

PSI Dark Matter Homework #1

Due: Apr. 12, 2012

1. Dark Matter Profiles:

- a) Make plots (using a plotting program of your choice) of the NFW, Iso-Core, and Einasto ($\alpha = 1.7$) galactic DM profiles using the functional forms listed in the notes. Put them on the same set of axes.
- b) We will see that the cosmic ray signal strength from given point in the sky from DM annihilation in our galaxy is proportional to $\langle\sigma v\rangle n_\chi^2$, where n_χ is the local DM number density. Assuming a constant cross section, make plots of the contributions to the signal strengths (as a function of the distance from the galactic center) for the NFW and Einasto profiles relative to the Iso profile.

2. Boltzmannology:

- a) The collision term for the Boltzmann equation for n_χ we derived in class for the $\chi\chi \rightarrow f\bar{f}$ process was

$$-\int \frac{d^3p_1}{(2\pi)^3} \frac{d^3p_2}{(2\pi)^3} g_\chi^2 \sigma v (f_1 f_2 - f_{1eq} f_{2eq}), \quad (1)$$

where “1” and “2” refer to the two χ particles. Assuming $f(E, t) = \xi(t) f_{eq}(E)$, re-express this quantity in terms of n_χ , $n_{\chi eq}$, and $\langle\sigma v\rangle$.

- b) Do the integral over p^0 in $\int d^4p \delta(p^2 - m^2) \theta(p^0)$. What does this imply for how the change in $(d\Pi) = d^3p/(2\pi)^3 2E$ under Lorentz transformations?
- c) Derive for yourself the rewriting of the Liouville term $(dn_\chi/dt + 3Hn_\chi)$ in terms of $Y_\chi = n_\chi/s$ as a function of $x = m/T$.

3. Cross Sections

- a) Consider the interaction

$$-\mathcal{L} \supset \frac{1}{\Lambda^2} \bar{\chi} \gamma^\mu \gamma^5 \chi \bar{f} \gamma_\mu f. \quad (2)$$

This could arise from integrating out a heavy vector boson that couples to both DM and the fermion f (which would give $\Lambda^2 = m_V^2/g_V^2$). Compute the squared matrix element for this process, summed over final spins and averaged over initial spins. You may assume the fermion f is massless, but keep the mass of χ around. Write your answer in terms of dot products involving the two incoming momenta p_1 and p_2 , and the outgoing momenta p_3 and p_4 .

- b) Evaluate this summed and squared matrix element in the center-of-mass frame.
- c) Turn this matrix element into σv by integrating it over the final-state phase space $(\int d\Pi_3 d\Pi_4 (2\pi)^4 \delta^{(4)}(\dots))$ and multiplying by the appropriate factors. You should be able to write everything in terms of the Mandelstam variable $s = (p_1 + p_2)^2$ and the mass of χ .

- d) Expand s to quadratic order in $v = p/E$ in the non-relativistic limit $v \ll 1$ (working in the CM frame). Plug this into your expression for σv . You should find that $\sigma v \propto v^2$. Compute the thermal average $\langle \sigma v \rangle$ following the non-relativistic prescription described in class.
- e) What would you get if you replaced $\bar{\chi} \gamma^\mu \gamma^5 \chi$ by $\bar{\chi} \gamma^\mu \chi$ in the interaction?
Hint: aside from tracking signs in a few Dirac traces, you basically shouldn't have to do any more work.