation a natural alternative to seesaw
- New particles must exist at the TeV scale in many cases
- Scalar decays probe neutrino mass generation and CP violation